

Reminder: Relativity of Simultaneity
in Special Relativity

Einstein
1907:
"On the
Principle
of Relativity
..."

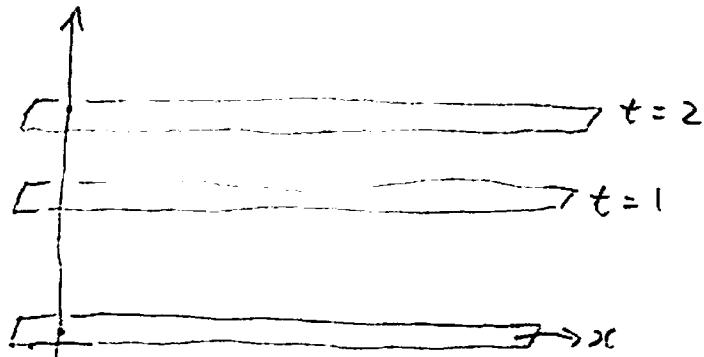
$$t' = \gamma(t - \frac{v}{c^2}x)$$

$$x' = \gamma(x - vt)$$

$$y' = y \quad z' = z$$

$$\gamma = \sqrt{1 + v^2/c^2}$$

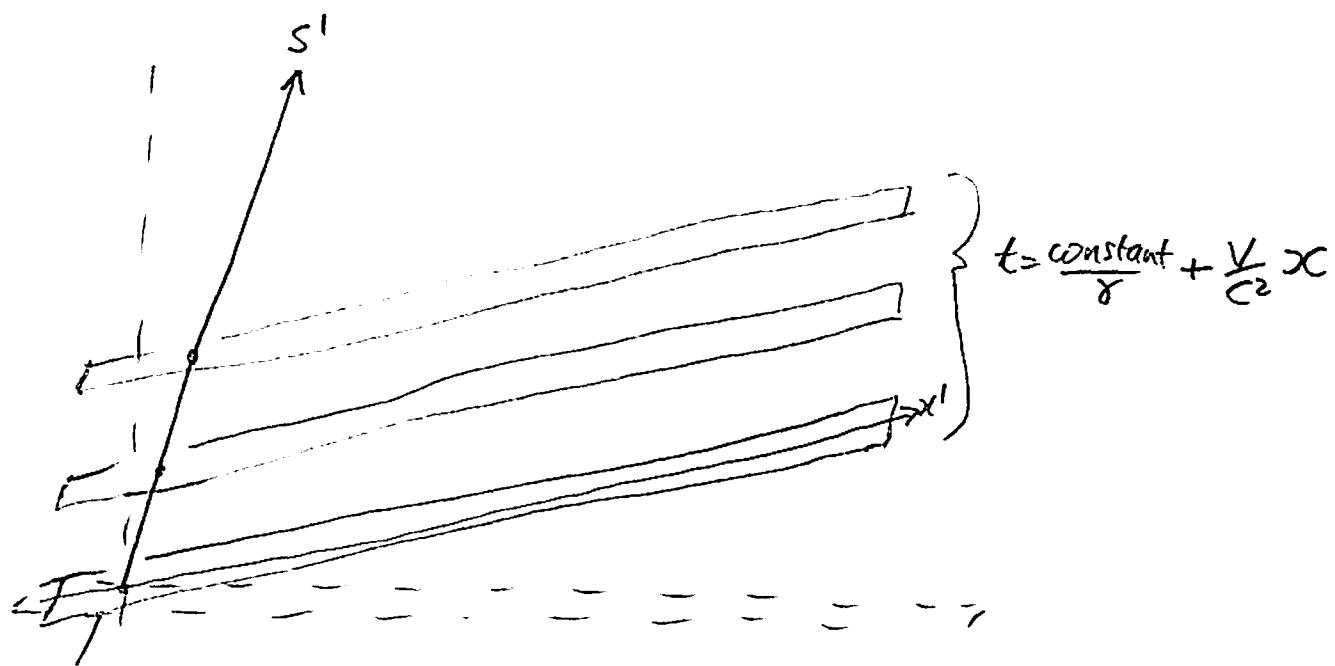
$$S(x, y, z, t)$$



Hypersurfaces of simultaneous events in $S'(x', y', z', t')$
satisfy $t' = \text{constant}$

$$\text{i.e. } \gamma(t - \frac{v}{c^2}x) = \text{constant}$$

$$\text{r.l. } t = \frac{\text{constant}}{\gamma} + \frac{v}{c^2}x$$



§18 ① - ②

Uniform
acceleration
 γ



No influence to
first order on
shape of body

Hence use Cartesian
coordinates as
spatial coordinates
of a uniformly
accelerated body

since Possible same for $+\gamma$, $-\gamma$
 effect (uniform dilation)

