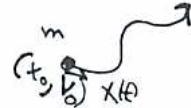


Chem. 1410
27 Aug. 07



Features of Classical Mechanics :

(entral purpose : find trajectory $x(t)$)

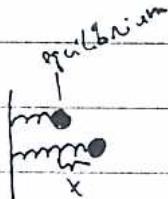
Basic Equation of motion (Newton)

$$F(x) = m \ddot{x}(t) = m \dot{v}(t) = \ddot{p}(t)$$

Example (1) Free particle $F=0 \Rightarrow x(t) = \frac{p_0}{m} t + x_0$

(2) $F=\text{const}$; e.g. $\downarrow g$, $F=mg$; now $x(t) = x_0 + \frac{p_0}{m} t + \frac{1}{2} \frac{F}{m} t^2$

(3) Linear restoring force (Harmonic Oscillator) $F(x) = -kx$; Now $m\ddot{x} = -kx$



$$\text{Or: } \ddot{x}_H = -\omega^2 x(t) \quad \omega \equiv \sqrt{k/m}$$

$$\text{Thus: } x(t) = \alpha \cos \omega t + \beta \sin \omega t = x_0 \cos \omega t + \frac{p_0}{m \omega} \sin \omega t \quad \leftarrow (x_0, p_0) \Leftrightarrow A, \delta$$

$$= A \sin(\omega t + \delta) = A [\sin \omega t \cos \delta + \cos \omega t \sin \delta]$$

Features of H.O. motion: $A = \text{amplitude}$ (magnitude of maximum displacement)

$$T = \text{period of oscillation}; \quad \omega T = 2\pi \quad \text{or} \quad T = \frac{2\pi}{\omega} = \frac{1}{\nu} \quad \leftarrow \text{frequency}$$

Principle of Energy Conservation: define potential energy $V(x) = - \int_0^x F(x') dx'$

\uparrow
Lower limit is arbitrary

thus, for $F=0$, $V(x)=0$

$$F = \text{const}, \quad V(x) = -Fx$$

$$F = -kx, \quad V(x) = \frac{1}{2} kx^2 \quad \textcircled{1}$$

Now consider: $E = \text{total energy} = \underbrace{\frac{1}{2}mV(t)^2}_{\text{Kinetic E}} + \underbrace{V(x(t))}_{\text{Potential E}}$

For free particle, $E = \frac{1}{2}\frac{P_0^2}{m} = \frac{1}{2}mV_0^2 = \text{const.}$

For constant force, $\frac{d}{dt}E = \frac{1}{2}m\left[\frac{P_0}{m} + \frac{F}{m}t\right] - F\left[x_0 + \frac{P_0}{m}t + \frac{1}{2}\frac{F}{m}t^2\right] = \frac{P_0^2}{2m} - Fx_0 = \text{constant}$

For H.O., $E = \frac{1}{2}m\left[A\omega\cos(\omega t + \delta)\right]^2 + \frac{1}{2}m\omega^2\left[A\sin(\omega t + \delta)\right]^2 = \frac{1}{2}m\omega^2A^2 = \frac{1}{2}kA^2$

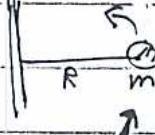
These are examples of Principle of Conservation of Energy (E)

Some basic characteristics of classical mechanics: ① Arbitrary E's allowed.

② Trajectories precisely specified

$$\omega, \vec{T} \uparrow \vec{x} \\ w = d\theta/dt$$

Classical Mechanical description of Rotational Motion.



