# DIMENSION REDUCTION FOR THIN FILMS WITH TRANSVERSALLY VARYING PRESTRAIN: THE OSCILLATORY AND THE NON-OSCILLATORY CASE 

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#### Abstract

We study the non-Euclidean (incompatible) elastic energy functionals in the description of prestressed thin films, at their singular limits ( $\Gamma$-limits) as $h \rightarrow 0$ in the film's thickness $h$. Firstly, we extend the prior results $[39,12,40]$ to arbitrary incompatibility metrics that depend on both the midplate and the transversal variables (the "non-oscillatory" case). Secondly, we analyze a more general class of incompatibilities, where the transversal dependence of the lower order terms is not necessarily linear (the "oscillatory" case), extending the results of [3, 47] to arbitrary metrics and higher order scalings. We exhibit connections between the two cases via projections of appropriate curvature forms on the polynomial tensor spaces. We also show the effective energy quantisation in terms of scalings as a power of $h$ and discuss the scaling regimes $h^{2}$ (Kirchhoff), $h^{4}$ (von Kármán) in the general case, as well as all possible (even powers) regimes for conformal metrics, thus paving the way to the subsequent complete analysis of the non-oscillatory setting in [34]. Thirdly, we prove the coercivity inequalities for the singular limits at $h^{2}$ - and $h^{4}$ - scaling orders, while disproving the full coercivity of the classical von Kármán energy functional at scaling $h^{4}$.


## 1. Introduction

The purpose of this paper is to further develop the analytical tools for understanding the mechanisms through which the local properties of a material lead to changes in its mechanical responses.

Motivated by the idea of imposing and controlling the prestrain (or "misfit") field in order to cause the plate to achieve a desired shape, our work is concerned with the analysis of thin elastic films exhibiting residual stress at free equilibria. Examples of this type of structures and their actuations include: plastically strained sheets, swelling or shrinking gels, growing tissues such as leaves, flowers or marine invertebrates, nanotubes, atomically thin graphene layers, etc. In the same vein, advancements in the construction of novel materials in thin film format require an analytical insight how the parameters affect the product and how to mimic the architectures found in nature.

In this paper, we will be concerned with the forward problem associated to the mentioned structures, based on the minimization of the elastic energy with incorporated inelastic effects.
1.1. The set-up of the problem. Let $\omega \subset \mathbb{R}^{2}$ be an open, bounded, connected set with Lipschitz boundary. We consider a family of thin hyperelastic sheets occupying the reference domains:

$$
\Omega^{h}=\omega \times\left(-\frac{h}{2}, \frac{h}{2}\right) \subset \mathbb{R}^{3}, \quad 0<h \ll 1 .
$$

A typical point in $\Omega^{h}$ is denoted by $x=\left(x_{1}, x_{2}, x_{3}\right)=\left(x^{\prime}, x_{3}\right)$. For $h=1$ we use the notation $\Omega=\Omega^{1}$ and view $\Omega$ as the referential rescaling of each $\Omega^{h}$ via: $\Omega^{h} \ni\left(x^{\prime}, x_{3}\right) \mapsto\left(x^{\prime}, x_{3} / h\right) \in \Omega$.

In this paper we study the limit behaviour, as $h \rightarrow 0$, of the energy functionals:

$$
\begin{equation*}
\mathcal{E}^{h}\left(u^{h}\right)=\frac{1}{h} \int_{\Omega^{h}} W\left(\nabla u^{h}(x) G^{h}(x)^{-1 / 2}\right) \mathrm{d} x \tag{1.1}
\end{equation*}
$$

defined on vector fields $u^{h} \in W^{1,2}\left(\Omega^{h}, \mathbb{R}^{3}\right)$, that are interpreted as deformations of $\Omega^{h}$. We view the limit in the vanishing thickness $h$ as the singular limit: indeed at $h=0$ the three-dimensional sheets $\Omega^{h}$ are reduced to the two-dimensional midplate $\omega$, and it is the goal of this paper to derive the energy of its (non-equidimensional) deformations $y: \omega \rightarrow \mathbb{R}^{3}$ relevant for the asymptotics of the minimizing sequences of the (equidimensional) deformations in (1.1). The sheets $\Omega^{h}$ are characterized by the smooth incompatibility (Riemann metric) tensors $G^{h} \in \mathcal{C}^{\infty}\left(\bar{\Omega}^{h}, \mathbb{R}_{\text {sym,pos }}^{3 \times 3}\right)$, satisfying the following structure assumption, referred to as "oscillatory":
(O)

$$
\begin{gathered}
\text { OsCILLATORY CASE : } \\
G^{h}(x)=\mathcal{G}^{h}\left(x^{\prime}, \frac{x_{3}}{h}\right) \quad \text { for all } x=\left(x^{\prime}, x_{3}\right) \in \Omega^{h}, \\
\mathcal{G}^{h}\left(x^{\prime}, t\right)=\overline{\mathcal{G}}\left(x^{\prime}\right)+h \mathcal{G}_{1}\left(x^{\prime}, t\right)+\frac{h^{2}}{2} \mathcal{G}_{2}\left(x^{\prime}, t\right)+o\left(h^{2}\right) \in \mathcal{C}^{\infty}\left(\bar{\Omega}, \mathbb{R}_{\mathrm{sym}, \mathrm{pos}}^{3 \times 3}\right),
\end{gathered}
$$

where $\overline{\mathcal{G}} \in \mathcal{C}^{\infty}\left(\bar{\omega}, \mathbb{R}_{\text {sym }, \text { pos }}^{3 \times 3}\right), \mathcal{G}_{1}, \mathcal{G}_{2} \in \mathcal{C}^{\infty}\left(\bar{\Omega}, \mathbb{R}_{\text {sym }}^{3 \times 3}\right)$ and $\int_{-1 / 2}^{1 / 2} \mathcal{G}_{1}\left(x^{\prime}, t\right) \mathrm{d} t=0$ for all $x^{\prime} \in \bar{\omega}$.
The requirement of $\overline{\mathcal{G}}$ being independent of the transversal variable $t \in(-1 / 2,1 / 2)$ is essential for the energy scaling order: $\inf \mathcal{E}^{h} \leq C h^{2}$. The zero mean requirement on $\mathcal{G}_{1}$ can be relaxed to requesting that $\int_{-1 / 2}^{1 / 2} \mathcal{G}_{1}\left(x^{\prime}, t\right)_{2 \times 2} \mathrm{~d} t$ be a linear strain with respect to the leading order midplate metric $\left(\overline{\mathcal{G}}_{1}\right)_{2 \times 2}$ (in case $(\overline{\mathcal{G}})_{2 \times 2}=I d_{2}$ the sufficient and necessary condition for this to happen is $\operatorname{curl}^{\top}$ curl $\int_{-1 / 2}^{1 / 2} \mathcal{G}_{1}\left(x^{\prime}, t\right)_{2 \times 2} \mathrm{~d} t=0$; this case has been studied in [3] where $\left.\overline{\mathcal{G}}=I d_{3}\right)$, and we also conjecture that it can be removed altogether, which will be the content of future work. In the present work, we assume the said condition in light of the special case (NO) below.

We refer to the family of films $\Omega^{h}$ prestrained by metrics in (O):

$$
\begin{equation*}
G^{h}(x)=\overline{\mathcal{G}}\left(x^{\prime}\right)+h \mathcal{G}_{1}\left(x^{\prime}, \frac{x_{3}}{h}\right)+\frac{h^{2}}{2} \mathcal{G}_{2}\left(x^{\prime}, \frac{x_{3}}{h}\right)+o\left(h^{2}\right) \quad \text { for all } \quad x=\left(x^{\prime}, x_{3}\right) \in \Omega^{h}, \tag{1.2}
\end{equation*}
$$

as "oscillatory", and note that this set-up includes a subcase of a single metric $G^{h}=G$, upon taking:

$$
\mathcal{G}_{1}\left(x^{\prime}, t\right)=t \overline{\mathcal{G}}_{1}\left(x^{\prime}\right), \quad \mathcal{G}_{2}\left(x^{\prime}, t\right)=t^{2} \overline{\mathcal{G}}_{2}\left(x^{\prime}\right) .
$$

We refer to this special case as "non-oscillatory"; formula (1.2) becomes then Taylor's expansion in:

$$
\text { (NO) }\left[\begin{array}{c}
\text { NON-OSCILLATORY CASE : } \\
G^{h}=G_{\mid \bar{\Omega}^{h}} \quad \text { for some } G \in \mathcal{C}^{\infty}\left(\bar{\Omega}, \mathbb{R}_{\mathrm{sym}, \mathrm{pos}}^{3 \times 3}\right), \\
G^{h}(x)=\overline{\mathcal{G}}\left(x^{\prime}\right)+x_{3} \partial_{3} G\left(x^{\prime}, 0\right)+\frac{x_{3}^{2}}{2} \partial_{33} G\left(x^{\prime}, 0\right)+o\left(x_{3}^{2}\right) \quad \text { for all } x=\left(x^{\prime}, x_{3}\right) \in \Omega^{h} .
\end{array}\right.
$$

Mechanically, the assumption (NO) describes thin sheets that have been cut out of a single specimen block $\Omega$, prestrained according to a fixed (though arbitrary) tensor $G$. As we shall see, the general
case (O) can be reduced to (NO) via the following effective metric:

$$
\begin{gather*}
\text { EfFECTIVE NON-OSCILLATORY CASE : } \\
\bar{G}^{h}(x)=\bar{G}(x)=\overline{\mathcal{G}}\left(x^{\prime}\right)+x_{3} \overline{\mathcal{G}}_{1}\left(x^{\prime}\right)+\frac{x_{3}^{2}}{2} \overline{\mathcal{G}}_{2}\left(x^{\prime}\right) \quad \text { for all } x=\left(x^{\prime}, x_{3}\right) \in \Omega^{h} \tag{EF}
\end{gather*}
$$

where: $\overline{\mathcal{G}}_{1}\left(x^{\prime}\right)_{2 \times 2}=12 \int_{-1 / 2}^{1 / 2} t \mathcal{G}_{1}\left(x^{\prime}, t\right)_{2 \times 2} \mathrm{~d} t, \quad \overline{\mathcal{G}}_{1}\left(x^{\prime}\right) e_{3}=-60 \int_{-1 / 2}^{1 / 2}\left(2 t^{3}-\frac{1}{2} t\right) \mathcal{G}_{1}\left(x^{\prime}, t\right) e_{3} \mathrm{~d} t$, and: $\quad \overline{\mathcal{G}}_{2}\left(x^{\prime}\right)_{2 \times 2}=30 \int_{-1 / 2}^{1 / 2}\left(6 t^{2}-\frac{1}{2}\right) \mathcal{G}_{2}\left(x^{\prime}, t\right)_{2 \times 2} \mathrm{~d} t$.

In (1.1), the homogeneous elastic energy density $W: \mathbb{R}^{3 \times 3} \rightarrow[0, \infty]$ is a Borel measurable function, assumed to satisfy the following properties:
(i) $W(R F)=W(F)$ for all $R \in S O(3)$ and $F \in \mathbb{R}^{3 \times 3}$,
(ii) $W(F)=0$ for all $F \in S O(3)$,
(iii) $W(F) \geq C \operatorname{dist}^{2}(F, S O(3))$ for all $F \in \mathbb{R}^{3 \times 3}$, with some uniform constant $C>0$,
(iv) there exists a neighbourhood $\mathcal{U}$ of $S O(3)$ such that $W$ is finite and $\mathcal{C}^{2}$ regular on $\mathcal{U}$.

We will be concerned with the regimes of curvatures of $G^{h}$ in (O) which yield the incompatibility rate, quantified by $\inf \mathcal{E}^{h}$, of order higher than $h^{2}$ in the plate's thickness $h$. With respect to the prior works in this context, the present paper proposes the following three new contributions.
1.2. New results of this work: Singular energies in the non-oscillatory case.
1.2.1. Kirchhoff scaling regime. We begin by deriving (in section 2), the $\Gamma$-limit of the rescaled energies $\frac{1}{h^{2}} \mathcal{E}^{h}$. In the setting of (NO), we obtain:

$$
\begin{aligned}
\mathcal{I}_{2}(y) & =\frac{1}{2} \| \text { Tensor }_{2}\left\|_{\mathcal{Q}_{2}}^{2}=\frac{1}{2}\right\| x_{3}\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2} x_{3} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2} \|_{\mathcal{Q}_{2}}^{2} \\
& =\frac{1}{24}\left\|\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}\right\|_{\mathcal{Q}_{2}}^{2} .
\end{aligned}
$$

We now explain the notation above. Firstly, $\|\cdot\|_{\mathcal{Q}_{2}}$ is a weighted $L^{2}$ norm in (2.8) on the space $\mathbb{E}$ of $\mathbb{R}_{\text {sym }}^{2 \times 2}$-valued tensor fields on $\Omega$. The weights in (2.6) are determined by the elastic energy $W$ together with the leading order metric coefficient $\overline{\mathcal{G}}$. The functional $\mathcal{I}_{2}$ is defined on the set of isometric immersions $\mathcal{Y}_{\overline{\mathcal{G}}_{2 \times 2}}=\left\{y \in W^{2,2}\left(\omega, \mathbb{R}^{3}\right) ;(\nabla y)^{\top} \nabla y=\overline{\mathcal{G}}_{2 \times 2}\right\}$; each such immersion generates the corresponding Cosserat vector $\vec{b}$, uniquely given by requesting: $\left[\partial_{1} y, \partial_{2} y, \vec{b}\right] \in S O(3) \overline{\mathcal{G}}^{1 / 2}$ on $\omega$. The family of energies obtained in this manner is parametrised by all matrix fields $S \in \mathcal{C}^{\infty}\left(\bar{\omega}, \mathbb{R}_{\text {sym }}^{2 \times 2}\right)$ and $T \in \mathcal{C}^{\infty}\left(\bar{\omega}, \mathbb{R}_{\text {sym }, \text { pos }}^{3 \times 3}\right)$, namely: $\mathcal{I}_{2}^{T, S}(y)=\frac{1}{24}\left\|\left((\nabla y)^{\top} \nabla \vec{b}(y)\right)_{\text {sym }}-S\right\|_{\mathcal{Q}_{2}}^{2}$ defined on the set $\mathcal{Y}_{T_{2 \times 2}}$, where one interprets $T$ as the leading order prestrain $\overline{\mathcal{G}}$ and $S$ as its first order correction $\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}$.

The energy $\mathcal{I}_{2}$ measures the bending quantity Tensor $_{2}$ which is linear in $x_{3}$, resulting in its reduction to the single nonlinear bending term, that equals the difference of the curvature form $\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}$ from the preferred curvature $\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}$. The same energy has been derived in [39, 12] under the assumption that $G$ is independent of $x_{3}$ and in [31] for a general manifold ( $M^{n}, g$ ) with any codimension submanifold $\left(N^{k}, g_{\mid N}\right)$ replacing the midplate $\omega \times\{0\}$. Since our derivation of $\mathcal{I}_{2}$ is a particular case of the result in case ( O ), we still state it here for completeness.

In section 3 we identify the necessary and sufficient conditions for $\min \mathcal{I}_{2}=0$ (when $\omega$ is simply connected), in terms of the vanishing of the Riemann curvatures $R_{1212}, R_{1213}, R_{1223}$ of $G$ at $x_{3}=0$. In this case, it follows that $\inf \mathcal{E}^{h} \leq C h^{4}$. For the discussed case (NO), the recent work [42] generalized the same statements for arbitrary dimension and codimension.
1.2.2. Von Kármán scaling regime. In section 6 we derive the $\Gamma$-limit of $\frac{1}{h^{4}} \mathcal{E}^{h}$, which is given by:

$$
\mathcal{I}_{4}(V, \mathbb{S})=\frac{1}{2} \| \text { Tensor }_{4} \|_{\mathcal{Q}_{2}}^{2}
$$

defined on the spaces of: finite strains $\mathscr{S}_{y_{0}}=\left\{\mathbb{S}=\lim _{n \rightarrow \infty, L^{2}}\left(\left(\nabla y_{0}\right)^{\top} \nabla w_{n}\right)_{\text {sym }} ; w_{n} \in W^{1,2}\left(\omega, \mathbb{R}^{3}\right)\right\}$ and first order infinitesimal isometries $\mathscr{V}_{y_{0}}=\left\{V \in W^{2,2}\left(\omega, \mathbb{R}^{3}\right) ;\left(\left(\nabla y_{0}\right)^{\top} \nabla V\right)_{\text {sym }}=0\right\}$ on the deformed midplate $y_{0}(\omega) \subset \mathbb{R}^{3}$. Here, $y_{0}$ is the unique smooth isometric immersion of $\overline{\mathcal{G}}_{2 \times 2}$ for which $\mathcal{I}_{2}\left(y_{0}\right)=0$; recall that it generates the corresponding Cosserat's vector $\vec{b}_{0}$.