\uparrow \uparrow
 acc. in acc. in
 $+x$ $-x$
 direction direction

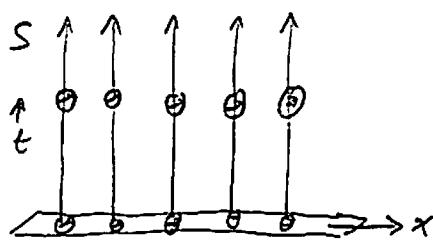
1.1. Effect(γ) = Effect($-\gamma$)

$$\text{Effect} = 1 + A\gamma + B\gamma^2 + C\gamma^3 + \dots$$

\uparrow $\underbrace{\quad}_{\text{vanishes}}$ $\underbrace{\quad}_{\text{2nd order small}}$
 No term is largest
 effect if $\gamma = 0$

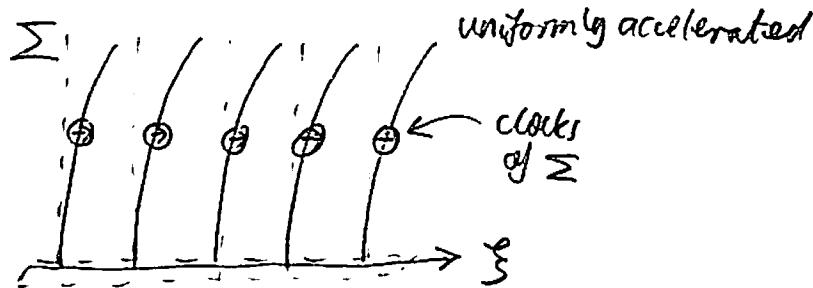
§18 ③

coordinate systems Σ , s , s' defined



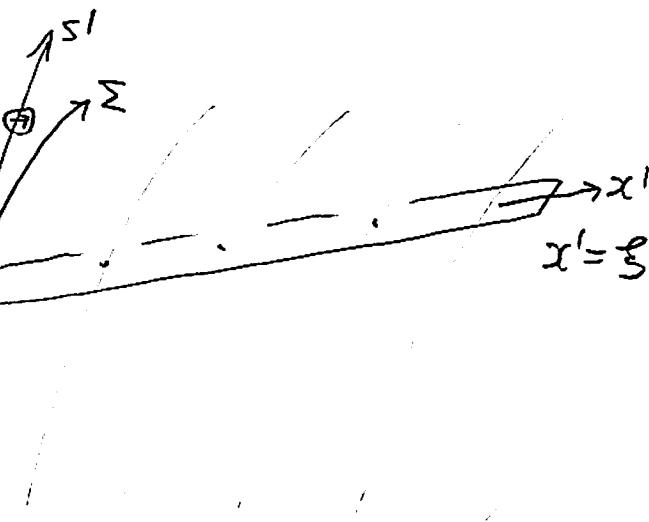
ordinary
inertial
system

inertial system
 s' momentarily
agrees
at spatial
origin



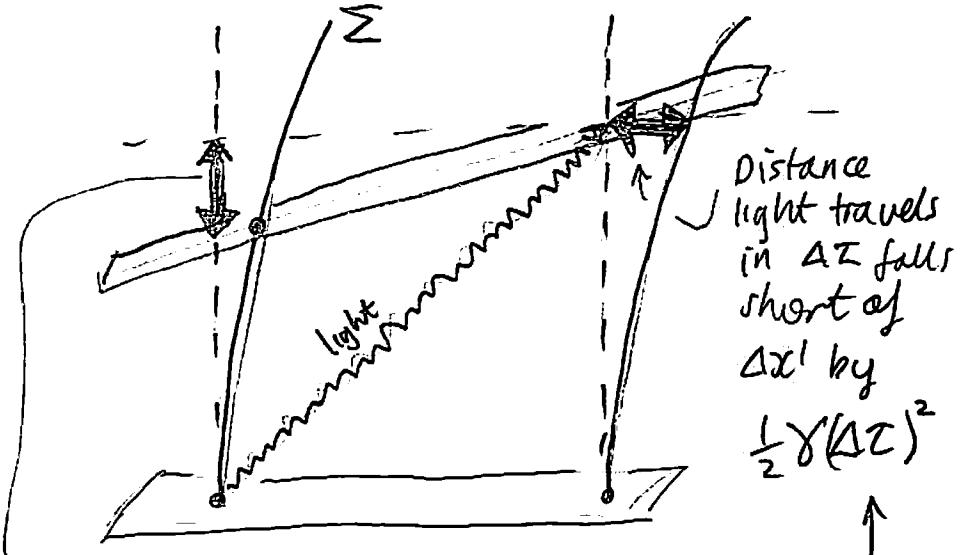
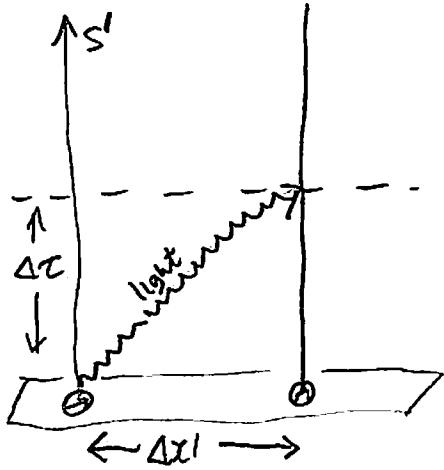
uniformly accelerated

clocks
of Σ



18.(4),(5)

Σ agrees with S'
for short time
intervals \Rightarrow speed of light in Σ
is universal constant C
- if measured over short
time intervals

over short time $\Delta\tau$ 

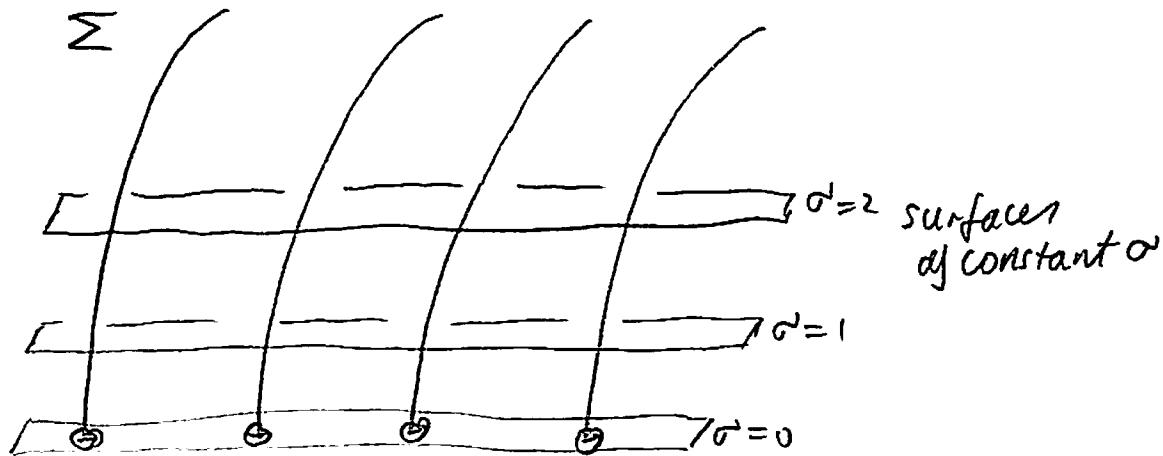
Light arrives early
in Σ by a relativity
of simultaneity

$$\text{term } \frac{\Delta x'(velocity)}{c} = \frac{\Delta x'}{c} \cdot \frac{1}{2} \gamma(\Delta\tau)^2 \leftarrow$$

Both
second
order
small
(in $\Delta\tau$)

