

# STAT MECH – PHASE TRANSITIONS – Illustrated with Ising model

Ising model – N spins on square lattice

$$E_v = -\sum_i^N H \mu s_i - J \sum_{i,j}' s_i s_j$$

$$s_i = \pm 1$$

$J > 0$ , favorable for spins to align  
 '  $\Rightarrow$  nearest neighbor only

Magnetization

$$\langle M \rangle = \sum_{i=1}^N \mu s_i$$

$T_c =$  curie temp

oftentimes written  $i < j$  or with a factor of  $\frac{1}{2}$

Ph.D. thesis of E. Ising (1925)  
 solved 1D

2D solved by Onsager (1944)

$J > 0$  ferromagnetic

$J < 0$  antiferromagnetic

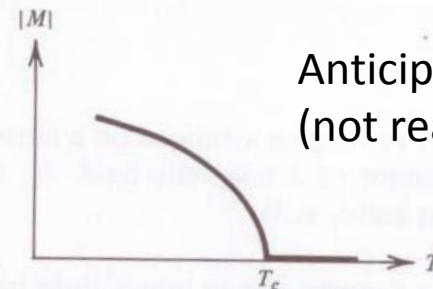


Fig. 5.2. Spontaneous magnetization.

(from Chandler)

suppose  $H = 0$

lowest energy =  $-DNJ$  ( $D$  = dimension)

$$(s_{N+1} = s_1)$$

$$\left. \begin{array}{l} 1D \quad 2 \\ 2D \quad 4 \\ 3D \quad 6 \end{array} \right\} \text{ \# nearest neighbor spins}$$

$$Q = \sum_v e^{-\beta E_v} = \sum_{s_{1=\pm 1}} \dots \sum_{s_{N=\pm 1}} e^{\left[ \beta \mu H \sum_{i=1}^N s_i + \beta J \sum_{ij} s_i s_j \right]}$$

1-D lattice

the interaction energy involves a single sum

$$-J \sum_{i=1}^N s_i s_{i+1}, \text{ with periodic boundary conditions}$$

$$H = 0 \rightarrow Q = 2N \left\{ \left[ \cosh(\beta J) \right]^N + \left[ \sinh(\beta J) \right]^N \right\} \\ \approx \left[ 2 \cosh(\beta J) \right]^N$$

2 spin-linear chain

	$s_1 s_2$	
$e^{\beta J(s_1 s_2 + s_2 s_3)}$	1 1	$e^{\beta(2)}$
	1 -1	$e^{-\beta(2)}$
	-1 1	$e^{-\beta(2)}$
	-1 -1	$e^{\beta(2)}$

No spontaneous magnetization in 1D

$$\begin{array}{ccccccc} \uparrow & \uparrow & \dots & \uparrow & \dots & \uparrow & \\ 1 & 2 & & & & N & \end{array} \quad E = -NJ$$

flip part of the chain

$$\begin{array}{ccccccc} \uparrow & \uparrow & \dots & \downarrow & \dots & \downarrow & \\ 1 & 2 & & & & N & \end{array} \quad E = (-N + 4)J \quad \Bigg| \quad \text{recall PBC}$$

small energy difference between a magnetic and non-magnetic systems

So, for large  $N$ , even at very low  $T$ , there is no net magnetization

(Actually, even if it were not for this issue, there is the problem of the degenerate ground state.)

For a 2-D spin system

...	↑	↑	↑	↑	...	the cost for flipping one half the spins energy goes as $\sqrt{N}$ when starting with fully aligned spin system.
	↑	↑	↑	↑		
...	↓	↓	↓	↓	...	
	↓	↓	↓	↓		

Onsager solved analytically the 2D Ising problem

$$H = 0, \quad Q = [2 \cosh(2\beta J) e^I]^N$$

$$I = \frac{1}{2\pi} \int_0^\pi d\varphi \ln \left\{ \frac{1}{2} [1 + (1 - \kappa^2 \sin^2 \varphi)]^{1/2} \right\}$$