The expression in $\mathrm{Tensor}_{4}$ is quite complicated but it has the structure of a quadratic polynomial in $x_{3}$. A key tool for identifying this expression, also in the general case ( O ), involves the subspaces $\left\{\mathbb{E}_{n} \subset \mathbb{E}\right\}_{n \geq 1}$ in (2.9), consisting of the tensorial polynomials in $x_{3}$ of order $n$. The bases of $\left\{\mathbb{E}_{n}\right\}$ are then naturally given in terms of the Legendre polynomials $\left\{p_{n}\right\}_{n \geq 0}$ on $\left(-\frac{1}{2}, \frac{1}{2}\right)$. Since Tensor ${ }_{4} \in \mathbb{E}_{2}$, we write the decomposition:

$$
\text { Tensor }_{4}=p_{0}\left(x_{3}\right) \text { Stretching }_{4}+p_{1}\left(x_{3}\right) \text { Bending }_{4}+p_{2}\left(x_{3}\right) \text { Curvature }{ }_{4},
$$

which, as shown in section 7 , results in:

$$
\begin{aligned}
\mathcal{I}_{4}(V, \mathbb{S})= & \frac{1}{2}\left(\| \text { Stretching }_{4}\left\|_{\mathcal{Q}_{2}}^{2}+\right\| \text { Bending }_{4}\left\|_{\mathcal{Q}_{2}}^{2}+\right\| \text { Curvature }_{4} \|_{\mathcal{Q}_{2}}^{2}\right) \\
= & \frac{1}{2}\left\|\mathbb{S}+\frac{1}{2}(\nabla V)^{\top} \nabla V+\frac{1}{24}\left(\nabla \vec{b}_{0}\right)^{\top} \nabla \vec{b}_{0}-\frac{1}{48} \partial_{33} G\left(x^{\prime}, 0\right)_{2 \times 2}\right\|_{\mathcal{Q}_{2}}^{2} \\
& +\frac{1}{24}\left\|\left[\left\langle\nabla_{i} \nabla_{j} V, \vec{b}_{0}\right\rangle\right]_{i, j=1,2}\right\|_{\mathcal{Q}_{2}}^{2}+\frac{1}{1440}\left\|\left[R_{i 3 j 3}\left(x^{\prime}, 0\right)\right]_{i, j=1,2}\right\|_{\mathcal{Q}_{2}}^{2} .
\end{aligned}
$$

Above, $\nabla_{i}$ denotes the covariant differentiation with respect to the metric $\overline{\mathcal{G}}$ and $R_{i 3 j 3}$ are the potentially non-zero curvatures of $G$ on $\omega$ at $x_{3}=0$.

The family of energies obtained in this manner is parametrised by all quadruples: vector fields $y_{0}, \vec{b}_{0} \in \mathcal{C}^{\infty}\left(\bar{\omega}, \mathbb{R}^{3}\right)$ satisfying $\operatorname{det}\left[\partial_{1} y_{0}, \partial_{2} y_{0}, \vec{b}_{0}\right]>0$, matrix fields $T \in \mathcal{C}^{\infty}\left(\bar{\omega}, \mathbb{R}_{\text {sym }}^{2 \times 2}\right)$, and numbers $r \in \mathbb{R}$ in the range (the left parentheses in the last interval below may be open or closed):

$$
\begin{equation*}
r \in \frac{2}{5}\left\{\left\|T-\left(\left(\nabla y_{0}\right)^{\top} \nabla \vec{d}_{0}\right)_{\mathrm{sym}}\right\|_{\mathcal{Q}_{2}}^{2} ; \vec{d}_{0} \in \mathcal{C}^{\infty}\left(\bar{\omega}, \mathbb{R}_{\mathrm{sym}}^{2 \times 2}\right)\right\}=\left[\frac{2}{5} \operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(T, \mathscr{S}_{y_{0}}\right),+\infty\right) . \tag{1.3}
\end{equation*}
$$

The functionals are then: $\mathcal{I}_{4}^{y_{0}, \vec{b}_{0}, T, r}(V, \mathbb{S})=\frac{1}{2}\left\|\mathbb{S}+\frac{1}{2}(\nabla V)^{\top} \nabla V-T\right\|_{\mathcal{Q}_{2}}^{2}+\frac{1}{24}\left\|\left[\left\langle\nabla_{i} \nabla_{j} V, \vec{b}_{0}\right\rangle\right]_{i j}\right\|_{\mathcal{Q}_{2}}^{2}+r$, defined on the linear space $\mathcal{V}_{y_{0}} \times \mathscr{S}_{y_{0}}$. Particular cases where the range of $r$ may be identified are:
(i) $y_{0}=i d_{2}$. Then $\mathscr{S}_{y_{0}}=\left\{\mathbb{S} \in L^{2}\left(\omega, \mathbb{R}_{\text {sym }}^{2 \times 2}\right)\right.$; $\left.\operatorname{curl}^{\top} \operatorname{curl} \mathbb{S}=0\right\}$ and the range of $r$ is defined by the appropriate norm of: $\operatorname{curl}^{\top} \operatorname{curl} T$.
(ii) Gauss curvature $\kappa\left(\left(\nabla y_{0}\right)^{\top} \nabla y_{0}\right)>0$ in $\bar{\omega}$. Then in [37] it is shown that $\mathscr{S}_{y_{0}}=L^{2}\left(\omega, \mathbb{R}_{\text {sym }}^{2 \times 2}\right)$. The range of possible $r$ (for any $T$ ) is then: $[0,+\infty)$.
When $y_{0}=i d_{2}$ (which occurs automatically when $\overline{\mathcal{G}}=I d_{3}$ ), then $\vec{b}_{0}=e_{3}$ and the first two terms in $\mathcal{I}_{4}$ reduce to the stretching and the linear bending contents of the classical von Kármán energy. The third term is purely metric-related and measures the non-immersability of $G$ relative to the present quartic scaling. These findings generalize the results of [12] valid for $x_{3}$-independent $G$ in (NO). We also point out that, following the same general principle in the $h^{2}$ - scaling regime, one may readily decompose:

$$
\text { Tensor }_{2}=p_{0}\left(x_{3}\right) \text { Stretching }_{2}+p_{1}\left(x_{3}\right) \text { Bending }_{2}
$$

since Tensor $_{2}$ is already a multiple of $x_{3}$, then Stretching $_{2}=0$ in the ultimate form of $\mathcal{I}_{2}$.

It is not hard to deduce (see section 8) that the necessary and sufficient conditions for having $\min \mathcal{I}_{4}=0$ are precisely that $R_{i j k l} \equiv 0$ on $\omega \times\{0\}$, for all $i, j, k, l=1 \ldots 3$. In section 9 we then analyze the conformal non-oscillatory metric $G=e^{2 \phi\left(x_{3}\right)} I d_{3}$ and show that different orders of vanishing of $\phi$ at $x_{3}=0$ correspond to different even orders of scaling of $\mathcal{E}^{h}$ as $h \rightarrow 0$ :

$$
\phi^{(k)}(0)=0 \quad \text { for } k=1 \ldots n-1 \quad \text { and } \quad \phi^{(n)}(0) \neq 0 \quad \Leftrightarrow \quad c h^{2 n} \leq \inf \mathcal{E}^{h} \leq C h^{2 n}
$$

with the lower bound: $\inf \mathcal{E}^{h} \geq c_{n} h^{n}\left\|\left[\partial_{3}^{(n-2)} R_{i 3 j 3}\left(x^{\prime}, 0\right)\right]_{i, j=1,2}\right\|_{\mathcal{Q}_{2}}^{2}$. These findings are consistent with and pave the way for the follow-up paper [34], which completes the scaling analysis of $\mathcal{E}^{h}$ in the non-oscillatory case, including the derivation of $\Gamma$-limits of $h^{-2 n} \mathcal{E}^{h}$ for all $n \geq 1$, and proving the energy quantisation in the sense that the even powers $2 n$ of $h$ are indeed the only possible ones (all of them are also attained).
1.3. New results of this work: Singular energies in the oscillatory case. We show that the analysis in the general case ( O ) may follow a similar procedure, where we first project the limiting quantity Tensor ${ }^{O}$ on an appropriate polynomial space and then decompose the projection along the respective Legendre basis. For the $\Gamma$-limit of $\frac{1}{h^{2}} \mathcal{E}^{h}$ in section 2, we show that:

$$
\begin{aligned}
& \text { Tensor }_{2}^{O}=x_{3}\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2}\left(\mathcal{G}_{1}\right)_{2 \times 2}=p_{0}\left(x_{3}\right) \text { Stretching }_{2}^{O}+p_{1}\left(x_{3}\right) \text { Bending }_{2}^{O}+\text { Excess }_{2}, \\
& \text { with } \text { Excess }_{2}=\text { Tensor }_{2}^{O}-\mathbb{P}_{1}\left(\text { Tensor }_{2}^{O}\right),
\end{aligned}
$$

where $\mathbb{P}_{1}$ denotes the orthogonal projection on $\mathbb{E}_{1}$. Consequently:

$$
\begin{aligned}
\mathcal{I}_{2}^{O}(y) & =\frac{1}{2}\left(\| \text { Stretching }_{2}^{O}\left\|_{\mathcal{Q}_{2}}^{2}+\right\| \text { Bending }_{2}^{O}\left\|_{\mathcal{Q}_{2}}^{2}+\right\| \text { Excess }_{2} \|_{\mathcal{Q}_{2}}^{2}\right) \\
& =\frac{1}{24}\left\|\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2}\left(\overline{\mathcal{G}}_{1}\right)_{2 \times 2}\right\|_{\mathcal{Q}_{2}}^{2}+\frac{1}{8} \operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(\left(\mathcal{G}_{1}\right)_{2 \times 2}, \mathbb{E}_{1}\right),
\end{aligned}
$$

where again Stretching $g_{2}^{O}=0$ in view of the assumed $\int_{-1 / 2}^{1 / 2} \mathcal{G}_{1} \mathrm{~d} x_{3}=0$. For the same reason:

$$
\text { Excess }_{2}=-\frac{1}{2}\left(\left(\mathcal{G}_{1}\right)_{2 \times 2}-\mathbb{P}_{1}\left(\left(\mathcal{G}_{1}\right)_{2 \times 2}\right)\right)=-\frac{1}{2}\left(\left(\mathcal{G}_{1}\right)_{2 \times 2}-12 \int_{-1 / 2}^{1 / 2} x_{3}\left(\mathcal{G}_{1}\right)_{2 \times 2} \mathrm{~d} x_{3}\right)
$$

and also: $\mathbb{P}_{1}\left(\left(\mathcal{G}_{1}\right)_{2 \times 2}\right)=x_{3}\left(\overline{\mathcal{G}}_{1}\right)_{2 \times 2}$ with $\left(\overline{\mathcal{G}}_{1}\right)_{2 \times 2}$ defined in (EF). The limiting oscillatory energy $\mathcal{I}_{2}^{O}$ consists thus of the bending term that coincides with $\mathcal{I}_{2}$ for the effective metric $\bar{G}$, plus the purely metric-related excess term. A special case of $\mathcal{I}_{2}^{O}$ when $\overline{\mathcal{G}}=I d_{3}$ and without analyzing the excess term, has been derived in [3], following the case with $G^{h}=I d_{3}+h \mathcal{G}_{1}\left(\frac{x_{3}}{h}\right)$ considered in [47]. An excess term has also been present in the work [28] on rods with misfit; we do not attempt to compare our results with studies of dimension reduction for rods; the literature there is abundant.

It is easy to observe that: $\min \mathcal{I}_{2}^{O}=0$ if and only if $\left(\mathcal{G}_{1}\right)_{2 \times 2}=x_{3}\left(\overline{\mathcal{G}}_{1}\right)_{2 \times 2}$ on $\omega \times\{0\}$. We show in section 5 that this automatically implies: $\inf \mathcal{E}^{h} \leq C h^{4}$. The $\Gamma$-limit of $\frac{1}{h^{4}} \mathcal{E}^{h}$ is further derived in sections 6 and 7 , by considering the decomposition:

$$
\begin{aligned}
& \text { Tensor }_{4}^{O}=p_{0}\left(x_{3}\right) \text { Stretching }_{4}^{O}+p_{1}\left(x_{3}\right) \text { Bending }_{4}^{O}+p_{2}\left(x_{3}\right) \text { Curvature }{ }_{4}^{O}+\text { Excess }_{4}, \\
& \text { with } \text { Excess }_{4}=\text { Tensor }_{4}^{O}-\mathbb{P}_{2}\left(\text { Tensor }_{4}^{O}\right) .
\end{aligned}
$$

It follows that:

$$
\begin{aligned}
\mathcal{I}_{4}^{O}(V, \mathbb{S})= & \frac{1}{2}\left(\| \text { Stretching }_{4}^{O}\left\|_{\mathcal{Q}_{2}}^{2}+\right\| \text { Bending }_{4}^{O}\left\|_{\mathcal{Q}_{2}}^{2}+\right\| \text { Curvature }_{4}^{O}\left\|_{\mathcal{Q}_{2}}^{2}+\right\| \text { Excess }_{4} \|_{\mathcal{Q}_{2}}^{2}\right) \\
= & \frac{1}{2}\left\|\mathbb{S}+\frac{1}{2}(\nabla V)^{\top} \nabla V+B_{0}\right\|_{\mathcal{Q}_{2}}^{2}+\frac{1}{24}\left\|\left[\left\langle\nabla_{i} \nabla_{j} V, \vec{b}_{0}\right\rangle\right]_{i, j=1,2}+B_{1}\right\|_{\mathcal{Q}_{2}}^{2} \\
& +\frac{1}{1440}\left\|\left[R_{i 3 j 3}\left(x^{\prime}, 0\right)\right]_{i, j=1,2}\right\|_{\mathcal{Q}_{2}}^{2} \\
& +\frac{1}{2} \operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(\frac{1}{4}\left(\mathcal{G}_{2}\right)_{2 \times 2}-\int_{0}^{x_{3}}\left[\nabla_{i}\left(\left(\mathcal{G}_{1} e_{3}\right)-\frac{1}{2}\left(\mathcal{G}_{1}\right)_{33} e_{3}\right)\right]_{i, j=1,2, \text { sym }} \mathrm{d} t, \mathbb{E}_{2}\right),
\end{aligned}
$$

where $R_{1313}, R_{1323}, R_{2323}$ are the respective Riemann curvatures of the effective metric $\bar{G}$ in (EF) at $x_{3}=0$. The corrections $B_{0}$ and $B_{1}$ coincide with the same expressions written for $\bar{G}$ under two extra constraints (see Theorem 7.5), that can be seen as the $h^{4}$-order counterparts of the $h^{2}$-order condition $\int_{-1 / 2}^{1 / 2} \mathcal{G}_{1} \mathrm{~d} x_{3}=0$ assumed throughout. In case these conditions are valid, the functional $\mathcal{I}_{4}^{O}$ is the sum of the effective stretching, bending and curvature in $\mathcal{I}_{4}$ for $\bar{G}$, plus the additional purely metric-related excess term.
1.4. New results of this work: coercivity of $\mathcal{I}_{2}$ and $\mathcal{I}_{4}$. We additionally analyze the derived limiting functionals by identifying their kernels, when nonempty. In section 4 we show that the kernel of $\mathcal{I}_{2}$ consists of the rigid motions of a single smooth deformation $y_{0}$ that solves:

$$
\left(\nabla y_{0}\right)^{\top} \nabla y_{0}=\overline{\mathcal{G}}_{2 \times 2}, \quad\left(\left(\nabla y_{0}\right)^{\top} \nabla \vec{b}_{0}\right)_{\text {sym }}=\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}
$$

Further, $\mathcal{I}_{2}(y)$ bounds from above the squared distance of an arbitrary $W^{2,2}$ isometric immersion $y$ of the midplate metric $\overline{\mathcal{G}}_{2 \times 2}$, from the indicated kernel of $\mathcal{I}_{2}$.