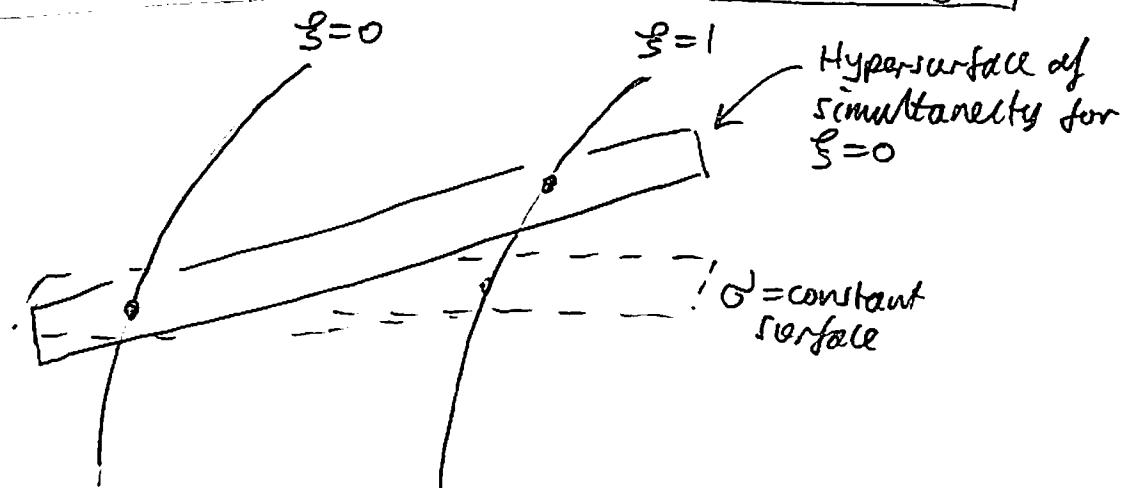
18.⑥

"Local time" = time read by clocks in Σ
 (initially set to 0 at $t=0$ of Σ)