$$\kappa = 2 \sinh(2\beta J) / \cosh^2(2\beta J)$$

depending on definition of  $J$ , need to divide by 2 to prevent double counting

Spontaneous magnetism for  $T < T_c = \frac{2.269J}{k}$

$$\sinh\left(\frac{2J}{kT_c}\right) = 1$$

near  $T_c$

$$C/N \sim \frac{8k}{\pi} \beta J \ln \left| \frac{1}{T - T_c} \right|$$

$$M/N \sim \text{constant} (T_c - T)^{1/8}, T < T_c$$

Corrected exponent

3D from numerical solution

near  $T_c$

$$C/N \sim \frac{1}{(T - T_c)^{1/8}}$$

$$M/N \sim (T_c - T)^{0.313}, T < T_c$$

critical exponents  
depend on dimensionality

## Connection between Ising and Lattice-Gas Models

Lattice gas

cell occupied  $n_i = 1$

cell empty  $n_i = 0$

particles in adjacent cells energy =  $-\varepsilon$

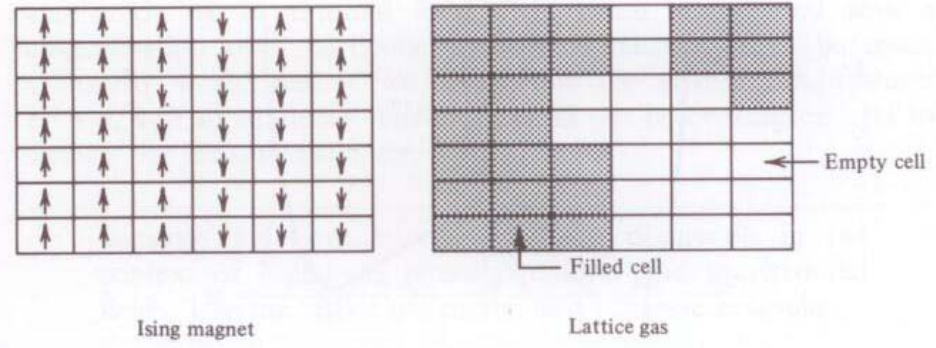


Fig. 5.3 from Chandler

$$E = -\varepsilon \sum_{i,j} n_i n_j$$

$$\Xi = \sum_{n_1, \dots, n_M} e^{\{\beta\mu \sum n_i + \beta\varepsilon \sum' n_i n_j\}}$$

sums over # cells

Lattice gas model is isomorphic with Ising model

$$s_i \rightarrow 2n_i - 1$$

spin up  $\rightarrow$  occ site

spin down  $\rightarrow$  empty site

mag field  $\rightarrow$  chemical potential

$$J \rightarrow \varepsilon / 4$$

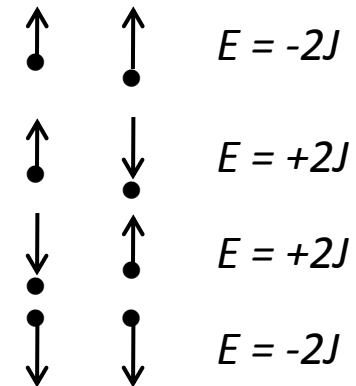
Ising model – **broken symmetry**

all spins up and all spins down same energy

$$\langle M \rangle = \frac{1}{Q} \sum_v \left( \sum_{i=1}^N \mu s_i \right) e^{-\beta E_v}$$

for every configuration with net up spin  
there is another with equal down spin

so why should we ever see magnetization?



$$Q = 2e^{2\beta J} + 2e^{-2\beta J}$$

$$\langle M \rangle = \frac{\mu [2e^{2\beta J} + 0e^{2\beta J} - 2e^{-2\beta J}]}{Q}$$

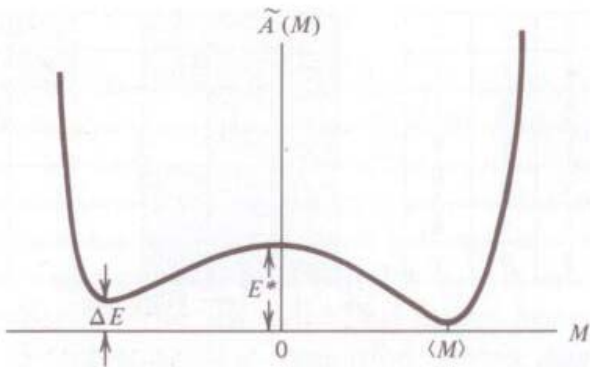


fig. 5.4 from Chandler

Fluctuations between  $\langle M \rangle$  and  $-\langle M \rangle$   
vanishingly small for large system

add weak field  $\rightarrow$  magnetize  
remove field  $\rightarrow$  spontaneous fluctuations  
do not destroy broken symmetry  
 $M$  can be viewed as an order parameter

$$\tilde{Q}(M) = \sum_{\nu} \Delta(M - M_{\nu}) e^{-\beta E_{\nu}}$$

$$\begin{aligned} \Delta(M - M_{\nu}) &= 1, & M &= M_{\nu} \\ &= 0, & M &\neq M_{\nu} \end{aligned}$$

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Question boils down to:

Does the system have long-range order?