In section 8 we consider the case of $\mathcal{I}_{4}$. We first identify (see Theorem 8.2) the zero-energy displacement-strain couples $(V, \mathbb{S})$. In particular, we show that the minimizing displacements are exactly the linearised rigid motions of the referential $y_{0}$. We then prove that the bending term in $\mathcal{I}_{4}$, which is solely a function of $V$, bounds from above the squared distance of an arbitrary $W^{2,2}$ displacement obeying $\left(\left(\nabla y_{0}\right)^{\top} \nabla V\right)_{\text {sym }}=0$, from the indicated minimizing set in $V$. On the other hand, the full coercivity result involving minimization in both $V$ and $\mathbb{S}$ is false. In Remark 8.4 we exhibit an example in the setting of the classical von Kármán functional, where $\mathcal{I}_{4}\left(V_{n}, \mathbb{S}_{n}\right) \rightarrow 0$ as $n \rightarrow \infty$, but the distance of $\left(V_{n}, \mathbb{S}_{n}\right)$ from the kernel of $\mathcal{I}_{4}$ remains uniformly bounded away from 0 . We note that this lack of coercivity is not prevented by the fact that the kernel is finite dimensional.
1.5. Other related works. Recently, there has been a sustained interest in studying shape formation driven by internal prestrain, through the experimental, modelling via formal methods, numerics, and analytical arguments $[27,23,16,30]$. General results have been derived in the abstract setting of Riemannian manifolds: in $[31,30] \Gamma$-convergence statements were proved for any dimension ambient manifold and codimension midplate, in the scaling regimes $\mathcal{O}\left(h^{2}\right)$ and $\mathcal{O}(1)$, respectively. In a work parallel to ours [42], the authors analyze scaling orders $o\left(h^{2}\right), \mathcal{O}\left(h^{4}\right)$ and $o\left(h^{4}\right)$, extending condition (3.6), Lemma 5.1 in (NO) case, and condition (8.2) to arbitrary manifolds. Although they do not identify the $\Gamma$-limits of the rescaled energies $\mathcal{E}^{h}$, they are able to provide the revealing lower bounds for $\inf \mathcal{E}^{h}$ in terms of the appropriate curvatures.

Higher energy scalings $\inf \mathcal{E}^{h} \sim h^{\beta}$ than the ones analyzed in the present paper may result from the interaction of the metric with boundary conditions or external forces, leading to the "wrinklinglike" effects. Indeed, our setting pertains to the "no wrinkling" regime where $\beta \geq 2$ and the prestrain metric admits a $W^{2,2}$ isometric immersion. While the systematic description of the singular limits at scalings $\beta<2$ is not yet available, the following studies are examples of the variety of emerging patterns. In [24, 9, 10], energies leading to the buckling- or compression- driven blistering in a thin film breaking away from its substrate and under clamped boundary conditions, are discussed ( $\beta=1$ ). Paper [6] displays dependence of the energy minimization on boundary conditions and classes of admissible deformations, while [7] discusses coarsening of folds in hanging drapes, where the energy identifies the number of generations of coarsening. In [49], wrinkling patterns are obtained, reproducing the experimental observations when a thin shell is placed on a liquid bath ( $\beta=1$ ), while [29] analyses wrinkling in the center of a stretched and twisted ribbon $(\beta=4 / 3)$. In [11, 50], energy levels of the origami patterns in paper crumpling are studied $(\beta=5 / 3)$. See also [43, 44, 45] for an analysis of the conical singularities $\left(\mathcal{E}^{h} \sim h^{2} \log (1 / h)\right)$. We remark that the mentioned papers do not address the dimension reduction, but rather analyze the chosen actual configuration of the prestrained sheet. Closely related is also the literature on shape selection in non-Euclidean plates, exhibiting hierarchical buckling patterns in zero-strain plates ( $\beta=2$ ), where the complex morphology is due to the non-smooth energy minimization [19, 20, 21].

Various geometrically nonlinear thin plate theories have been used to analyze the self-similar structures with metric asymptotically flat at infinity [5], a disk with edge-localized growth [16], the shape of a long leaf [41], or torn plastic sheets [48]. In [13, 14] a variant of the Föppl-von Kármán equilibrium equations has been formally derived from finite incompressible elasticity, via the multiplicative decomposition of deformation gradient [46] similar to ours. See also models related to wrinkling of paper in areas with high ink or paint coverage, grass blades, sympatelous (meaning "with fused petals") flowers and studies of movement of micro-organisms that share certain characteristics of animals and plants, called the euglenids $[14,8,4]$. The forward and inverse problems in the self-folding of thin sheets of patterned hydrogel bilayers are discussed in [2].

On the frontiers of experimental modeling of shape formation, we mention the halftone gel lithography method for polymeric materials that can swell by imbibing fluids [26, 25, 51, 15]. By blocking the ability of portions of plate to swell or causing them to swell inhomogeneously, it is possible to have the plate assume a variety of deformed shapes. Even more sophisticated techniques of biomimetic $4 d$ printing allow for engineering of the 3 d shape-morphing systems that mimic nastic plant motions where organs such as tendrils, leaves and flowers respond to the environmental stimuli [22]. Optimal control in such systems has been studied in [23], see also [1].

In $[35,36,38]$, derivations similar to the results of the present paper were carried out under a different assumption on the asymptotic behavior of the prestrain, which also implied energy scaling $h^{\beta}$ in non-even regimes of $\beta>2$. In [35] it was shown that the resulting Euler-Lagrange equations of the residual energy are identical to those describing the effects of growth in elastic plates [41]. In [38], a model with a Monge-Ampère constraint was derived and analysed from various aspects.

We finally mention the paper [34], completed after the submission of the present article, which resolves the scaling analysis of $\mathcal{E}^{h}$ together with the derivation of $\Gamma$-limits of $h^{-2 n} \mathcal{E}^{h}$, for all $n \geq 1$. There, we identify equivalent conditions for the validity of the scalings $h^{2 n}$ in terms of vanishing of the Riemann curvatures, for an arbitrary non-oscillatory metric $G$, up to appropriate orders and in terms of the matched isometry expansions. We also establish the asymptotic behaviour of the
minimizing immersions of $G$ as $h \rightarrow 0$ and prove the energy quantisation, in the sense that the even powers $2 n$ of $h$ are indeed the only possible ones (all of them are also attained).
1.6. Notation. Given a matrix $F \in \mathbb{R}^{n \times n}$, we denote its transpose by $F^{\top}$ and its symmetric part by $F_{\text {sym }}=\frac{1}{2}\left(F+F^{\top}\right)$. The space of symmetric $n \times n$ matrices is denoted by $\mathbb{R}_{\text {sym }}^{n \times n}$, whereas $\mathbb{R}_{\text {sym,pos }}^{n \times n}$ stands for the space of symmetric, positive definite $n \times n$ matrices. By $S O(n)=\left\{R \in \mathbb{R}^{n \times n} ; R^{\top}=\right.$ $R^{-1}$ and $\left.\operatorname{det} R=1\right\}$ we mean the group of special rotations; its tangent space at $I d_{n}$ consists of skew-symmetric matrices: $T_{I d_{n}} S O(n)=s o(n)=\left\{F \in \mathbb{R}^{n \times n} ; F_{\text {sym }}=0\right\}$. We use the matrix norm $|F|=\left(\operatorname{trace}\left(F^{\top} F\right)\right)^{1 / 2}$, which is induced by the inner product $\left\langle F_{1}: F_{2}\right\rangle=\operatorname{trace}\left(F_{1}^{\top} F_{2}\right)$. The $2 \times 2$ principal minor of $F \in \mathbb{R}^{3 \times 3}$ is denoted by $F_{2 \times 2}$. Conversely, for a given $F_{2 \times 2} \in \mathbb{R}^{2 \times 2}$, the $3 \times 3$ matrix with principal minor equal $F_{2 \times 2}$ and all other entries equal to 0 , is denoted by $F_{2 \times 2}^{*}$. Unless specified otherwise, all limits are taken as the thickness parameter $h$ vanishes: $h \rightarrow 0$. By $C$ we denote any universal positive constant, independent of $h$. We use the Einstein summation convention over repeated lower and upper indices running from 1 to 3 .
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## 2. Compactness and $\Gamma$-Limit under $C h^{2}$ energy bound

Define the matrix fields $\bar{A} \in \mathcal{C}^{\infty}\left(\bar{\omega}, \mathbb{R}_{\text {sym }, \text { pos }}^{3 \times 3}\right)$ and $A^{h}, A_{1}, A_{2} \in \mathcal{C}^{\infty}\left(\bar{\Omega}, \mathbb{R}_{\text {sym }}^{3 \times 3}\right)$ so that, uniformly for all $\left(x^{\prime}, x_{3}\right) \in \Omega^{h}$ there holds:

$$
A^{h}\left(x^{\prime}, x_{3}\right)=\mathcal{G}^{h}\left(x^{\prime}, x_{3}\right)^{1 / 2}=\bar{A}\left(x^{\prime}\right)+h A_{1}\left(x^{\prime}, \frac{x_{3}}{h}\right)+\frac{h^{2}}{2} A_{2}\left(x^{\prime}, \frac{x_{3}}{h}\right)+o\left(h^{2}\right)
$$

Equivalently, $\bar{A}, A_{1}, A_{2}$ solve the following system of equations:

$$
\begin{equation*}
\bar{A}^{2}=\overline{\mathcal{G}}, \quad 2\left(\bar{A} A_{1}\right)_{\text {sym }}=\mathcal{G}_{1}, \quad 2 A_{1}^{2}+2\left(\bar{A} A_{2}\right)_{\text {sym }}=\mathcal{G}_{2} \quad \text { in } \bar{\Omega} . \tag{2.1}
\end{equation*}
$$

Under the assumption (O), condition (iii) on $W$ easily implies:

$$
\begin{aligned}
\frac{1}{h} \int_{\Omega^{h}} \operatorname{dist}^{2}\left(\nabla u^{h}(x) \bar{A}\left(x^{\prime}\right)^{-1}, S O(3)\right) \mathrm{d} x & \leq \frac{C}{h} \int_{\Omega^{h}} \operatorname{dist}^{2}\left(\nabla u^{h}(x) A^{h}(x)^{-1}, S O(3)\right)+h^{2} \mathrm{~d} x \\
& \leq C\left(\mathcal{E}^{h}\left(u^{h}\right)+h^{2}\right) .
\end{aligned}
$$

Consequently, the results of [12] automatically give the following compactness properties of any sequence of deformations with the quadratic energy scaling:

Theorem 2.1. Assume (O). Let $\left\{u^{h} \in W^{1,2}\left(\Omega^{h}, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0}$ be a sequence of deformations satisfying:

$$
\begin{equation*}
\mathcal{E}^{h}\left(u^{h}\right) \leq C h^{2} . \tag{2.2}
\end{equation*}
$$

Then the following properties hold for the rescalings $y^{h} \in W^{1,2}\left(\Omega, \mathbb{R}^{3}\right)$ given by:

$$
y^{h}\left(x^{\prime}, x_{3}\right)=u^{h}\left(x^{\prime}, h x_{3}\right)-f_{\Omega^{h}} u^{h} \mathrm{~d} x .
$$

(i) There exist $y \in W^{2,2}\left(\omega, \mathbb{R}^{3}\right)$ and $\vec{b} \in W^{1,2} \cap L^{\infty}\left(\omega, \mathbb{R}^{3}\right)$ such that, up to a subsequence:

$$
y^{h} \rightarrow y \quad \text { strongly in } W^{1,2}\left(\Omega, \mathbb{R}^{3}\right) \quad \text { and } \quad \frac{1}{h} \partial_{3} y^{h} \rightarrow \vec{b} \quad \text { strongly in } L^{2}\left(\Omega, \mathbb{R}^{3}\right), \quad \text { as } h \rightarrow 0 .
$$

(ii) The limit deformation $y$ realizes the reduced midplate metric on $\omega$ :

$$
\begin{equation*}
(\nabla y)^{\top} \nabla y=\overline{\mathcal{G}}_{2 \times 2} . \tag{2.3}
\end{equation*}
$$

In particular $\partial_{1} y, \partial_{2} y \in L^{\infty}\left(\omega, \mathbb{R}^{3}\right)$ and the unit normal $\vec{\nu}=\frac{\partial_{1} y \times \partial_{2} y}{\left|\partial_{1} y \times \partial_{2} y\right|}$ to the surface $y(\omega)$ satisfies: $\vec{\nu} \in W^{1,2} \cap L^{\infty}\left(\omega, \mathbb{R}^{3}\right)$. The limit displacement $\vec{b}$ is the Cosserat field defined via:

$$
\vec{b}=(\nabla y)\left(\overline{\mathcal{G}}_{2 \times 2}\right)^{-1}\left[\begin{array}{l}
\overline{\mathcal{G}}_{13}  \tag{2.4}\\
\overline{\mathcal{G}}_{23}
\end{array}\right]+\frac{\sqrt{\operatorname{det} \overline{\mathcal{G}}}}{\sqrt{\operatorname{det} \overline{\mathcal{G}}_{2 \times 2}}} \vec{\nu} .
$$

Recall that the results in [12] also give:

$$
\begin{equation*}
\liminf _{h \rightarrow 0} \frac{1}{h^{2}} \frac{1}{h} \int_{\Omega^{h}} W\left(\nabla u^{h} \overline{\mathcal{G}}^{-1 / 2}\right) \geq \frac{1}{24} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime}, \nabla y\left(x^{\prime}\right)^{\top} \nabla \vec{b}\left(x^{\prime}\right)\right) \mathrm{d} x^{\prime} \tag{2.5}
\end{equation*}
$$

with the curvature integrand $(\nabla y)^{\top} \nabla \vec{b}$ quantified by the quadratic forms:

$$
\begin{align*}
& \mathcal{Q}_{2}\left(x^{\prime}, F_{2 \times 2}\right)=\min \left\{\mathcal{Q}_{3}\left(\bar{A}\left(x^{\prime}\right)^{-1} \tilde{F} \bar{A}\left(x^{\prime}\right)^{-1}\right) ; \tilde{F} \in \mathbb{R}^{3 \times 3} \text { with } \tilde{F}_{2 \times 2}=F_{2 \times 2}\right\},  \tag{2.6}\\
& \mathcal{Q}_{3}(F)=D^{2} W\left(I d_{3}\right)(F, F)
\end{align*}
$$

The form $\mathcal{Q}_{3}$ is defined for all $F \in \mathbb{R}^{3 \times 3}$, while $\mathcal{Q}_{2}\left(x^{\prime}, \cdot\right)$ are defined on $F_{2 \times 2} \in \mathbb{R}^{2 \times 2}$. Both forms $\mathcal{Q}_{3}$ and all $\mathcal{Q}_{2}$ are nonnegative definite and depend only on the symmetric parts of their arguments, in view of the assumptions on the elastic energy density $W$. Clearly, the minimization problem in (2.6) has a unique solution among symmetric matrices $\tilde{F}$ which for each $x^{\prime} \in \omega$ is described by the linear function $F_{2 \times 2} \mapsto c\left(x^{\prime}, F_{2 \times 2}\right) \in \mathbb{R}^{3}$ in:

$$
\begin{equation*}
\mathcal{Q}_{2}\left(x^{\prime}, F_{2 \times 2}\right)=\min \left\{\mathcal{Q}_{3}\left(\bar{A}\left(x^{\prime}\right)^{-1}\left(F_{2 \times 2}^{*}+c \otimes e_{3}\right) \bar{A}\left(x^{\prime}\right)^{-1}\right) ; c \in \mathbb{R}^{3}\right\} . \tag{2.7}
\end{equation*}
$$

The energy in the right hand side of (2.5) is a Kirchhoff-like fully nonlinear bending, which in case of $\bar{A} e_{3}=e_{3}$ reduces to the classical bending content relative to the second fundamental form $(\nabla y)^{\top} \nabla \vec{b}=(\nabla y)^{\top} \nabla \vec{\nu}$ on the deformed surface $y(\omega)$.