18.⑦

Local time does not respect simultaneity



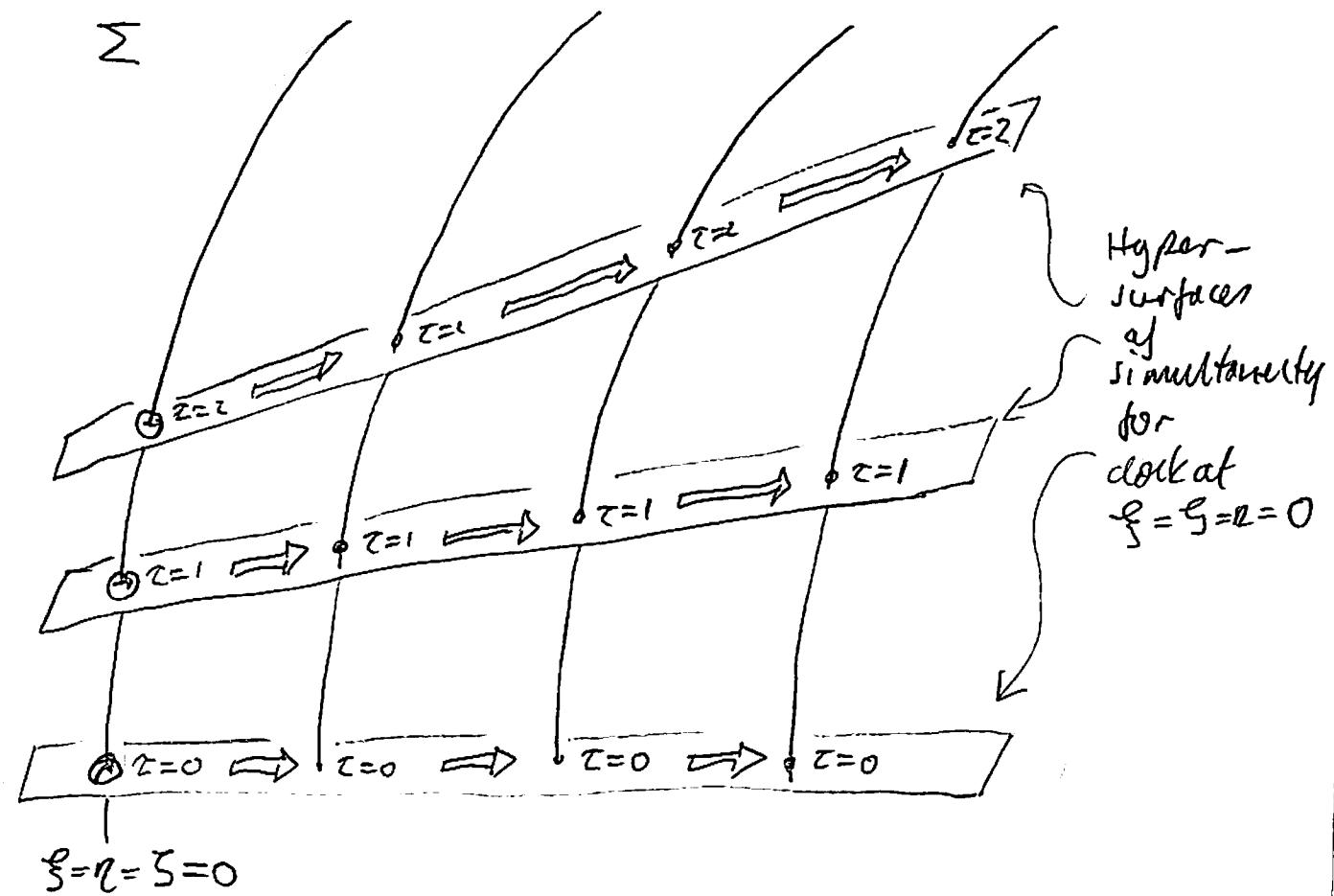
18.⑧

THE time

 Σ of the
system

Time read by clock at origin