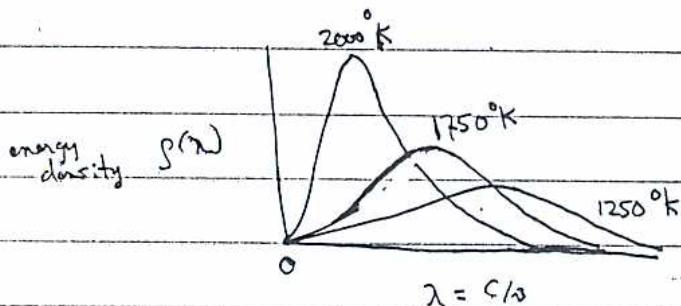
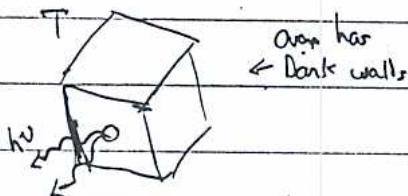
linear	rotational	$ T $
m	$I (= mR^2)$, moment of inertia	
v	ω , angular velocity	"Newton's Eq"
P	\vec{T} , " momentum	$\dot{T} = \vec{T}$
F	T , torque	

For constant torque: $\theta(t) = \theta_0 + \omega_0 t + \frac{1}{2}\frac{T}{I}t^2$

(rotational) KE = $\frac{1}{2}I\dot{\theta}(t)^2 = \frac{T^2t^2}{2I}$ if $\omega_0 = 0$

Failures of Classical Physics [~1900]

(I) Blackbody Radiation



Some

Features: (1) $\lambda_{\max} T = \text{const}$, where λ_{\max} = wavelength at which intensity is max. [Wien's Law]

$$(2) \int_0^{\infty} d\lambda g(\lambda) = \sigma T^4 \quad [\text{Stefan's Law}]$$

$$5.67 \times 10^{-8} \frac{\text{W}}{\text{m}^2 \text{K}^4}$$

Theories [Raleigh-Jeans, Planck...]

Aside: Classical Equilibrium Thm.: At thermal equilibrium [$T = \text{temp}, \Omega$], energy is partitioned into each molecule according to: $\frac{1}{2} k_B T$ in KE for each degree of freedom; and...

$$\frac{1}{2} k_B T \text{ in PE for each vibrational degree of freedom}$$

$$\rightarrow k_B = 1.38 \times 10^{-23} \text{ J/K}$$

[Thus, \rightarrow for a 1D H.O., $E = k_B T$]

(3)

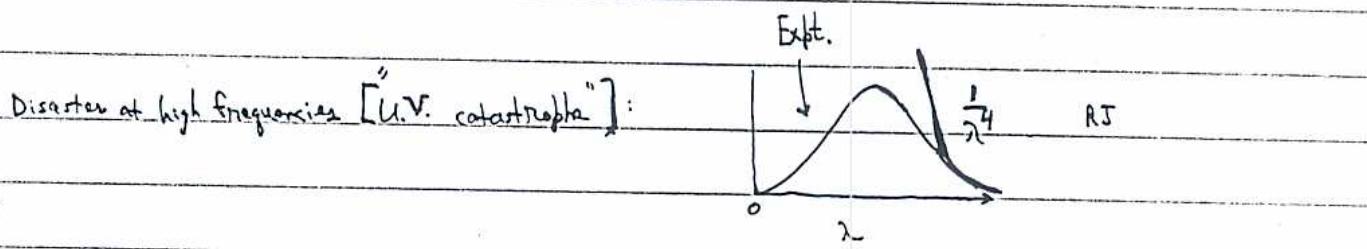
Now, Rayleigh-Jeans... Assumed: (i) electromagnetic field = collection of h.o.s; one for each possible frequency of light

(ii) classical equipartition thm. [Each mode gets kT energy, indep. of frequency]

Then: (RJ) $\rho_N = \frac{8\pi}{\lambda^4} \cdot k_B T$

energy/mode
 \downarrow
 density of E.M.
 modes w/ wavelength λ
 in the cavity [per unit volume]

← energy
 (wavelength)(volume) } "energy density"



Planck: for each field oscillator (frequency ν), restrict allowed energies to: $\frac{h\nu}{2\pi}$ [Quantization]

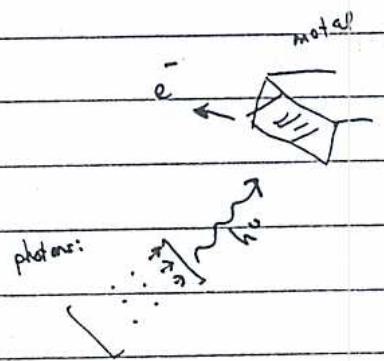
Analysis then shows: $E(\nu) = \frac{h\nu}{h\nu/k_B T - 1}$ [$\nu = \frac{c}{\lambda}$] ↑ high frequency
modes are not excited! [no photons in those modes]