In the present setting, we start with an observation about projections on polynomial subspaces of $L^{2}$. Consider the following Hilbert space, with its norm:

$$
\begin{equation*}
\mathbb{E} \doteq\left(L^{2}\left(\Omega, \mathbb{R}_{\mathrm{sym}}^{2 \times 2}\right),\|\cdot\|_{\mathcal{Q}_{2}}\right), \quad\|F\|_{\mathcal{Q}_{2}}=\left(\int_{\Omega} \mathcal{Q}_{2}\left(x^{\prime}, F(x)\right) \mathrm{d} x\right)^{1 / 2} \tag{2.8}
\end{equation*}
$$

associated to the scalar product (with obvious notation):

$$
\left\langle F_{1}, F_{2}\right\rangle_{\mathcal{Q}_{2}}=\int_{\Omega} \mathcal{L}_{2, x^{\prime}}\left(F_{1}(x), F_{2}(x)\right) \mathrm{d} x .
$$

We define $\mathbb{P}_{1}$ and $\mathbb{P}_{2}$, respectively, as the orthogonal projections onto the following subspaces of $\mathbb{E}$ :

$$
\begin{align*}
& \mathbb{E}_{1}=\left\{x_{3} \mathcal{F}_{1}\left(x^{\prime}\right)+\mathcal{F}_{0}\left(x^{\prime}\right) ; \mathcal{F}_{1}, \mathcal{F}_{0} \in L^{2}\left(\omega, \mathbb{R}_{\mathrm{sym}}^{2 \times 2}\right)\right\}  \tag{2.9}\\
& \mathbb{E}_{2}=\left\{x_{3}^{2} \mathcal{F}_{2}\left(x^{\prime}\right)+x_{3} \mathcal{F}_{1}\left(x^{\prime}\right)+\mathcal{F}_{0}\left(x^{\prime}\right) ; \mathcal{F}_{2}, \mathcal{F}_{1}, \mathcal{F}_{0} \in L^{2}\left(\omega, \mathbb{R}_{\mathrm{sym}}^{2 \times 2}\right)\right\},
\end{align*}
$$

obtained by projecting each $F\left(x^{\prime}, \cdot\right)$ on the appropriate polynomial subspaces of $L^{2}(-1 / 2,1 / 2)$ whose orthonormal bases consist of the Legendre polynomials $\left\{p_{i}\right\}_{i=0}^{\infty}$. The first three polynomials are:

$$
p_{0}(t)=1, \quad p_{1}(t)=\sqrt{12} t, \quad p_{2}(t)=\sqrt{5}\left(6 t^{2}-\frac{1}{2}\right) .
$$

Lemma 2.2. For every $F \in \mathbb{E}$, we have:

$$
\begin{aligned}
& \mathbb{P}_{1}(F)=12\left(\int_{-1 / 2}^{1 / 2} x_{3} F \mathrm{~d} x_{3}\right) x_{3}+\left(\int_{-1 / 2}^{1 / 2} F \mathrm{~d} x_{3}\right), \\
& \mathbb{P}_{2}(F)=\left(\int_{-1 / 2}^{1 / 2}\left(180 x_{3}^{2}-15\right) F \mathrm{~d} x_{3}\right) x_{3}^{2}+12\left(\int_{-1 / 2}^{1 / 2} x_{3} F \mathrm{~d} x_{3}\right) x_{3}+\left(\int_{-1 / 2}^{1 / 2}\left(-15 x_{3}^{2}+\frac{9}{4}\right) F \mathrm{~d} x_{3}\right)
\end{aligned}
$$

Moreover, the distances from spaces $\mathbb{E}_{1}$ and $\mathbb{E}_{2}$ are given by:

$$
\begin{aligned}
\operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(F, \mathbb{E}_{1}\right)= & \int_{\omega}\left(\int_{-1 / 2}^{1 / 2} \mathcal{Q}_{2}\left(x^{\prime}, F\right) \mathrm{d} x_{3}-12 \mathcal{Q}_{2}\left(x^{\prime}, \int_{-1 / 2}^{1 / 2} x_{3} F \mathrm{~d} x_{3}\right)-\mathcal{Q}_{2}\left(x^{\prime}, \int_{-1 / 2}^{1 / 2} F \mathrm{~d} x_{3}\right)\right) \mathrm{d} x^{\prime}, \\
\operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(F, \mathbb{E}_{2}\right)= & \int_{\omega}\left(\int_{-1 / 2}^{1 / 2} \mathcal{Q}_{2}\left(x^{\prime}, F\right) \mathrm{d} x_{3}-180 \mathcal{Q}_{2}\left(x^{\prime}, \int_{-1 / 2}^{1 / 2}\left(x_{3}^{2}-\frac{1}{12}\right) F \mathrm{~d} x_{3}\right)\right. \\
& \left.\quad-12 \mathcal{Q}_{2}\left(x^{\prime}, \int_{-1 / 2}^{1 / 2} x_{3} F \mathrm{~d} x_{3}\right)-\mathcal{Q}_{2}\left(x^{\prime}, \int_{-1 / 2}^{1 / 2} F \mathrm{~d} x_{3}\right)\right) \mathrm{d} x^{\prime} .
\end{aligned}
$$

Proof. The Lemma results by a straightforward calculation:

$$
\begin{aligned}
\operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(F, \mathbb{E}_{1}\right) & =\|F\|_{\mathcal{Q}_{2}}^{2}-\left\|\mathbb{P}_{1}(F)\right\|_{\mathcal{Q}_{2}}^{2}=\|F\|_{\mathcal{Q}_{2}}^{2}-\left(\left\|\int_{-1 / 2}^{1 / 2} p_{1} F \mathrm{~d} x_{3}\right\|_{\mathcal{Q}_{2}}^{2}+\left\|\int_{-1 / 2}^{1 / 2} p_{0} F \mathrm{~d} x_{3}\right\|_{\mathcal{Q}_{2}}^{2}\right), \\
\operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(F, \mathbb{E}_{2}\right) & =\|F\|_{\mathcal{Q}_{2}}^{2}-\left\|\mathbb{P}_{2}(F)\right\|_{\mathcal{Q}_{2}}^{2} \\
& =\|F\|_{\mathcal{Q}_{2}}^{2}-\left(\left\|\int_{-1 / 2}^{1 / 2} p_{2} F \mathrm{~d} x_{3}\right\|_{\mathcal{Q}_{2}}^{2}+\left\|\int_{-1 / 2}^{1 / 2} p_{1} F \mathrm{~d} x_{3}\right\|_{\mathcal{Q}_{2}}^{2}+\left\|\int_{-1 / 2}^{1 / 2} p_{0} F \mathrm{~d} x_{3}\right\|_{\mathcal{Q}_{2}}^{2}\right),
\end{aligned}
$$

where we have used that: $\mathbb{P}_{1}(F)=p_{1} \int_{-1 / 2}^{1 / 2} p_{1} F \mathrm{~d} x_{3}+p_{0} \int_{-1 / 2}^{1 / 2} p_{0} F \mathrm{~d} x_{3}$ and similarly: $\mathbb{P}_{2}(F)=$ $p_{2} \int_{-1 / 2}^{1 / 2} p_{2} F \mathrm{~d} x_{3}+p_{1} \int_{-1 / 2}^{1 / 2} p_{1} F \mathrm{~d} x_{3}+p_{0} \int_{-1 / 2}^{1 / 2} p_{0} F \mathrm{~d} x_{3}$.

Theorem 2.3. In the setting of Theorem 2.1, $\lim _{\inf }^{h \rightarrow 0}{ }^{\frac{1}{h^{2}} \mathcal{E}^{h}\left(u^{h}\right) \text { is bounded from below by: }}$

$$
\begin{aligned}
\mathcal{I}_{2}^{O}(y) & =\frac{1}{2} \int_{\Omega} \mathcal{Q}_{2}\left(x^{\prime}, x_{3} \nabla y\left(x^{\prime}\right)^{\top} \nabla \vec{b}\left(x^{\prime}\right)-\frac{1}{2} \mathcal{G}_{1}(x)_{2 \times 2}\right) \mathrm{d} x \\
& =\frac{1}{24} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime},\left(\nabla y\left(x^{\prime}\right)^{\top} \nabla \vec{b}\left(x^{\prime}\right)\right)_{\text {sym }}-\frac{1}{2} \overline{\mathcal{G}}_{1}\left(x^{\prime}\right)_{2 \times 2}\right) \mathrm{d} x^{\prime}+\frac{1}{8} \operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(\left(\mathcal{G}_{1}\right)_{2 \times 2}, \mathbb{E}_{1}\right),
\end{aligned}
$$

where $\overline{\mathcal{G}}_{1}$ is as in (EF). In the non-oscillatory case (NO) this formula becomes:

$$
\mathcal{I}_{2}(y)=\frac{1}{24} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime},\left(\nabla y\left(x^{\prime}\right)^{\top} \nabla \vec{b}\left(x^{\prime}\right)\right)_{\text {sym }}-\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}\right) \mathrm{d} x^{\prime}
$$

The first term in $\mathcal{I}_{2}^{O}$ coincides with $\mathcal{I}_{2}$ for the effective metric $\bar{G}$ in (EF).
Proof. The argument follows the proof of [12, Theorem 2.1] and thus we only indicate its new ingredients. Applying the compactness analysis for the $x_{3}$-independent metric $\bar{G}$, one obtains the sequence $\left\{R^{h} \in L^{2}(\omega, S O(3))\right\}_{h \rightarrow 0}$ of approximating rotation-valued fields, satisfying:

$$
\begin{equation*}
\frac{1}{h} \int_{\Omega^{h}}\left|\nabla u^{h}(x) \bar{A}\left(x^{\prime}\right)^{-1}-R^{h}\left(x^{\prime}\right)\right|^{2} \mathrm{~d} x \leq C h^{2} . \tag{2.10}
\end{equation*}
$$

Define now the family $\left\{S^{h} \in L^{2}\left(\Omega, \mathbb{R}^{3 \times 3}\right)\right\}_{h \rightarrow 0}$ by:

$$
S^{h}\left(x^{\prime}, x_{3}\right)=\frac{1}{h}\left(R^{h}\left(x^{\prime}\right)^{\top} \nabla u^{h}\left(x^{\prime}, h x_{3}\right) A^{h}\left(x^{\prime}, h x_{3}\right)^{-1}-I d_{3}\right) .
$$

According to [12], the same quantities, written for the metric $\bar{G}$ rather than $G^{h}$ :

$$
\bar{S}^{h}\left(x^{\prime}, x_{3}\right)=\frac{1}{h}\left(R^{h}\left(x^{\prime}\right)^{\top} \nabla u^{h}\left(x^{\prime}, h x_{3}\right) \bar{A}\left(x^{\prime}\right)^{-1}-I d_{3}\right),
$$

converge weakly in $L^{2}\left(\Omega, \mathbb{R}^{3 \times 3}\right)$ to $\bar{S}$, such that:

$$
\begin{equation*}
\left(\bar{A}\left(x^{\prime}\right) \bar{S}\left(x^{\prime}, x_{3}\right) \bar{A}\left(x^{\prime}\right)\right)_{2 \times 2}=\bar{s}\left(x^{\prime}\right)+x_{3} \nabla y\left(x^{\prime}\right)^{\top} \nabla \vec{b}\left(x^{\prime}\right), \tag{2.11}
\end{equation*}
$$

with some appropriate $\bar{s} \in L^{2}\left(\omega, \mathbb{R}^{2 \times 2}\right)$. Observe that:

$$
S^{h}\left(x^{\prime}, x_{3}\right)=\bar{S}^{h}\left(x^{\prime}, x_{3}\right)+\frac{1}{h} R^{h}\left(x^{\prime}\right)^{\top} \nabla^{h} u\left(x^{\prime}, h x_{3}\right)\left(A^{h}\left(x^{\prime}, h x_{3}\right)^{-1}-\bar{A}\left(x^{\prime}\right)^{-1}\right)
$$

and that the term $R^{h}\left(x^{\prime}\right)^{\top} \nabla u^{h}\left(x^{\prime}, h x_{3}\right)$ converges strongly in $L^{2}\left(\Omega, \mathbb{R}^{3 \times 3}\right)$ to $\bar{A}\left(x^{\prime}\right)$. On the other hand, the remaining factor converges uniformly on $\Omega$ as $h \rightarrow 0$, because:

$$
\begin{equation*}
\frac{1}{h}\left(A^{h}\left(x^{\prime}, h x_{3}\right)^{-1}-\bar{A}\left(x^{\prime}\right)^{-1}\right)=-\bar{A}\left(x^{\prime}\right)^{-1} A_{1}\left(x^{\prime}, x_{3}\right) \bar{A}\left(x^{\prime}\right)^{-1}+\mathcal{O}(h) \tag{2.12}
\end{equation*}
$$

Concluding, $S^{h}$ converge weakly in $L^{2}\left(\Omega, \mathbb{R}^{3 \times 3}\right)$ to $S$, satisfying by (2.11):

$$
\begin{equation*}
\left(\bar{A}\left(x^{\prime}\right) S\left(x^{\prime}, x_{3}\right) \bar{A}\left(x^{\prime}\right)\right)_{2 \times 2}=\bar{s}\left(x^{\prime}\right)+x_{3} \nabla y\left(x^{\prime}\right)^{\top} \nabla \vec{b}\left(x^{\prime}\right)-\left(\bar{A}\left(x^{\prime}\right) A_{1}\left(x^{\prime}, x_{3}\right)\right)_{2 \times 2} . \tag{2.13}
\end{equation*}
$$

Consequently, using the definition of $S^{h}$ and frame invariance of $W$ and Taylor expanding $W$ at $I d_{3}$ on the set $\left\{\left|S^{h}\right|^{2} \leq 1 / h\right\}$, we obtain:

$$
\begin{aligned}
\liminf _{h \rightarrow 0} \mathcal{E}^{h}\left(u^{h}\right) & =\liminf _{h \rightarrow 0} \frac{1}{h^{2}} \int_{\Omega} W\left(I d_{3}+h S^{h}(x)\right) \mathrm{d} x \\
& \geq \liminf _{h \rightarrow 0} \frac{1}{2} \int_{\left\{\left|S^{h}\right|^{2} \leq 1 / h\right\}} \mathcal{Q}_{3}\left(S^{h}(x)\right)+o\left(\left|S^{h}\right|^{2}\right) \mathrm{d} x \\
& \geq \frac{1}{2} \int_{\Omega} \mathcal{Q}_{3}(S(x)) \mathrm{d} x \geq \frac{1}{2} \int_{\Omega} \mathcal{Q}_{2}\left(x^{\prime},\left(\bar{A}\left(x^{\prime}\right) S(x) \bar{A}\left(x^{\prime}\right)\right)_{2 \times 2}\right) \mathrm{d} x .
\end{aligned}
$$

Further, recalling (2.13) and (2.1) we get:

$$
\begin{aligned}
& \liminf _{h \rightarrow 0} \mathcal{E}^{h}\left(u^{h}\right) \geq \frac{1}{2}\left\|\bar{s}_{\text {sym }}+x_{3}\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2}\left(\mathcal{G}_{1}\right)_{2 \times 2}\right\|_{\mathcal{Q}_{2}}^{2} \\
& \quad= \frac{1}{2}\left\|\mathbb{P}_{1}\left(\bar{s}_{\text {sym }}+x_{3}\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2}\left(\mathcal{G}_{1}\right)_{2 \times 2}\right)\right\|_{\mathcal{Q}_{2}}^{2}+\frac{1}{8}\left\|\left(\mathcal{G}_{1}\right)_{2 \times 2}-\mathbb{P}_{1}\left(\left(\mathcal{G}_{1}\right)_{2 \times 2}\right)\right\|_{\mathcal{Q}_{2}}^{2} \\
& \quad= \frac{1}{2}\left\|\bar{s}_{\text {sym }}\right\|_{\mathcal{Q}_{2}}^{2}+\frac{1}{2}\left\|x_{3}\left(\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-6 \int_{-1 / 2}^{1 / 2} t\left(\mathcal{G}_{1}\right)_{2 \times 2} \mathrm{~d} t\right)\right\|_{\mathcal{Q}_{2}}^{2}+\frac{1}{8} \operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(\left(\mathcal{G}_{1}\right)_{2 \times 2}, \mathbb{E}_{1}\right) \\
& \quad \geq \frac{1}{24}\left\|\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-6 \int_{-1 / 2}^{1 / 2} t\left(\mathcal{G}_{1}\right)_{2 \times 2} \mathrm{~d} t\right\|_{\mathcal{Q}_{2}}^{2}+\frac{1}{8} \operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(\left(\mathcal{G}_{1}\right)_{2 \times 2}, \mathbb{E}_{1}\right)=\mathcal{I}_{2}^{O}(y),
\end{aligned}
$$

where we have used the fact that $\mathcal{Q}_{2}\left(x^{\prime}, \cdot\right)$ is a function of its symmetrized argument and Lemma 2.2. The formula for $\mathcal{I}_{2}$ in case (NO) is immediate.