of Σ propagated to all events
by Einstein synchrony

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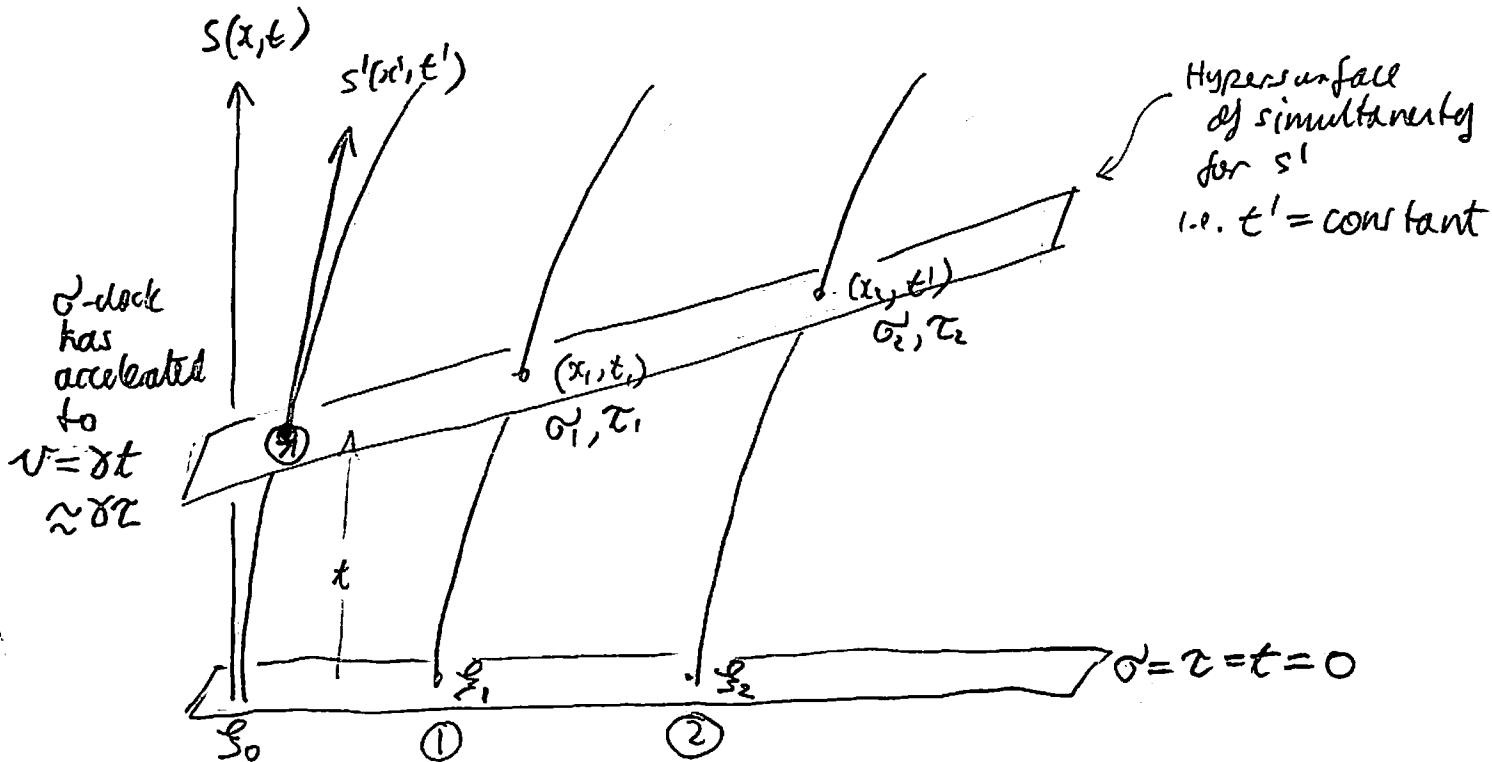


