P8.1) A strong absorption band in the infrared region of the electromagnetic spectrum is observed at $\tilde{\nu} = 2170 \text{ cm}^{-1}$ for ${}^{12}\text{C}{}^{16}\text{O}$. Assuming that the harmonic potential applies, calculate the fundamental frequency ν in units of inverse seconds, the vibrational period in seconds, and the zero point energy for the molecule in joules and electron-volts.

$$\nu = \bar{\nu}c = 2170 \text{ cm}^{-1} \times 3.00 \times 10^{10} \text{ cm} \text{ s}^{-1} = 6.51 \times 10^{13} \text{ s}^{-1}$$

$$T = \frac{1}{\nu} = \frac{1}{6.51 \times 10^{13} \text{ s}^{-1}} = 1.54 \times 10^{-14} \text{ s}$$

$$E = \frac{1}{2}h\nu = \frac{1}{2} \times 6.626 \times 10^{-34} \text{ Js} \times 6.51 \times 10^{13} \text{ s}^{-1} = 2.16 \times 10^{-20} \text{ J} \times \frac{6.241 \times 10^{18} \text{ eV}}{\text{J}} = 0.134 \text{ eV}$$

P8.7

A measurement of the vibrational energy levels of ¹²C¹⁶O gives the relationship $\tilde{\nu}(n) = 2170.21 \left(n + \frac{1}{2}\right) \text{ cm}^{-1} - 13.461 \left(n + \frac{1}{2}\right)^2 \text{ cm}^{-1}$ where *n* is the vibrational quantum

number. The fundamental vibrational frequency is $\tilde{\nu}_0 = 2170.21 \text{ cm}^{-1}$. From these data, calculate the depth D_e of the Morse potential for ${}^{12}\text{C}{}^{16}\text{O}$. Calculate the bond energy of the molecule.

We convert the expression to one in terms of energy

$$E(n) = 2170.21 \times hc\left(n + \frac{1}{2}\right) - 13.461 \times hc\left(n + \frac{1}{2}\right)^2$$

and compare it with Equation (19.5) for the Morse potential.

$$E_n = h\nu \left(n + \frac{1}{2}\right) - \frac{\left(h\nu\right)^2}{4D_e} \left(n + \frac{1}{2}\right)^2$$

from which we see that

$$13.461 \times hc = \frac{(h\nu)^2}{4D_e}$$

$$D_e = \frac{h\nu^2}{4 \times 13.461 \times c} = \frac{hc\tilde{\nu}^2}{4 \times 13.461}$$

$$= \frac{6.626 \times 10^{-34} \,\mathrm{J}\,\mathrm{s} \times 2.998 \times 10^{10} \,\mathrm{cm}\,\mathrm{s}^{-1} \times \left(2170 \,\mathrm{cm}^{-1}\right)^2}{4 \times 13.461} = 1.738 \times 10^{-18} \,\mathrm{J}$$

$$D_0 = D_e - \frac{1}{2} hc\tilde{\nu}_0 = 1.738 \times 10^{-18} \,\mathrm{J} - \frac{1}{2} \times 6.626 \times 10^{-34} \,\mathrm{J}\,\mathrm{s} \times 2.998 \times 10^{10} \,\mathrm{cm}\,\mathrm{s}^{-1} \times 2170 \,\mathrm{cm}^{-1}$$

$$= 1.717 \times 10^{-18} \,\mathrm{J}$$

The bond energy per mole is $D_0 N_A = 1.034 \times 10^3 \text{ kJ mol}^{-1}$.

P8.16

Calculate the moment of inertia, the magnitude of the rotational angular momentum, and the energy in the J = 1 rotational state for ${}^{1}\text{H}_{2}$ in which the bond length of ${}^{1}\text{H}_{2}$ is 74.6 pm. The atomic mass of ${}^{1}\text{H}$ is 1.007825 amu.

$$J = \mu r_0^2 = \frac{1.007825^2 \text{ amu}}{2 \times 1.007825} \times 1.6605402 \times 10^{-27} \text{kg amu}^{-1} \times (74.6 \times 10^{-12} \text{ m})^2$$

= 4.657×10⁻⁴⁸ kg m²
 $|J| = \sqrt{J(J+1)}\hbar = \sqrt{2} \times 1.0554 \times 10^{-34} \text{ J s} = 1.4914 \times 10^{-34} \text{ J s}$
= 1.49×10⁻³⁴ kg m s⁻²
 $E_J = \frac{J(J+1)\hbar^2}{2I} = \frac{2 \times (1.0554 \times 10^{-34} \text{ J s})^2}{2 \times 4.657 \times 10^{-48} \text{ kg m}^2} = 2.39 \times 10^{-21} \text{ J}$

(0) i) For ICN N=#atoms=3, and, since it is a linear molecule $n_v = \#$ vibrational degrees of freedom (normal modes of vibration)= 3N-5 = 4.

ii) For CH₄, N=5, and, since it is a nonlinear molecule, $n_v = 3N-6 = 9$.