Our next result is the upper bound, parallel to the lower bound in Theorem 2.3:

Theorem 2.4. Assume (O). For every isometric immersion $y \in W^{2,2}\left(\omega, \mathbb{R}^{3}\right)$ of the reduced midplate metric $\overline{\mathcal{G}}_{2 \times 2}$ as in (2.3), there exists a sequence $\left\{u^{h} \in W^{1,2}\left(\Omega^{h}, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0}$ such that the sequence $\left\{y^{h}\left(x^{\prime}, x_{3}\right)=u^{h}\left(x^{\prime}, h x_{3}\right)\right\}_{h \rightarrow 0}$ converges in $W^{1,2}\left(\Omega, \mathbb{R}^{3}\right)$ to $y$ and:

$$
\begin{equation*}
\lim _{h \rightarrow 0} \frac{1}{h^{2}} \mathcal{E}^{h}\left(u^{h}\right)=\mathcal{I}_{2}^{O}(y) \tag{2.14}
\end{equation*}
$$

Automatically, $\frac{1}{h} \partial_{3} y^{h}$ converges in $L^{2}\left(\Omega, \mathbb{R}^{3}\right)$ to $\vec{b} \in W^{1,2} \cap L^{\infty}\left(\omega, \mathbb{R}^{3}\right)$ as in (2.4).
Proof. Given an admissible $y$, we define $\vec{b}$ by (2.4) and also define the matrix field:

$$
\begin{equation*}
Q=\left[\partial_{1} y, \partial_{2} y, \vec{b}\right] \in W^{1,2} \cap L^{\infty}\left(\omega, \mathbb{R}^{3 \times 3}\right) \tag{2.15}
\end{equation*}
$$

It follows that $Q\left(x^{\prime}\right) \bar{A}\left(x^{\prime}\right)^{-1} \in S O(3)$ on $\omega$. The recovery sequence $u^{h}$ satisfying (2.14) is then constructed via a diagonal argument, applied to the explicit deformation fields below. Again, we only indicate the new ingredients with respect to the proof in [12, Theorem 3.1].

Recalling the notion of the linear vector association $c\left(x^{\prime}, F_{2 \times 2}\right)$ that is minimizing in the right hand side of the formula (2.7), we define the vector field $\vec{d} \in L^{2}\left(\Omega, \mathbb{R}^{3}\right)$ by:

$$
\begin{align*}
\vec{d}\left(x^{\prime}, x_{3}\right)=Q\left(x^{\prime}\right)^{\mathrm{\top},-1}( & \frac{x_{3}^{2}}{2}\left(c\left(x^{\prime}, \nabla y\left(x^{\prime}\right)^{\top} \nabla \vec{b}\left(x^{\prime}\right)\right)-\frac{1}{2}\left[\begin{array}{c}
\nabla|\vec{b}|^{2}\left(x^{\prime}\right) \\
0
\end{array}\right]\right) \\
& -\frac{1}{2} c\left(x^{\prime}, \int_{0}^{x_{3}} \mathcal{G}_{1}\left(x^{\prime}, t\right)_{2 \times 2} \mathrm{~d} t\right)  \tag{2.16}\\
& \left.+\int_{0}^{x_{3}} \mathcal{G}_{1}\left(x^{\prime}, t\right) \mathrm{d} t e_{3}-\frac{1}{2} \int_{0}^{x_{3}} \mathcal{G}_{1}\left(x^{\prime}, t\right)_{33} \mathrm{~d} t e_{3}\right) .
\end{align*}
$$

In view of (2.7), the above definition is equivalent to the vector field $\partial_{3} \vec{d} \in L^{2}\left(\Omega, \mathbb{R}^{3}\right)$ being, for each $\left(x^{\prime}, x_{3}\right) \in \Omega$, the unique solution to:

$$
\begin{aligned}
& \mathcal{Q}_{2}\left(x^{\prime}, x_{3} \nabla y\left(x^{\prime}\right)^{\top} \nabla \vec{b}\left(x^{\prime}\right)-\frac{1}{2} \mathcal{G}_{1}\left(x^{\prime}, x_{3}\right)_{2 \times 2}\right) \\
& \quad=\mathcal{Q}_{3}\left(\bar{A}\left(x^{\prime}\right)^{-1}\left(Q\left(x^{\prime}\right)^{\top}\left[x_{3} \partial_{1} \vec{b}\left(x^{\prime}\right), x_{3} \partial_{2} \vec{b}\left(x^{\prime}\right), \partial_{3} \vec{d}\left(x^{\prime}, x_{3}\right)\right]-\frac{1}{2} \mathcal{G}_{1}\left(x^{\prime}, x_{3}\right)\right) \bar{A}\left(x^{\prime}\right)^{-1}\right) .
\end{aligned}
$$

One then approximates $y, \vec{b}$ by sequences $\left\{y^{h} \in W^{2, \infty}\left(\omega, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0},\left\{b^{h} \in W^{1, \infty}\left(\omega, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0}$ respectively, and request them to satisfy conditions exactly as in the proof of [12, Theorem 3.1]. The warping field $\vec{d}$ is approximated by $d^{h}\left(x^{\prime}, x_{3}\right)=\int_{0}^{x_{3}} \vec{d}^{h}\left(x^{\prime}, t\right) \mathrm{d} t$, where:

$$
\bar{d}^{h} \rightarrow \bar{d}=\partial_{3} \vec{d} \quad \text { strongly in } L^{2}\left(\Omega, \mathbb{R}^{3}\right) \quad \text { and } \quad h\left\|\bar{d}^{h}\right\|_{W^{1, \infty}\left(\Omega, \mathbb{R}^{3}\right)} \rightarrow 0 \quad \text { as } h \rightarrow 0
$$

Finally, we define:

$$
\begin{equation*}
u^{h}\left(x^{\prime}, x_{3}\right)=y^{h}\left(x^{\prime}\right)+x_{3} b^{h}\left(x^{\prime}\right)+h^{2} d^{h}\left(x^{\prime}, \frac{x_{3}}{h}\right), \tag{2.17}
\end{equation*}
$$

so that, with the right approximation error, there holds:

$$
\nabla u^{h}\left(x^{\prime}, x_{3}\right) \approx Q\left(x^{\prime}\right)+h\left[\frac{x_{3}}{h} \partial_{1} \vec{b}\left(x^{\prime}\right), \frac{x_{3}}{h} \partial_{2} \vec{b}\left(x^{\prime}\right), \partial_{3} \vec{d}\left(x^{\prime}, \frac{x_{3}}{h}\right)\right] .
$$

Using Taylor's expansion of $W$, the definition (2.16) and the controlled blow-up rates of the approximating sequences, the construction is done.

We conclude this section by noting the following easy direct consequence of Theorems 2.3 and 2.4:

Corollary 2.5. If the set of $W^{2,2}\left(\omega, \mathbb{R}^{3}\right)$ isometric immersions of $\overline{\mathcal{G}}_{2 \times 2}$ is nonempty, then the functional $\mathcal{I}_{2}^{O}$ attains its infimum and:

$$
\lim _{h \rightarrow 0} \frac{1}{h^{2}} \inf \mathcal{E}^{h}=\min \mathcal{I}_{2}^{0}
$$

The infima in the left hand side are taken over $W^{1,2}\left(\Omega^{h}, \mathbb{R}^{3}\right)$ deformations $u^{h}$, whereas the minima in the right hand side are taken over $W^{2,2}\left(\omega, \mathbb{R}^{3}\right)$ isometric immersions y of $\overline{\mathcal{G}}_{2 \times 2}$.

## 3. Identification of the $C h^{2}$ scaling regime

In this section, we identify the equivalent conditions for $\inf \mathcal{E}^{h} \sim h^{2}$ in terms of curvatures of the metric tensor $\bar{G}$ in case (NO). We begin by expressing the integrand tensor in the residual energy $\mathcal{I}_{2}$ in terms of the shape operator on the deformed midplate. Recall that we always use the Einstein summation convention over repeated indices running from 1 to 3 .
Lemma 3.1. In the the non-oscillatory setting (NO), let $y \in W^{2,2}\left(\omega, \mathbb{R}^{3}\right)$ be an isometric immersion of the metric $\overline{\mathcal{G}}_{2 \times 2}$, so that (2.3) holds on $\omega$. Define the Cosserat vector $\vec{b}$ according to (2.4). Then:

$$
\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}=\frac{1}{\sqrt{\overline{\mathcal{G}}^{33}}} \Pi_{y}+\frac{1}{\overline{\mathcal{G}}^{33}}\left[\begin{array}{cc}
\Gamma_{11}^{3} & \Gamma_{12}^{3}  \tag{3.1}\\
\Gamma_{12}^{3} & \Gamma_{22}^{3}
\end{array}\right]\left(x^{\prime}, 0\right),
$$

for all $x^{\prime} \in \omega$. Above, $\overline{\mathcal{G}}^{33}=\left\langle\overline{\mathcal{G}}^{-1} e_{3}, e_{3}\right\rangle$, whereas $\Pi_{y}=(\nabla y)^{\top} \nabla \vec{\nu} \in W^{1,2}\left(\omega, \mathbb{R}_{\text {sym }}^{2 \times 2}\right)$ is the second fundamental form of the surface $y(\omega) \subset \mathbb{R}^{3}$, and $\left\{\Gamma_{k l}^{i}\right\}_{i, k, l=1 \ldots 3}$ are the Christoffel symbols of $G$ :

$$
\Gamma_{k l}^{i}=\frac{1}{2} G^{i m}\left(\partial_{l} G_{m k}+\partial_{k} G_{m l}-\partial_{m} G_{k l}\right) .
$$

Proof. The proof is an extension of the arguments in [12, Theorem 5.3], which we modify for the case of $x_{3}$-dependent metric $G$. Firstly, the fact that $Q^{\top} Q=\overline{\mathcal{G}}$ with $Q$ defined in (2.15), yields:

$$
\begin{equation*}
\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}=\left(\left[\partial_{i} \overline{\mathcal{G}}_{j 3}\right]_{i, j=1,2}\right)_{\text {sym }}-\left[\left\langle\partial_{i j} y, \vec{b}\right\rangle\right]_{i, j=1,2} \tag{3.2}
\end{equation*}
$$

Also, $\partial_{i} \overline{\mathcal{G}}=2\left(\left(\partial_{i} Q\right)^{\top} Q\right)_{\text {sym }}$ for $i=1,2$, results in:

$$
\begin{equation*}
\left\langle\partial_{i j} y, \partial_{k} y\right\rangle=\frac{1}{2}\left(\partial_{i} G_{k j}+\partial_{j} G_{i k}-\partial_{k} G_{i j}\right) \tag{3.3}
\end{equation*}
$$

and:

$$
(\nabla y)^{\top} \partial_{i j} y=\Gamma_{i j}^{m}\left(x^{\prime}, 0\right)\left[\begin{array}{l}
\overline{\mathcal{G}}_{m 1} \\
\overline{\mathcal{G}}_{m 2}
\end{array}\right] \quad \text { for } i, j=1,2
$$

Consequently, we obtain the formula:

$$
\begin{aligned}
{\left[\overline{\mathcal{G}}_{13}, \overline{\mathcal{G}}_{23}\right]\left(\overline{\mathcal{G}}_{2 \times 2}\right)^{-1}(\nabla y)^{\top} \partial_{i j} y } & =\left[\overline{\mathcal{G}}_{13}, \overline{\mathcal{G}}_{23},\left[\overline{\mathcal{G}}_{13}, \overline{\mathcal{G}}_{23}\right]\left(\overline{\mathcal{G}}_{2 \times 2}\right)^{-1}\left[\begin{array}{c}
\overline{\mathcal{G}}_{13} \\
\overline{\mathcal{G}}_{23}
\end{array}\right]\right]\left[\begin{array}{c}
\Gamma_{i j}^{1} \\
\Gamma_{i j}^{2} \\
\Gamma_{i j}^{3}
\end{array}\right]\left(x^{\prime}, 0\right) \\
& =\overline{\mathcal{G}}_{m 3} \Gamma_{i j}^{m}\left(x^{\prime}, 0\right)-\frac{1}{\overline{\mathcal{G}}^{33}} \Gamma_{i j}^{3}\left(x^{\prime}, 0\right) .
\end{aligned}
$$

Computing the normal vector $\vec{\nu}$ from (2.4) and noting that $\operatorname{det} \overline{\mathcal{G}}_{2 \times 2} / \operatorname{det} \overline{\mathcal{G}}=\overline{\mathcal{G}}^{33}$, we get:

$$
\begin{aligned}
\Pi_{i j} & =-\left\langle\partial_{i j} y, \vec{\nu}\right\rangle=-\sqrt{\overline{\mathcal{G}}^{33}}\left(\left\langle\partial_{i j} y, \vec{b}\right\rangle-\left[\overline{\mathcal{G}}_{13}, \overline{\mathcal{G}}_{23}\right]\left(\overline{\mathcal{G}}_{2 \times 2}\right)^{-1}(\nabla y)^{\top} \partial_{i j} y\right) \\
& =\sqrt{\overline{\mathcal{G}}^{33}}\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\operatorname{sym}, i j}-\frac{1}{\sqrt{\overline{\mathcal{G}}^{33}}} \Gamma_{i j}^{3}\left(x^{\prime}, 0\right)-\frac{\sqrt{\overline{\mathcal{G}}^{33}}}{2} \partial_{3} G_{i j}\left(x^{\prime}, 0\right), \quad \text { for } i, j=1,2,
\end{aligned}
$$

which completes the proof of (3.1).
The key result of this section is the following:
Theorem 3.2. The energy scaling beyond the Kirchhoff regime:

$$
\lim _{h \rightarrow 0} \frac{1}{h^{2}} \inf \mathcal{E}^{h}=0
$$

is equivalent to the following conditions:
(i) in the oscillatory case ( $O$ )
$\left[\begin{array}{rl}\left(\mathcal{G}_{1}\right)_{2 \times 2} \in \mathbb{E}_{1} \text { or equivalently there holds: } \\ \qquad \mathcal{G}_{1}\left(x^{\prime}, x_{3}\right)_{2 \times 2}=x_{3} \overline{\mathcal{G}}_{1}\left(x^{\prime}\right)_{2 \times 2} & \text { for all }\left(x^{\prime}, x_{3}\right) \in \bar{\Omega} .\end{array}\right.$
Moreover, condition (3.5) below must be satisfied with $G$ replaced by the effective metric $\left[\bar{G}\right.$ in $(E F)$. This condition involves only $\overline{\mathcal{G}}$ and $\left(\overline{\mathcal{G}}_{1}\right)_{2 \times 2}$ terms of $\bar{G}$.
(ii) in the non-oscillatory case (NO)
$\left[\begin{array}{l}\text { There exists } y_{0} \in W^{2,2}\left(\omega, \mathbb{R}^{3}\right) \text { satisfying (2.3) and such that: } \\ \qquad \Pi_{y_{0}}\left(x^{\prime}\right)=-\frac{1}{\sqrt{\overline{\mathcal{G}}^{33}}}\left[\begin{array}{cc}\Gamma_{11}^{3} & \Gamma_{12}^{3} \\ \Gamma_{12}^{3} & \Gamma_{22}^{3}\end{array}\right]\left(x^{\prime}, 0\right) \quad \text { for all } x^{\prime} \in \omega,\end{array}\right.$
where $\Pi_{y_{0}}$ is the second fundamental form of the surface $y_{0}(\omega)$ and $\left\{\Gamma_{j k}^{i}\right\}$ are the Christoffel symbols of the metric $G$.