18. ⑨

Relation between
local time σ and
THT time Σ

$$\sigma = \Sigma \left(1 + \frac{\gamma \xi}{c^2} \right)$$

for small times after clock set
(i.e. drop all terms in $(\text{time})^2$)



condition for hypersurface: $t' = \gamma (x - \frac{v}{c^2} x) = \text{constant}$

$$\text{i.e. } t'_1 = t'_2 \quad t_1 - \frac{v}{c^2} x_1 = t_2 - \frac{v}{c^2} x_2$$

$$\therefore t_2 - t_1 = \frac{v}{c^2} (x_2 - x_1)$$

$$\begin{array}{l} \downarrow \\ v \approx \Delta \Sigma \quad t_1 \approx \sigma_1, t_2 \approx \sigma_2 \\ x_2 - x_1 \approx x'_2 - x'_1 \approx \xi_2 - \xi_1 \end{array}$$

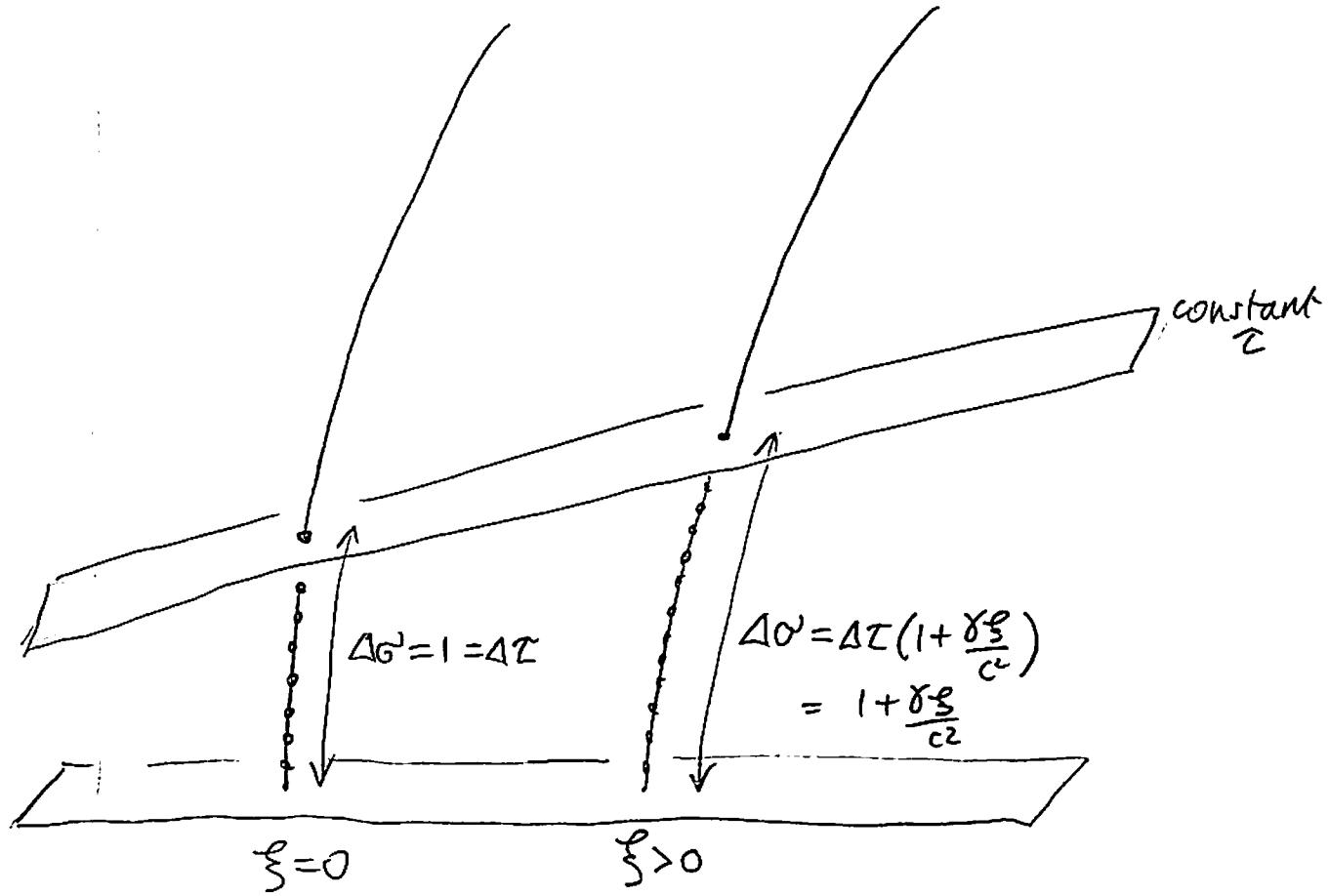
$$\sigma_2 - \sigma_1 = \frac{\Delta \Sigma}{c^2} (\xi_2 - \xi_1)$$