(1) From class notes, we know: $\vec{x}_1 = \vec{X} - \frac{m_2}{M}\vec{r}$, $\vec{x}_2 = \vec{X} + \frac{m_1}{M}\vec{r}$. Thus:

$$\begin{split} T &= \frac{1}{2} m_1 \vec{x}_1^2 + \frac{1}{2} m_2 \vec{x}_2^2 \\ &= \frac{1}{2} m_1 \left(\dot{\vec{X}} - \frac{m_2}{M} \dot{\vec{r}} \right)^2 + \frac{1}{2} m_2 \left(\dot{\vec{X}} + \frac{m_1}{M} \dot{\vec{r}} \right)^2 \\ &= \frac{1}{2} m_1 \left(\dot{\vec{X}}^2 - \frac{2m_2}{M} \dot{\vec{r}} \cdot \dot{\vec{X}} + \frac{m_2^2}{M^2} \dot{\vec{r}}^2 \right) + \frac{1}{2} m_2 \left(\dot{\vec{X}}^2 + \frac{2m_1}{M} \dot{\vec{r}} \cdot \dot{\vec{X}} + \frac{m_1^2}{M^2} \dot{\vec{r}}^2 \right) \\ &= \frac{1}{2} M \vec{\vec{X}}^2 + \frac{1}{2} \mu \vec{\vec{r}}^2 \quad , QED. \end{split}$$

Note: Here $\vec{a} \cdot \vec{b}$ is the dot product between two 3D vectors \vec{a}, \vec{b} , and $\vec{a}^2 = \vec{a} \cdot \vec{a}$.

(2) a). Assume the desired energy eigenfunction can be written in the product form $\psi(r, \theta, \phi) = R_l(r)Y_{lm}(\theta, \phi)$ where Y_{lm} is the standard spherical harmonic function, and $R_l(r)$ is a radial function to be determined. [Note: We anticipate that it will depend on the value of *l* in the companion spherical harmonic.]

Note the following useful version of the 3D Laplacian operator:

$$\nabla^2 = \frac{1}{r} \frac{\partial^2 (r)}{\partial r^2} + \frac{\hat{\Lambda}^2}{r^2} ,$$

where the "Legendrian"

$$\hat{\Lambda}^2 = \frac{1}{\sin\theta} \frac{\partial}{\partial\theta} (\sin\theta \frac{\partial}{\partial\theta}) + \frac{1}{\sin^2\theta} \frac{\partial^2}{\partial\phi^2}$$

operates (obviously) on the angles (θ, ϕ) . Substituting all this into the Schrödinger Eq.:

$$\left[\frac{-\hbar^2}{2\mu}\left\{\frac{1}{r}\frac{\partial^2(r)}{\partial r^2} + \frac{\hat{\Lambda}^2}{r^2}\right\} + v(r)\right]R_l(r)Y_{lm}(\theta,\phi) = ER_l(r)Y_{lm}(\theta,\phi) \quad [A1]$$

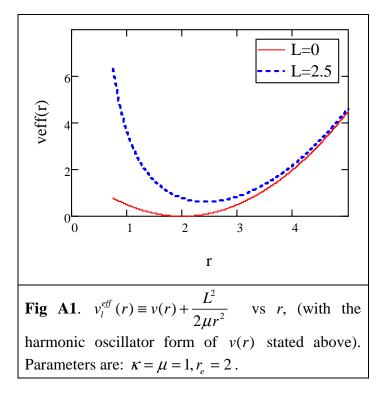
Now, using the fact that the spherical harmonics are eigenfunction of the Legendrian, namely $\hat{\Lambda}^2 Y_{lm}(\theta, \phi) = -l(l+1)Y_{lm}(\theta, \phi)$, Eq. A1 implies the radial wave equation:

$$\frac{-\hbar^2}{2\mu}\frac{1}{r}\frac{\partial^2(rR_l(r))}{\partial r^2} + \left(v(r) + \frac{\hbar^2l(l+1)}{2\mu r^2}\right)R_l(r) = ER_l(r)$$

Finally, defining the reduced radial wavefunction $\chi_l(r)$ according to $rR_l(r) = \chi_l(r)$, we obtain a "standard" 1D Schrödinger Eq. for the reduced radial wave function:

$$\frac{-\hbar^2}{2\mu}\frac{\partial^2 \chi_l(r)}{\partial r^2} + v_l^{eff}(r)\chi_l(r) = E\chi_l(r),$$

with the effective radial potential $v_l^{eff}(r) \equiv v(r) + \frac{\hbar^2 l(l+1)}{2\mu r^2}$, *QED*.



b) Adding the repulsive "centrifugal" potential term $\frac{\hbar^2 l(l+1)}{2\mu r^2}$ to $v(r) = \frac{1}{2}\kappa(r-r_{eq})^2$ pushes the equilibrium position of the potential well $v_l^{eff}(r)$ out to a larger value; cf. Fig. A1. Thus, increasing *l* increases the equilibrium separation of the two particles in the diatomic.