The isometric immersion $y_{0}$ in (3.5) is automatically smooth (up to the boundary) and it is unique up to rigid motions. Further, on a simply connected midplate $\omega$, condition (3.5) is equivalent to:

$$
\begin{align*}
& \text { The following Riemann curvatures of the metric } G \text { vanish on } \omega \times\{0\} \text { : }  \tag{3.6}\\
& \qquad R_{1212}\left(x^{\prime}, 0\right)=R_{1213}\left(x^{\prime}, 0\right)=R_{1223}\left(x^{\prime}, 0\right)=0 \quad \text { for all } x^{\prime} \in \omega .
\end{align*}
$$

Above, the Riemann curvatures of a given metric $G$ are:

$$
R_{i k l m}=\frac{1}{2}\left(\partial_{k l} G_{i m}+\partial_{i m} G_{k l}-\partial_{k m} G_{i l}-\partial_{i l} G_{k m}\right)+G_{n p}\left(\Gamma_{k l}^{n} \Gamma_{i m}^{p}-\Gamma_{k m}^{n} \Gamma_{i l}^{p}\right)
$$

Proof. By Corollary 2.5, it suffices to determine the equivalent conditions for $\min \mathcal{I}_{2}^{O}=0$ and $\min \mathcal{I}_{2}=0$. In case (O), the linearity of $x_{3} \mapsto \mathcal{G}_{1}\left(x^{\prime}, x_{3}\right)_{2 \times 2}$ is immediate, while condition (3.5) follows in both cases (O) and (NO) by Lemma 3.1. Note that the Christoffel symbols $\left\{\Gamma_{j k}^{i}\right\}$ depend only on $\overline{\mathcal{G}}$ and $\partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}$ in the Taylor expansion of $G$. This completes the proof of (i) and (ii).

Regularity of $y_{0}$ is an easy consequence, via the bootstrap argument, of the continuity equation:

$$
\begin{equation*}
\partial_{i j} y_{0}=\sum_{m=1}^{2} \gamma_{i j}^{m} \partial_{m} y_{0}-\left(\Pi_{y_{0}}\right)_{i j} \vec{\nu}_{0} \quad \text { for } i, j=1,2, \tag{3.7}
\end{equation*}
$$

where $\left\{\gamma_{i j}^{m}\right\}_{i, j, m=1 \ldots 2}$ denote the Christoffel symbols of $\overline{\mathcal{G}}_{2 \times 2}$ on $\omega$. Uniqueness of $y_{0}$ is a consequence of (3.5), due to uniqueness of isometric immersion with prescribed second fundamental form.

To show (3.6), we argue as in the proof of [12, Theorem 5.5]. The compatibility of $\bar{G}_{2 \times 2}$ and $\Pi_{y_{0}}$ is equivalent to the satisfaction of the related Gauss-Codazzi-Mainardi equations. By an explicit
calculation, we see that the two Codazzi-Mainardi equations become:

$$
\begin{array}{r}
\left(\partial_{2} \Gamma_{11}^{3}-\partial_{1} \Gamma_{12}^{3}\right)-\frac{1}{2}\left(\frac{\partial_{2} G^{33}}{G^{33}} \Gamma_{11}^{3}-\frac{\partial_{1} G^{33}}{G^{33}} \Gamma_{12}^{3}\right)+\frac{1}{G^{33}} G^{m 3}\left(\Gamma_{2 m}^{3} \Gamma_{11}^{3}-\Gamma_{1 m}^{3} \Gamma_{12}^{3}\right) \\
=\left(\sum_{m=1}^{2} \Gamma_{1 m}^{3} \Gamma_{12}^{m}-\sum_{m=1}^{2} \Gamma_{2 m}^{3} \Gamma_{11}^{m}\right)+\frac{G^{32}}{G^{33}}\left(\Gamma_{11}^{3} \Gamma_{22}^{3}-\left(\Gamma_{12}^{3}\right)^{2}\right), \\
\left(\partial_{2} \Gamma_{12}^{3}-\partial_{1} \Gamma_{22}^{3}\right)-\frac{1}{2}\left(\frac{\partial_{2} G^{33}}{G^{33}} \Gamma_{12}^{3}-\frac{\partial_{1} G^{33}}{G^{33}} \Gamma_{22}^{3}\right)+\frac{1}{G^{33}} G^{m 3}\left(\Gamma_{2 m}^{3} \Gamma_{12}^{3}-\Gamma_{1 m}^{3} \Gamma_{22}^{3}\right) \\
=\left(\sum_{m=1}^{2} \Gamma_{1 m}^{3} \Gamma_{22}^{m}-\sum_{m=1}^{2} \Gamma_{2 m}^{3} \Gamma_{12}^{m}\right)-\frac{G^{31}}{G^{33}}\left(\Gamma_{11}^{3} \Gamma_{22}^{3}-\left(\Gamma_{12}^{3}\right)^{2}\right),
\end{array}
$$

and are equivalent to $R_{121}^{3}=R_{221}^{3}=0$ on $\omega \times\{0\}$. The Gauss equation is, in turn, equivalent to $R_{1212}=0$ exactly as in [12]. The simultaneous vanishing of $R_{121}^{3}, R_{221}^{3}, R_{1212}$ is equivalent with the vanishing of $R_{1212}, R_{1213}$ and $R_{1223}$, which proves the claim in (3.6).

## 4. Coercivity of the limiting energy $\mathcal{I}_{2}$

In this section we quantify the statement in Theorem 3.2 and prove that when either of $\mathcal{I}_{2}$ or $\mathcal{I}_{2}^{O}$ can be minimized to zero, the effective energy $\mathcal{I}_{2}(y)$ measures then the distance of a given isometric immersion $y$ from the kernel: $\operatorname{ker} \mathcal{I}_{2}=\left\{Q y_{0}+d ; Q \in S O(3), d \in \mathbb{R}^{3}\right\}$.

Assume that the set of $W^{2,2}\left(\omega, \mathbb{R}^{3}\right)$ isometric immersions $y$ of $\overline{\mathcal{G}}_{2 \times 2}$ is nonempty, which in view of Theorems 2.3 and 2.4 is equivalent to: $\inf \mathcal{E}^{h} \leq C h^{2}$. For each such $y$, the continuity equation (3.7) combined with Lemma 3.1 gives the following formula, valid for all $i, j=1,2$ :

$$
\begin{equation*}
\partial_{i j} y=\sum_{m=1,2} \gamma_{i j}^{m} \partial_{m} y-\sqrt{\overline{\mathcal{G}}^{33}}\left(\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}\right)_{i j} \vec{\nu}+\frac{\Gamma_{i j}^{3}}{\sqrt{\overline{\mathcal{G}}^{33}}} \vec{\nu} \quad \text { on } \omega . \tag{4.1}
\end{equation*}
$$

Another consequence of (3.7) is:

$$
\left|\nabla^{2} y\right|^{2}=\left|\Pi_{y}\right|^{2}+\sum_{i, j=1,2}\left\langle\overline{\mathcal{G}}_{2 \times 2}:\left[\gamma_{i j}^{1}, \gamma_{i j}^{2}\right]^{\otimes 2}\right\rangle \quad \text { on } \omega .
$$

By Lemma 3.1 and since $|\nabla y|^{2}=$ trace $\overline{\mathcal{G}}_{2 \times 2}$, this yields the bound:

$$
\begin{equation*}
\left\|y-f_{\omega} y\right\|_{W^{2,2}\left(\omega, \mathbb{R}^{3}\right)}^{2} \leq C\left(\mathcal{I}_{2}(y)+1\right) \tag{4.2}
\end{equation*}
$$

where $C$ is a constant independent of $y$. Clearly, when condition (3.6) does not hold, so that $\min \mathcal{I}_{2}>0$, the right hand side $C\left(\mathcal{I}_{2}(y)+1\right)$ above may be replaced by $C \mathcal{I}_{2}(y)$. On the other hand, in presence of (3.6), the bound (4.2) can be refined to the following coercivity result:

Theorem 4.1. Assume the curvature condition (3.6) on a metric $G$ as in (NO), and let $y_{0}$ be the unique (up to rigid motions in $\mathbb{R}^{3}$ ) isometric immersion of $\overline{\mathcal{G}}_{2 \times 2}$ satisfying (3.5). Then, for all $y \in W^{2,2}\left(\omega, \mathbb{R}^{3}\right)$ such that $(\nabla y)^{\top} \nabla y=\overline{\mathcal{G}}_{2 \times 2}$, there holds:

$$
\begin{equation*}
\operatorname{dist}_{W^{2,2}\left(\omega, \mathbb{R}^{3}\right)}^{2}\left(y,\left\{R y_{0}+c ; R \in S O(3), c \in \mathbb{R}^{3}\right\}\right) \leq C \mathcal{I}_{2}(y) \tag{4.3}
\end{equation*}
$$

with a constant $C>0$ that depends on $G, \omega$ and $W$ but is independent of $y$.

Proof. Without loss of generality, we set $f_{\omega} y=f_{\omega} y_{0}=0$. For any $R \in S O$ (3), identity (4.1) implies:

$$
\begin{aligned}
\int_{\omega}\left|\nabla^{2} y-\nabla^{2}\left(R y_{0}\right)\right|^{2} \mathrm{~d} x^{\prime} \leq C( & \int_{\omega}\left|\nabla y-\nabla\left(R y_{0}\right)\right|^{2} \mathrm{~d} x^{\prime} \\
& \left.+\int_{\omega}\left|\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}\right|^{2} \mathrm{~d} x^{\prime}+\int_{\omega}\left|\vec{\nu}-R \vec{\nu}_{0}\right|^{2} \mathrm{~d} x^{\prime}\right) \\
\leq C( & \left.\int_{\omega}\left|\nabla y-\nabla\left(R y_{0}\right)\right|^{2} \mathrm{~d} x^{\prime}+\int_{\omega}\left|\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}\right|^{2} \mathrm{~d} x^{\prime}\right)
\end{aligned}
$$

where we used $\mathcal{I}_{2}\left(R y_{0}\right)=0$ and the fact that $\int_{\omega}\left|\vec{\nu}-R \vec{\nu}_{0}\right|^{2} \mathrm{~d} x^{\prime} \leq C \int_{\omega}\left|\nabla y-\nabla\left(R y_{0}\right)\right|^{2} \mathrm{~d} x^{\prime}$ following, in particular, from $\left|\partial_{1} y \times \partial_{2} y\right|=\left|\partial_{1}\left(R y_{0}\right) \times \partial_{2}\left(R y_{0}\right)\right|=\sqrt{\operatorname{det} \overline{\mathcal{G}}_{2 \times 2}}$. Also, the non-degeneracy of quadratic forms $\mathcal{Q}_{2}\left(x^{\prime}, \cdot\right)$ in (2.6), implies the uniform bound:

$$
\int_{\omega}\left|\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}\right|^{2} \mathrm{~d} x^{\prime} \leq C \mathcal{I}_{2}(y) .
$$

Taking $R \in S O(3)$ as in Lemma 4.2 below, (4.3) directly follows in view of (4.4).
The next weak coercivity estimate has been the essential part of Theorem 4.1:
Lemma 4.2. Let $y$ and $y_{0}$ be as in Theorem 4.1. Then there exists $R \in S O(3)$ such that:

$$
\begin{equation*}
\int_{\omega}\left|\nabla y-R \nabla y_{0}\right|^{2} \mathrm{~d} x^{\prime} \leq C \int_{\omega}\left|\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}\right|^{2} \mathrm{~d} x^{\prime} \tag{4.4}
\end{equation*}
$$

with a constant $C>0$ that depends on $G, \omega$ but it is independent of $y$.
Proof. Consider the natural extensions $u$ and $u_{0}$ of $y$ and $y_{0}$, namely:

$$
u\left(x^{\prime}, x_{3}\right)=y\left(x^{\prime}\right)+x_{3} \vec{b}\left(x^{\prime}\right), \quad u_{0}\left(x^{\prime}, x_{3}\right)=y_{0}\left(x^{\prime}\right)+x_{3} \vec{b}_{0}\left(x^{\prime}\right) \quad \text { for all }\left(x^{\prime}, x_{3}\right) \in \Omega^{h} .
$$

Clearly, $u \in W^{1,2}\left(\Omega^{h}, \mathbb{R}^{3}\right)$ and $u_{0} \in \mathcal{C}^{1}\left(\bar{\Omega}^{h}, \mathbb{R}^{3}\right)$ satisfies $\operatorname{det} \nabla u_{0}>0$ for $h$ sufficiently small. Write:

$$
\omega=\bigcup_{k=1}^{N} \omega_{k}, \quad \Omega^{h}=\bigcup_{k=1}^{N} \Omega_{k}^{h}
$$

as the union of $N \geq 1$ open, bounded, connected domains with Lipschitz boundary, such that on each $\left\{\Omega_{k}^{h}=\omega_{k} \times\left(-\frac{h}{2}, \frac{h}{2}\right)\right\}_{k=1}^{N}$, the deformation $u_{0 \mid \Omega_{k}^{h}}$ is a $\mathcal{C}^{1}$ diffeomorphism onto its image $\mathcal{U}_{k}^{h} \subset \mathbb{R}^{3}$.

1. We first prove (4.4) under the assumption $N=1$. Call $v=u \circ u_{0}^{-1} \in W^{1,2}\left(\mathcal{U}^{h}, \mathbb{R}^{3}\right)$ and apply the geometric rigidity estimate [17] for the existence of $R \in S O(3)$ satisfying:

$$
\begin{equation*}
\int_{\mathcal{U}^{h}}|\nabla v-R|^{2} \mathrm{~d} z \leq C \int_{\mathcal{U}^{h}} \operatorname{dist}^{2}(\nabla v, S O(3)) \mathrm{d} z \tag{4.5}
\end{equation*}
$$

with a constant $C$ depending on a particular choice of $h$ (and ultimately $k$, when $N>1$ ), but independent of $v$. Since $\nabla v\left(u_{0}(x)\right)=\nabla u(x)\left(\nabla u_{0}(x)\right)^{-1}$ for all $x \in \Omega^{h}$, we get:

$$
\begin{align*}
\int_{\mathcal{U}^{h}}|\nabla v-R|^{2} \mathrm{~d} z & =\int_{\Omega^{h}}\left(\operatorname{det} \nabla u_{0}\right)\left|\left(\nabla u-R \nabla u_{0}\right)\left(\nabla u_{0}\right)^{-1}\right|^{2} \mathrm{~d} x \geq C \int_{\Omega^{h}}\left|\nabla u-R \nabla u_{0}\right|^{2} \mathrm{~d} x \\
& =C \int_{\Omega^{h}}\left|\left[\partial_{1} y, \partial_{2} y, \vec{b}\right]-R\left[\partial_{1} y_{0}, \partial_{2} y_{0}, \vec{b}_{0}\right]\right|^{2}+x_{3}^{2}\left|\nabla \vec{b}-R \nabla \overrightarrow{b_{0}}\right|^{2} \mathrm{~d} x  \tag{4.6}\\
& \geq C h \int_{\omega}\left|\nabla y-R \nabla y_{0}\right|^{2} \mathrm{~d} x^{\prime} .
\end{align*}
$$

Likewise, the change of variables in the right hand side of (4.5) gives:

$$
\begin{equation*}
\int_{\mathcal{U}^{h}} \operatorname{dist}^{2}(\nabla v, S O(3)) \mathrm{d} z \leq C \int_{\Omega^{h}} \operatorname{dist}^{2}\left((\nabla u)\left(\nabla u_{0}\right)^{-1}, S O(3)\right) \mathrm{d} x . \tag{4.7}
\end{equation*}
$$