Thus:

(Planck) $\boxed{\rho_N = \frac{8\pi hc}{\lambda^5} \frac{1}{e^{hc/\lambda k_B T} - 1}}$; $c = \text{speed of light} = 3 \times 10^8 \text{ cm/sec}$

$h = \text{Planck's const.} = 6.625 \times 10^{-34} \text{ erg-sec}$

Photo-electron effect [Einstein]



Expt: (1) no e^- for $\nu < \nu_0$

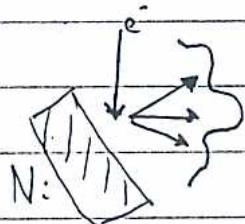
$$(2) \text{ for } \nu > \nu_0, e^- \text{ off w k.E. } \propto (\nu - \nu_0) \Rightarrow \frac{1}{2}mv^2 = h\nu - h\nu_0$$

$\frac{h\nu}{\text{work funcn.}}$

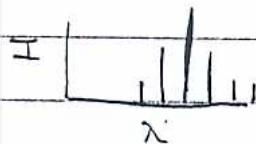
(3) e^- off instantaneously for $\nu > \nu_0$, even at low intensities

of metal

Davison-Germer Expt. [1925] : Interference [Diffraction] in electron-surface scattering \Rightarrow Matter is wavelike

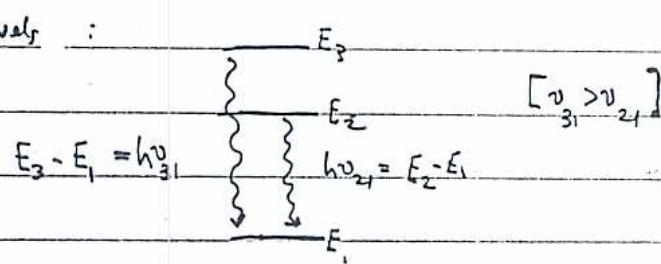


$$\text{De Broglie: } \lambda = h/p \text{ for any particle}$$



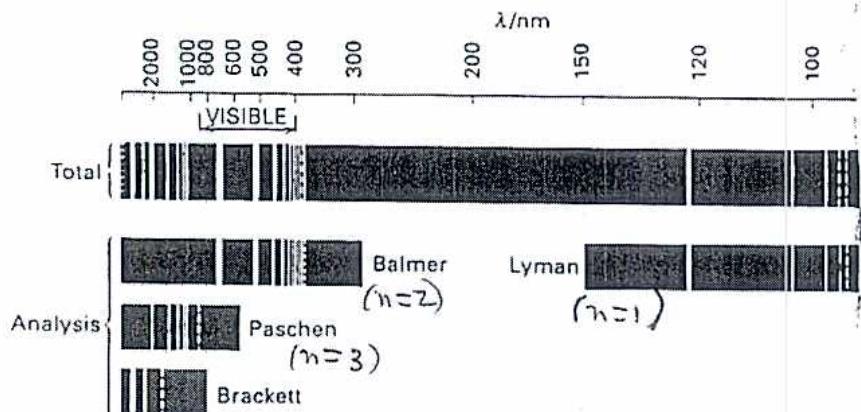
See Fig. 11.5-11.6 in Atkins

Atomic + Molecular Spectra - always composed of discrete lines



Suggests quantization of atomic/molecular E levels

pass electric discharge thru $H_2(g)$ $\rightarrow H\cdot + H\cdot$
 consider emission spectra of the hydrogen atom [Balmer, Lyman...]
 highly organized!