$$\begin{array}{l} \downarrow \\ \text{Set point ① to origin: } \xi_1 = 0, \xi_2 = \xi \text{ etc} \\ \sigma_1 = \Sigma \end{array}$$

$$\sigma = \Sigma \left[1 + \frac{\gamma \xi}{c^2} \right]$$

19. Slowing of clocks in uniformly accelerating frame Σ / Homogeneous gravitational field

← Deeper in field. \Rightarrow
smaller $\bar{\Phi} = \gamma \xi$



Energy in Φ
which has grown Δt

$$m = \frac{c^2}{2} E$$

mass E

growth Δt

growth Δt

locally measured

measured energy

Adiabatically

locally measured

$$\frac{\Phi}{E}$$

by

locality

measured

energy

is this

$$\text{by } \frac{\partial \Phi}{\partial t} = \frac{c^2}{2}$$

quantities Φ, E

locally measured

must adjust

to

time

$t + \Delta t$

time

time

time

field energy

$$0 = m \alpha \left(\frac{c^2}{2} + 1 \right) \int \frac{d\Phi}{dt} + m \alpha^2 n \left(\frac{c^2}{2} \right) d\omega$$

on matter

work done

Rate

time

Energy

time

\therefore light bent by gravity

$$\text{of light} = C \left(\frac{c^2}{2} + 1 \right)$$

"of non-shielding
of matter"

$$*\bar{E} \times \bar{\Delta} = \frac{2P}{\bar{H}\bar{P}} \left(\frac{c^2}{2} + 1 \right)$$

$$\frac{\partial}{\partial P} \downarrow \quad \frac{\partial}{\partial P} \downarrow \\ = \bar{n} \left(\frac{c^2}{2} + 1 \right) = \frac{2P}{\bar{H}\bar{P}}$$

local time

$(2, n, 5, 2)$

equations in

Maxwell's

"for static
& stationary"

$$*\bar{E} \times \bar{\Delta} = \frac{2P}{\bar{H}\bar{P}} \quad *\bar{H} \times \bar{\Delta} = \left(\frac{2P}{\bar{E}\bar{P}} + 3\bar{n} \right) \frac{1}{2}$$

local time

$(2, n, 5, 6)$

equations in

Maxwell's

$$\left(\frac{c^2}{2} + 1 \right) = \bar{n} \quad \left(\frac{c^2}{2} + 1 \right) \bar{H} = \bar{H} \quad \left(\frac{c^2}{2} + 1 \right) \bar{E} = \bar{E}$$

$$\bar{E} \times \bar{\Delta} = \frac{2P}{\bar{H}\bar{P}} \quad \bar{H} \times \bar{\Delta} = \left(\frac{2P}{\bar{E}\bar{P}} + \bar{n} \right) \frac{1}{2}$$

$(7, n, 3, 7)$

equations in

Maxwell's

Properties of Result

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