Since $(\nabla u)^{\top} \nabla u\left(x^{\prime}, 0\right)=\left(\nabla u_{0}\right)^{\top} \nabla u_{0}\left(x^{\prime}, 0\right)=\overline{\mathcal{G}}\left(x^{\prime}\right)$, by polar decomposition it follows that: $\nabla u\left(x^{\prime}, 0\right)=$ $Q\left(x^{\prime}\right)=\bar{R} \overline{\mathcal{G}}^{1 / 2}$ and $\nabla u_{0}\left(x^{\prime}, 0\right)=Q_{0}\left(x^{\prime}\right)=\bar{R}_{0} \overline{\mathcal{G}}^{1 / 2}$ for some $\bar{R}, \bar{R}_{0} \in S O(3)$. The notation $Q, Q_{0}$ is consistent with that introduced in (2.15). Observe further:

$$
\begin{aligned}
\nabla u\left(x^{\prime}, x_{3}\right) & =Q+x_{3}\left[\partial_{1} \vec{b}, \partial_{2} \vec{b}, 0\right]=\bar{R} \overline{\mathcal{G}}^{1 / 2}\left(I d_{3}+x_{3} \overline{\mathcal{G}}^{-1} Q^{\top}\left[\partial_{1} \vec{b}, \partial_{2} \vec{b}, 0\right]\right) \\
& =\bar{R} \overline{\mathcal{G}}^{1 / 2}\left(I d_{3}+x_{3} \overline{\mathcal{G}}^{-1}\left(\left((\nabla y)^{\top} \nabla \vec{b}\right)^{*}+e_{3} \otimes[\nabla \vec{b}, 0]^{\top} \vec{b}\right)\right),
\end{aligned}
$$

and similarly:

$$
\nabla u_{0}\left(x^{\prime}, x_{3}\right)=\bar{R}_{0} \overline{\mathcal{G}}^{1 / 2}\left(I d_{3}+x_{3} \overline{\mathcal{G}}^{-1}\left(\left(\left(\nabla y_{0}\right)^{\top} \nabla \vec{b}_{0}\right)^{*}+e_{3} \otimes\left[\nabla \vec{b}_{0}, 0\right]^{\top} \vec{b}_{0}\right)\right)
$$

Consequently, the integrand in the right hand side of (4.7) becomes:

$$
\begin{align*}
& (\nabla u)\left(\nabla u_{0}\right)^{-1} \\
& \quad=\bar{R} \overline{\mathcal{G}}^{1 / 2}\left(I d_{3}+x_{3} \overline{\mathcal{G}}^{-1} S\left(I d_{3}+x_{3} \overline{\mathcal{G}}^{-1}\left(\left(\nabla y_{0}\right)^{\top} \nabla \vec{b}_{0}\right)^{*}+e_{3} \otimes\left[\nabla \vec{b}_{0}, 0\right]^{\top} \vec{b}_{0}\right)^{-1}\right) \overline{\mathcal{G}}^{-1 / 2} \bar{R}_{0}^{\top}, \tag{4.8}
\end{align*}
$$

where:

$$
S=\left((\nabla y)^{\top} \nabla \vec{b}-\left(\nabla y_{0}\right)^{\top} \nabla \vec{b}_{0}\right)^{*}+e_{3} \otimes[\nabla \vec{b}, 0]^{\top} \vec{b}-e_{3} \otimes\left[\nabla \vec{b}_{0}, 0\right]^{\top} \vec{b}_{0}=\left((\nabla y)^{\top} \nabla \vec{b}-\left(\nabla y_{0}\right)^{\top} \nabla \vec{b}_{0}\right)_{\mathrm{sym}}^{*}
$$

The last equality follows from the easy facts that, for $i, j=1,2$, we have:

$$
\begin{aligned}
& \left\langle\partial_{i} \vec{b}, \vec{b}\right\rangle=\left\langle\partial_{i} \vec{b}_{0}, \vec{b}_{0}\right\rangle=\frac{1}{2} \partial_{i} \overline{\mathcal{G}}_{33} \\
& \left\langle\partial_{i} y, \partial_{j} \vec{b}\right\rangle-\left\langle\partial_{j} y, \partial_{i} \vec{b}\right\rangle=\left\langle\partial_{i} y_{0}, \partial_{j} \vec{b}_{0}\right\rangle-\left\langle\partial_{j} y_{0}, \partial_{i} \vec{b}_{0}\right\rangle=\partial_{j} \overline{\mathcal{G}}_{i 3}-\partial_{i} \overline{\mathcal{G}}_{j 3} .
\end{aligned}
$$

Thus, (4.7) and (4.8) imply:

$$
\begin{align*}
\int_{\mathcal{U}^{h}} \operatorname{dist}^{2}(\nabla v, S O(3)) \mathrm{d} z & \leq C \int_{\Omega^{h}}\left|(\nabla u)\left(\nabla u_{0}\right)^{-1}-\bar{R} \bar{R}_{0}^{\top}\right|^{2} \mathrm{~d} x \leq C \int_{\Omega^{h}}\left|x_{3} S\left(x^{\prime}, x_{3}\right)\right|^{2} \mathrm{~d} x \\
& \leq C \int_{\omega}\left|\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\left(\left(\nabla y_{0}\right)^{\top} \nabla \vec{b}_{0}\right)_{\text {sym }}\right|^{2} \mathrm{~d} x^{\prime}  \tag{4.9}\\
& =C \int_{\omega}\left|\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}\right|^{2} \mathrm{~d} x^{\prime}
\end{align*}
$$

with a constant $C$ that depends on $G, \omega$ and $h$, but not on $y$. We conclude (4.4) in view of (4.5), (4.6) and (4.9).
2. To prove (4.4) in case $N>1$, let $k, s: 1 \ldots N$ be such that $\omega_{k} \cap \omega_{s} \neq \emptyset$. Define:

$$
F=\left(\int_{\Omega_{k}^{h} \cap \Omega_{s}^{h}} \operatorname{det} \nabla u_{0} \mathrm{~d} x\right)^{-1} \int_{\Omega_{k}^{h} \cap \Omega_{s}^{h}}\left(\operatorname{det} \nabla u_{0}\right)(\nabla u)\left(\nabla u_{0}\right)^{-1} \mathrm{~d} x \in \mathbb{R}^{3 \times 3}
$$

Denote by $R_{k}, R_{s} \in S O(3)$ the corresponding rotations in (4.4) on $\omega_{k}, \omega_{s}$. For $i \in\{k, s\}$ we have:

$$
\begin{aligned}
\left|F-R_{i}\right|^{2} & =\left|\left(\int_{\Omega_{k}^{h} \cap \Omega_{s}^{h}} \operatorname{det} \nabla u_{0} \mathrm{~d} x\right)^{-1} \int_{\Omega_{k}^{h} \cap \Omega_{s}^{h}}\left(\operatorname{det} \nabla u_{0}\right)\left(\nabla u-R_{i} \nabla u_{0}\right)\left(\nabla u_{0}\right)^{-1} \mathrm{~d} x\right|^{2} \\
& \leq C \int_{\Omega_{k}^{h} \cap \Omega_{s}^{h}}\left|\nabla u-R_{i} \nabla u_{0}\right|^{2} \mathrm{~d} x \leq C \int_{\Omega_{i}^{h}}\left|\nabla u-R_{i} \nabla u_{0}\right|^{2} \mathrm{~d} x \\
& \leq \int_{\omega_{i}}\left|\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}\right|^{2} \mathrm{~d} x^{\prime},
\end{aligned}
$$

where for the sake of the last bound we applied the intermediate estimate in (4.6) to the left hand side of (4.5), as discussed in the previous step. Consequently:

$$
\left|R_{k}-R_{s}\right|^{2} \leq C \int_{\omega}\left|\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}\right|^{2} \mathrm{~d} x^{\prime},
$$

and thus:

$$
\begin{aligned}
\int_{\omega_{k}}\left|\nabla y-R_{s} \nabla y_{0}\right|^{2} \mathrm{~d} x^{\prime} & \leq 2\left(\int_{\omega_{k}}\left|\nabla y-R_{k} \nabla y_{0}\right|^{2} \mathrm{~d} x^{\prime}+\int_{\omega_{k}}\left|R_{k}-R_{s}\right|^{2}\left|\nabla y_{0}\right|^{2} \mathrm{~d} x^{\prime}\right) \\
& \leq C \int_{\omega}\left|\left((\nabla y)^{\top} \nabla \vec{b}\right)_{\text {sym }}-\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}\right|^{2} \mathrm{~d} x^{\prime} .
\end{aligned}
$$

This shows that one can take one and the same $R=R_{1}$ on each $\left\{\omega_{k}\right\}_{k=1}^{N}$, at the expense of possibly increasing the constant $C$ by a controlled factor depending only on $N$. The proof of (4.4) is done.

Remark 4.3. A similar reasoning as in the proof of Lemma 4.2, yields a quantitative version of the uniqueness of isometric immersion with a prescribed second fundamental form compatible to the metric by the Gauss-Codazzi-Mainardi equations. More precisely, given a smooth metric $g$ in $\omega \subset \mathbb{R}^{2}$, for every two isometric immersions $y_{1}, y_{2} \in W^{2,2}\left(\omega, \mathbb{R}^{3}\right)$ of $g$, there holds:

$$
\min _{R \in S O(3)} \int_{\omega}\left|\nabla y_{1}-R \nabla y_{2}\right|^{2} \mathrm{~d} x^{\prime} \leq C \int_{\omega}\left|\Pi_{y_{1}}-\Pi_{y_{2}}\right|^{2} \mathrm{~d} x^{\prime}
$$

with a constant $C>0$, depending on $g$ and $\omega$ but independent of $y_{1}$ and $y_{2}$.

## 5. Higher order energy scalings

In this and the next sections we assume that:

$$
\begin{equation*}
\lim _{h \rightarrow 0} \frac{1}{h^{2}} \inf \mathcal{E}^{h}=0 \tag{5.1}
\end{equation*}
$$

Recall that by Theorem 3.2 this condition is equivalent to the existence of a (automatically smooth and unique up to rigid motions) vector field $y_{0}: \bar{\omega} \rightarrow \mathbb{R}^{3}$ satisfying:

$$
\begin{equation*}
\left(\nabla y_{0}\right)^{\top} \nabla y_{0}=\overline{\mathcal{G}}_{2 \times 2} \quad \text { and } \quad\left(\left(\nabla y_{0}\right)^{\top} \nabla \vec{b}_{0}\right)_{\text {sym }}=\frac{1}{2}\left(\overline{\mathcal{G}}_{1}\right)_{2 \times 2} \quad \text { on } \omega, \tag{5.2}
\end{equation*}
$$

where in the oscillatory case $(\mathrm{O})$ the symmetric $x^{\prime}$-dependent matrix $\mathcal{G}_{1}$ is given in (EF) and there must be $\left(\mathcal{G}_{1}\right)_{2 \times 2}=x_{3}\left(\overline{\mathcal{G}}_{1}\right)_{2 \times 2}$, whereas in the non-oscillatory (NO) case $\overline{\mathcal{G}}_{1}\left(x^{\prime}\right)$ is simply $\partial_{3} G\left(x^{\prime}, 0\right)$. The (smooth) Cosserat field $\vec{b}_{0}: \bar{\omega} \rightarrow \mathbb{R}^{3}$ in (2.4) is uniquely given by requesting that:

$$
Q_{0} \doteq\left[\partial_{1} y_{0}, \partial_{2} y_{0}, \vec{b}_{0}\right] \quad \text { satisfies: } \quad Q_{0}^{\top} Q_{0}=\overline{\mathcal{G}}, \quad \operatorname{det} Q_{0}>0 \quad \text { on } \omega,
$$

with notation similar to (2.15). We now introduce the new vector field $\vec{d}_{0}: \bar{\Omega} \rightarrow \mathbb{R}^{3}$ through:

$$
\begin{equation*}
Q_{0}^{\top}\left[x_{3} \partial_{1} \vec{b}_{0}\left(x^{\prime}\right), x_{3} \partial_{2} \vec{b}_{0}\left(x^{\prime}\right), \partial_{3} \vec{d}_{0}\left(x^{\prime}, x_{3}\right)\right]-\frac{1}{2} \mathcal{G}_{1}\left(x^{\prime}, x_{3}\right) \in s o(3), \tag{5.3}
\end{equation*}
$$

justified by (5.2) and in agreement with the construction (2.16) of second order terms in the recovery sequence for the Kirchhoff limiting energies. Explicitly, we have:

$$
\vec{d}_{0}\left(x^{\prime}, x_{3}\right)=Q_{0}\left(x^{\prime}\right)^{\top,-1}\left(\int_{0}^{x_{3}} \mathcal{G}_{1}\left(x^{\prime}, t\right) \mathrm{d} t e_{3}-\frac{1}{2} \int_{0}^{x_{3}} \mathcal{G}_{1}\left(x^{\prime}, t\right)_{33} \mathrm{~d} t e_{3}-\frac{x_{3}^{2}}{2}\left[\begin{array}{c}
\left(\nabla \vec{b}_{0}\right)^{\top} \vec{b}_{0}\left(x^{\prime}\right) \\
0
\end{array}\right]\right)
$$

In what follows, the smooth matrix field in (5.3) will be referred to as $P_{0}: \bar{\Omega} \rightarrow \mathbb{R}^{3 \times 3}$, namely:

$$
\begin{equation*}
P_{0}\left(x^{\prime}, x_{3}\right)=\left[x_{3} \partial_{1} \vec{b}_{0}\left(x^{\prime}\right), x_{3} \partial_{2} \vec{b}_{0}\left(x^{\prime}\right), \partial_{3} \vec{d}_{0}\left(x^{\prime}, x_{3}\right)\right] . \tag{5.4}
\end{equation*}
$$

In the non-oscillatory case (NO), the above formulas become:

$$
\begin{align*}
& \vec{d}_{0}=\frac{x_{3}^{2}}{2} \tilde{d}_{0}\left(x^{\prime}\right), \quad P_{0}\left(x^{\prime}, x_{3}\right)=x_{3}\left[\partial_{1} \vec{b}_{0}, \partial_{2} \vec{b}_{0}, \tilde{d}_{0}\right]\left(x^{\prime}\right), \\
& \text { where: } \quad \tilde{d}_{0}\left(x^{\prime}\right)=Q_{0}\left(x^{\prime}\right)^{\mathrm{T},-1}\left(\partial_{3} G\left(x^{\prime}, 0\right) e_{3}-\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{33} e_{3}-\left[\begin{array}{c}
\left(\nabla \vec{b}_{0}\right)^{\top} \vec{b}_{0}\left(x^{\prime}\right) \\
0
\end{array}\right]\right) . \tag{5.5}
\end{align*}
$$

We also note that the assumption $\int_{-1 / 2}^{1 / 2} \mathcal{G}_{1}\left(x^{\prime}, t\right) \mathrm{d} t=0$ implies:

$$
\begin{equation*}
\int_{-1 / 2}^{1 / 2} P_{0}\left(x^{\prime}, x_{3}\right) \mathrm{d} x_{3}=0 \quad \text { for all } x^{\prime} \in \bar{\omega} \tag{5.6}
\end{equation*}
$$

With the aid of $\vec{d}_{0}$ we now construct the sequence of deformations with low energy:
Lemma 5.1. Assume (O). Then (5.1) implies:

$$
\inf \mathcal{E}^{h} \leq C h^{4} .
$$

Proof. Define the sequence of smooth maps $\left\{u^{h}: \bar{\Omega}^{h} \rightarrow \mathbb{R}^{3}\right\}_{h \rightarrow 0}$ by:

$$
\begin{equation*}
u^{h}\left(x^{\prime}, x_{3}\right)=y_{0}\left(x^{\prime}\right)+x_{3} \vec{b}_{0}\left(x^{\prime}\right)+h^{2} \vec{d}_{0}\left(x^{\prime}, \frac{x_{3}}{h}\right) \tag{5.7}
\end{equation*}
$$

In order to compute $\nabla u^{h}\left(A^{h}\right)^{-1}$, recall the expansion of $\left(A^{h}\right)^{-1}$, so that:

$$
\begin{equation*}
\nabla u^{h}(x) A^{h}(x)^{-1}=Q_{0}\left(x^{\prime}\right) \bar{A}\left(x^{\prime}\right)^{-1}\left(I d_{3}+h S^{h}(x)+\mathcal{O}\left(h^{2}\right)\right), \tag{5.8}
\end{equation*}
$$

where for every $x=\left(x^{\prime}, x_{3}\right) \in \Omega^{h}$ :

$$
S^{h}(x)=\bar{A}\left(x^{\prime}\right)^{-1}\left(Q_{0}\left(x^{\prime}\right)^{\top} P_{0}\left(x^{\prime}, \frac{x_{3}}{h}\right)-\bar{A}\left(x^{\prime}\right) A_{1}\left(x^{\prime}, \frac{x_{3}}{h}\right)\right) \bar{A}\left(x^{\prime}\right)^{-1} .
$$

By frame invariance of the energy density $W$ and since $Q_{0}\left(x^{\prime}\right) \bar{A}\left(x^{\prime}\right)^{-1} \in S O(3)$, we obtain:

$$
\begin{aligned}
W\left(\nabla u^{h}(x) A^{h}(x)^{-1}\right) & =W\left(I d_{3}+h S^{h}(x)+\mathcal{O}\left(h^{2}\right)\right)=W\left(I d_{3}+h S^{h}(x)_{\mathrm{sym}}+\mathcal{O}\left(h^{2}\right)\right) \\
& =W\left(I d_{3}+\mathcal{O}\left(h^{2}\right)\right)=\mathcal{O}\left(h^{4}\right),
\end{aligned}
$$

where we also used the fact that $S^{h}(x)_{\text {sym }}=0$ following directly from the definition (5.3). This implies that $\mathcal{E}^{h}\left(u^{h}\right)=\mathcal{O}\left(h^{4}\right)$ as well, proving the claim.