## Pair correlation function

$$c_{ij} = \langle s_i s_j \rangle - \langle s_i \rangle \langle s_j \rangle$$

= 0 if spin at  $i$  uncorrelated with that of  $j$ .

$$\sum_{j=2}^N c_{1j} = \text{\# of spins correlated with spin 1.}$$

$$\text{Susceptibility } \chi = \frac{1}{N} \left( \frac{\partial \langle M \rangle}{\partial (\beta H)} \right)_{\beta}$$

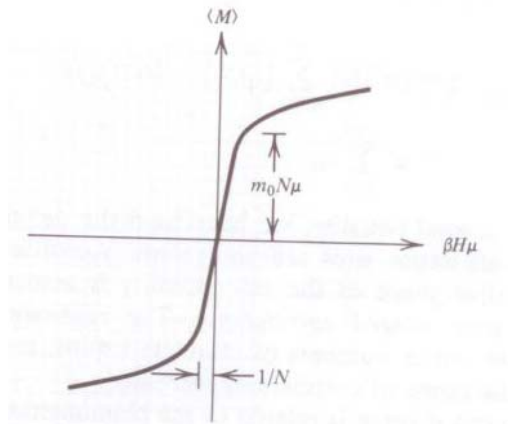
$$\begin{aligned} \delta M &= M - \langle M \rangle \\ &= \mu \sum_{i=1}^N [s_i - \langle s_i \rangle] \end{aligned}$$

$$\chi = \frac{1}{N} \langle (\delta M)^2 \rangle$$

$$\chi = \frac{\mu^2}{N} \sum_{i,j} [\langle s_i s_j \rangle - \langle s_i \rangle \langle s_j \rangle] = \mu^2 \sum_{i=1}^N c_{1j}$$

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using fact all sites  
are equivalent



$\langle M \rangle$  for finite  $N$   
fig. from Chandler

$$T < T_c$$

$\langle M \rangle$  jumps at  $H = 0$ , so  $\frac{\partial \langle M \rangle}{\partial H}$  diverges.

suppose  $N$  large but not  $\infty$ ,  $H = 0$ , and  $T < T_c$

$$\langle s_i \rangle = 0$$

$$\sum_{j=1}^N \langle s_i s_j \rangle = N m_0$$

$$\chi = N m_0 \mu^2$$

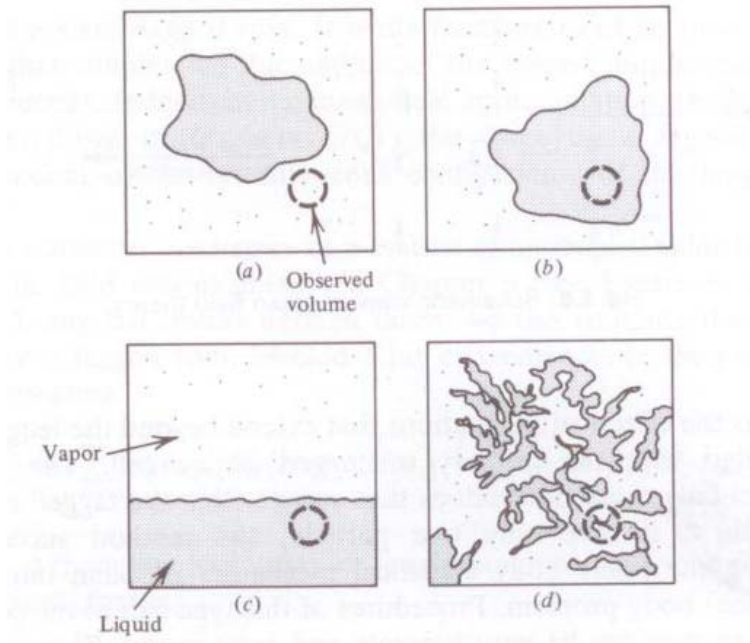
(0  $\Rightarrow$  zero field)

Under the stated conditions  
the choice of  $s_1$  (+ or -)  
biases the other spins

divergence of  $\chi \Leftrightarrow$  macroscopic fluctuations

quenched by applying a small symmetry-breaking field

$\chi$  also diverges near the critical point  
now no difference between spin up/down  
(barrier disappears)



Fluctuations for liq.-vapor equilibria  
(from Chandler). (a), (b),  $T \ll T_c$ ; (c)  $T < T_c$ ,  
with gravitational field; (d)  $T \approx T_c$