5th [Atkins]
 5th Edition [\rightarrow 13.1 in 6th Ed.]
 [8th Ed. p. 10.1]

13.2 The spectrum of atomic hydrogen.
 The observed spectrum and its resolution into overlapping series are shown. Note that the Balmer series lies in the visible region

All lines in this emission spectrum are consistent w/ the following E-level diagram

$$E_n = -13.6 \frac{eV}{n^2} ; n=1, 2, 3, \dots, \infty$$

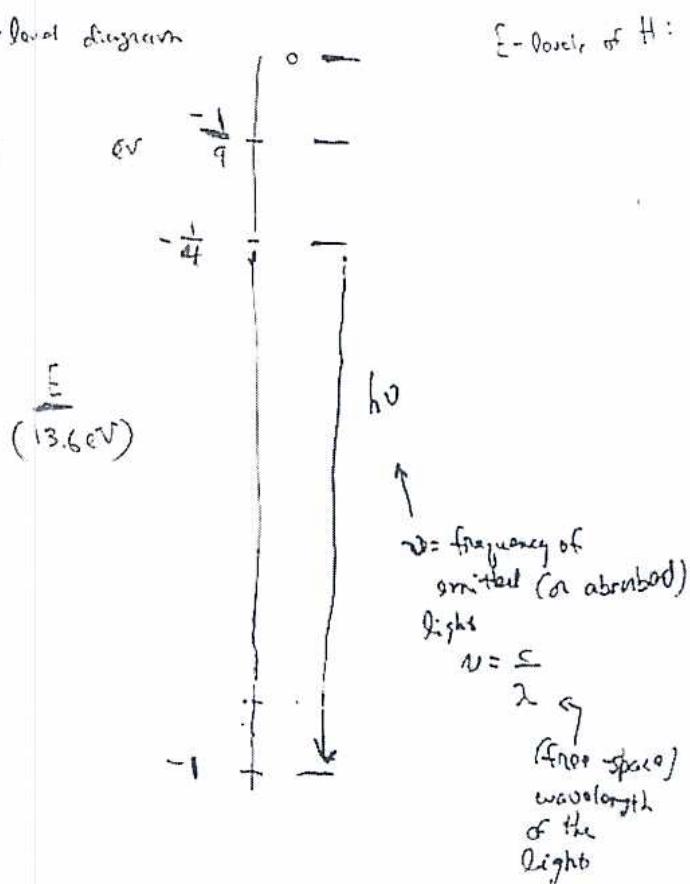
Quantitatively, for any line in the spectrum

$$h\nu = 13.6 eV \left(\frac{1}{n^2} - \frac{1}{m^2} \right)$$

$m, n = 1, 2, 3$
and $m > n$

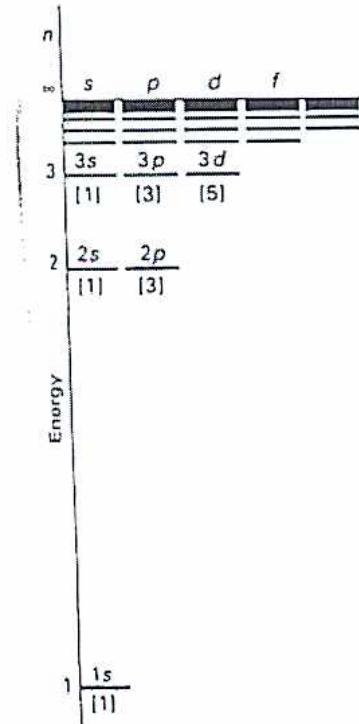
J. Rydberg [1890]

$$[1 eV = 1.6 \times 10^{12} \text{ ergs}]$$



Energy levels, states & degeneracies

In some systems, there are two or more distinguishable states w/ the same energy, e.g. in the H-atom:



13.8 The energy levels of the hydrogen atom showing the subshells and (in square brackets) the numbers of orbitals in each subshell. All orbitals of a given shell have the same energy in hydrogenic atoms.

Note also: spin degeneracy = 2 ($\uparrow\downarrow$)
for each spatial orbital

so total degeneracy of energy levels:

	spatial degeneracy		spin degeneracy	total
(i) $n=1$	1		2	2
	= 2		4	8
	= 3		9	18

generally: for n principle quantum #

$$n^2$$

$$2 \times n^2$$

of electrons
that can
have

energy E_n