Lemma 5.2. Assume (O) and (5.1). For an open, Lipschitz subset $\mathcal{V} \subset \omega$, denote:

$$
\mathcal{V}^{h}=\mathcal{V} \times\left(-\frac{h}{2}, \frac{h}{2}\right), \quad \mathcal{E}^{h}\left(u^{h}, \mathcal{V}^{h}\right)=\frac{1}{h} \int_{\mathcal{V}^{h}} W\left(\nabla u^{h}\left(A^{h}\right)^{-1}\right) \mathrm{d} x .
$$

If $y_{0}$ is injective on $\mathcal{V}$, then for every $u^{h} \in W^{1,2}\left(\mathcal{V}^{h}, \mathbb{R}^{3}\right)$ there exists $\bar{R}^{h} \in S O(3)$ such that:

$$
\begin{equation*}
\frac{1}{h} \int_{\mathcal{V}^{h}}\left|\nabla u^{h}(x)-\bar{R}^{h}\left(Q_{0}\left(x^{\prime}\right)+h P_{0}\left(x^{\prime}, \frac{x_{3}}{h}\right)\right)\right|^{2} \mathrm{~d} x \leq C\left(\mathcal{E}^{h}\left(u^{h}, \mathcal{V}^{h}\right)+h^{3}\left|\mathcal{V}^{h}\right|\right) \tag{5.9}
\end{equation*}
$$

with the smooth correction matrix field $P_{0}$ in (5.4). The constant $C$ in (5.9) is uniform for all subdomains $\mathcal{V}^{h} \subset \Omega^{h}$ which are bi-Lipschitz equivalent with controlled Lipschitz constants.

Proof. The proof, similar to [40, Lemma 2.2], is a combination of the change of variable argument in Lemma 4.2 and the low energy deformation construction in Lemma 5.1. Observe first that:

$$
Q_{0}\left(x^{\prime}\right)+h P_{0}\left(x^{\prime}, \frac{x_{3}}{h}\right)=\nabla Y^{h}\left(x^{\prime}, x_{3}\right)+\mathcal{O}\left(h^{2}\right),
$$

where by $Y^{h}: \bar{\Omega}^{h} \rightarrow \mathbb{R}^{3}$ we denote the smooth vector fields in (5.7). It is clear that for sufficiently small $h>0$, each $Y_{\mid \mathcal{V}^{h}}^{h}$ is a smooth diffeomorphism onto its image $\mathcal{U}^{h} \subset \mathbb{R}^{3}$, satisfying uniformly: $\operatorname{det} \nabla Y^{h}>c>0$. We now consider $v^{h}=u^{h} \circ\left(Y^{h}\right)^{-1} \in W^{1,2}\left(\mathcal{U}^{h}, \mathbb{R}^{3}\right)$. By the rigidity estimate [17]:

$$
\begin{equation*}
\int_{\mathcal{U}^{h}}\left|\nabla v^{h}-\bar{R}^{h}\right|^{2} \mathrm{~d} z \leq C \int_{\mathcal{U}^{h}} \operatorname{dist}^{2}\left(\nabla v^{h}, S O(3)\right) \mathrm{d} z, \tag{5.10}
\end{equation*}
$$

for some rotation $\bar{R}^{h} \in S O(3)$. Noting that: $\left(\nabla v^{h}\right) \circ Y^{h}=\left(\nabla u^{h}\right)\left(\nabla Y^{h}\right)^{-1}$ in the set $\mathcal{V}^{h}$, the change of variable formula yields for the left hand side in (5.10):

$$
\begin{aligned}
\int_{\mathcal{U}^{h}}\left|\nabla v^{h}-\bar{R}^{h}\right|^{2} \mathrm{~d} z & =\int_{\mathcal{V}^{h}}\left(\operatorname{det} \nabla Y^{h}\right)\left|\left(\nabla u^{h}\right)\left(\nabla Y^{h}\right)^{-1}-\bar{R}^{h}\right|^{2} \mathrm{~d} x \\
& \geq c \int_{\mathcal{V}^{h}}\left|\nabla u^{h}-\bar{R}^{h}\left(Q_{0}\left(x^{\prime}\right)+h P_{0}\left(x^{\prime}, \frac{x_{3}}{h}\right)+\mathcal{O}\left(h^{2}\right)\right)\right|^{2} \mathrm{~d} x \\
& \geq c \int_{\mathcal{V}^{h}}\left|\nabla u^{h}-\bar{R}^{h}\left(Q_{0}\left(x^{\prime}\right)+h P_{0}\left(x^{\prime}, \frac{x_{3}}{h}\right)\right)\right|^{2} \mathrm{~d} x-c \int_{\mathcal{V}^{h}} \mathcal{O}\left(h^{4}\right) \mathrm{d} x .
\end{aligned}
$$

Similarly, the right hand side in (5.10) can be estimated by:

$$
\begin{aligned}
\int_{\mathcal{U}^{h}} \operatorname{dist}^{2}\left(\nabla v^{h}, S O(3)\right) \mathrm{d} z & =\int_{\mathcal{V}^{h}}\left(\operatorname{det} \nabla Y^{h}\right) \operatorname{dist}^{2}\left(\left(\nabla u^{h}\right)\left(\nabla Y^{h}\right)^{-1}, S O(3)\right) \mathrm{d} x \\
& \leq C \int_{\mathcal{V}^{h}} \operatorname{dist}^{2}\left(\left(\nabla u^{h}\right)\left(A^{h}\right)^{-1} \cdot A^{h}\left(\nabla Y^{h}\right)^{-1}, S O(3)\right) \mathrm{d} x \\
& \leq C \int_{\mathcal{V}^{h}} \operatorname{dist}^{2}\left(\left(\nabla u^{h}\right)\left(A^{h}\right)^{-1}, S O(3)\left(\nabla Y^{h}\right)\left(A^{h}\right)^{-1}\right) \mathrm{d} x .
\end{aligned}
$$

Recall that from (5.8) we have: $\left(\nabla Y^{h}\right)\left(A^{h}\right)^{-1} \in S O(3)\left(I d_{3}+h S^{h}+\mathcal{O}\left(h^{2}\right)\right) \subset S O(3)\left(I d_{3}+\mathcal{O}\left(h^{2}\right)\right)$, since $S^{h} \in s o(3)$. Consequently, the above bound becomes:

$$
\begin{aligned}
\int_{\mathcal{U}^{h}} \operatorname{dist}^{2}\left(\nabla v^{h}, S O(3)\right) \mathrm{d} z & \leq C \int_{\mathcal{V}^{h}} \operatorname{dist}^{2}\left(\left(\nabla u^{h}\right)\left(A^{h}\right)^{-1}, S O(3)\left(I d_{3}+\mathcal{O}\left(h^{2}\right)\right)\right) \mathrm{d} x \\
& \leq C \int_{\mathcal{V}^{h}} \operatorname{dist}^{2}\left(\left(\nabla u^{h}\right)\left(A^{h}\right)^{-1}, S O(3)\right)+\mathcal{O}\left(h^{4}\right) \mathrm{d} x
\end{aligned}
$$

The estimate (5.9) follows now in view of (5.10) and by the lower bound on energy density $W$.

The well-known approximation technique [17] combined with the arguments in [40, Corollary 2.3], yield the following approximation result that can be seen as a higher order counterpart of (2.10):

Corollary 5.3. Assume (O) and (5.1). Then, for any sequence $\left\{u^{h} \in W^{1,2}\left(\Omega^{h}, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0}$ satisfying: $\mathcal{E}^{h}\left(u^{h}\right) \leq C h^{4}$, there exists a sequence of rotation-valued maps $R^{h} \in W^{1,2}(\omega, S O(3))$, such that with $P_{0}$ defined in (5.4) we have:

$$
\begin{align*}
& \frac{1}{h} \int_{\Omega^{h}}\left|\nabla u^{h}(x)-R^{h}\left(x^{\prime}\right)\left(Q_{0}\left(x^{\prime}\right)+h P_{0}\left(x^{\prime}, \frac{x_{3}}{h}\right)\right)\right|^{2} \mathrm{~d} x \leq C h^{4} \\
& \int_{\omega}\left|\nabla R^{h}\left(x^{\prime}\right)\right|^{2} \mathrm{~d} x^{\prime} \leq C h^{2} \tag{5.11}
\end{align*}
$$

## 6. Compactness and $\Gamma$-Limit under $C h^{4}$ energy bound

In this section, we derive the $\Gamma$-convergence result for the energy functionals $\mathcal{E}^{h}$ in the von Kármán scaling regime. The general form of the limiting energy $\mathcal{I}_{4}^{O}$ will be further discussed and split into the stretching, bending, curvature and excess components in section 7 . We begin by stating the compactness result, that is the higher order version of Theorem 2.1.

Theorem 6.1. Assume (O) and (5.1). Fix $y_{0}$ solving (5.2) and normalize it to have: $\int_{\omega} y_{0} \mathrm{~d} x^{\prime}=0$. Then, for any sequence of deformations $\left\{u^{h} \in W^{1,2}\left(\Omega^{h}, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0}$ satisfying:

$$
\begin{equation*}
\mathcal{E}^{h}\left(u^{h}\right) \leq C h^{4}, \tag{6.1}
\end{equation*}
$$

there exists a sequence $\left\{\bar{R}^{h} \in S O(3)\right\}_{h \rightarrow 0}$ such that the following convergences (up to a subsequence) below, hold for $y^{h} \in W^{1,2}\left(\Omega, \mathbb{R}^{3}\right)$ :

$$
y^{h}\left(x^{\prime}, x_{3}\right)=\left(\bar{R}^{h}\right)^{\top}\left(u^{h}\left(x^{\prime}, h x_{3}\right)-f_{\Omega^{h}} u^{h} \mathrm{~d} x\right)
$$

(i) $y^{h} \rightarrow y_{0}$ strongly in $W^{1,2}\left(\Omega, \mathbb{R}^{3}\right)$ and $\frac{1}{h} \partial_{3} y^{h} \rightarrow \vec{b}_{0}$ strongly in $L^{2}\left(\Omega, \mathbb{R}^{3}\right)$, as $h \rightarrow 0$.
(ii) There exists $V \in W^{2,2}\left(\omega, \mathbb{R}^{3}\right)$ and $\mathbb{S} \in L^{2}\left(\omega, \mathbb{R}_{\text {sym }}^{2 \times 2}\right)$ such that, as $h \rightarrow 0$ :

$$
\begin{aligned}
& V^{h}\left(x^{\prime}\right)=\frac{1}{h} \int_{-1 / 2}^{1 / 2} y^{h}\left(x^{\prime}, x_{3}\right)-\left(y_{0}\left(x^{\prime}\right)+h x_{3} \vec{b}_{0}\left(x^{\prime}\right)\right) \mathrm{d} x_{3} \rightarrow V \quad \text { strongly in } W^{1,2}\left(\omega, \mathbb{R}^{3}\right) \\
& \frac{1}{h}\left(\left(\nabla y_{0}\right)^{\top} \nabla V^{h}\right)_{\text {sym }} \rightharpoonup \mathbb{S} \quad \text { weakly in } L^{2}\left(\omega, \mathbb{R}^{2 \times 2}\right) .
\end{aligned}
$$

(iii) The limiting displacement $V$ satisfies: $\left(\left(\nabla y_{0}\right)^{\top} \nabla V\right)_{\text {sym }}=0$ in $\omega$.

We omit the proof because it follows as in [40, Theorem 3.1] in view of condition (5.6). We only recall the definitions used in the sequel. The rotations $\bar{R}^{h}$ are given by:

$$
\bar{R}^{h}=\mathbb{P}_{S O(3)} f_{\Omega^{h}} \nabla u^{h}(x) Q_{0}\left(x^{\prime}\right)^{-1} \mathrm{~d} x
$$

and (5.11) implies that they satisfy, for some limiting rotation $\bar{R}$ :

$$
\begin{equation*}
\int_{\omega}\left|R^{h}\left(x^{\prime}\right)-\bar{R}^{h}\right|^{2} \mathrm{~d} x^{\prime} \leq C h^{2} \quad \text { and } \quad \bar{R}^{h} \rightarrow \bar{R} \in S O(3) \tag{6.2}
\end{equation*}
$$

Consequently:

$$
\begin{equation*}
S^{h}=\frac{1}{h}\left(\left(\bar{R}^{h}\right)^{\top} R^{h}\left(x^{\prime}\right)-I d_{3}\right) \rightharpoonup S \quad \text { weakly in } W^{1,2}\left(\omega, \mathbb{R}^{3 \times 3}\right) \tag{6.3}
\end{equation*}
$$

The field $S \in W^{1,2}(\omega, s o(3))$ is such that $\left(\nabla y_{0}\right)^{\top} \nabla V=\left(Q_{0}^{\top} S Q_{0}\right)_{2 \times 2} \in s o(2)$, which allows for defining a new vector field $\vec{p} \in W^{1,2}\left(\omega, \mathbb{R}^{3}\right)$ through:
(6.4) $[\nabla V, \vec{p}]=S Q_{0} \quad$ or equivalently: $\vec{p}\left(x^{\prime}\right)=-Q_{0}\left(x^{\prime}\right)^{\top,-1}\left[\begin{array}{c}\nabla V\left(x^{\prime}\right)^{\top} \vec{b}_{0}\left(x^{\prime}\right) \\ 0\end{array}\right] \quad$ for all $x^{\prime} \in \omega$.

Finally, by (5.11) we note the uniform boundedness of the fields $\left\{Z^{h} \in L^{2}\left(\Omega, \mathbb{R}^{3 \times 3}\right)\right\}_{h \rightarrow 0}$ below, together with their convergence (up to as subsequence) as $h \rightarrow 0$ :

$$
\begin{equation*}
Z^{h}(x)=\frac{1}{h^{2}}\left(\nabla u^{h}\left(x^{\prime}, h x_{3}\right)-R^{h}\left(x^{\prime}\right)\left(Q_{0}\left(x^{\prime}\right)+h P_{0}\left(x^{\prime}, x_{3}\right)\right)\right) \rightharpoonup Z \quad \text { weakly in } L^{2}\left(\Omega, \mathbb{R}^{3 \times 3}\right) \tag{6.5}
\end{equation*}
$$

Rearranging terms and using the previously established convergences, it can be shown that:

$$
\begin{equation*}
\mathbb{S}\left(x^{\prime}\right)=\left(Q_{0}^{\top}\left(x^{\prime}\right) \bar{R}^{\top} \int_{-1 / 2}^{1 / 2} Z\left(x^{\prime}, x_{3}\right) \mathrm{d} x_{3}\right)_{2 \times 2, \mathrm{sym}}-\frac{1}{2} \nabla V\left(x^{\prime}\right)^{\top} \nabla V\left(x^{\prime}\right) \quad \text { for all } x^{\prime} \in \omega \tag{6.6}
\end{equation*}
$$

Theorem 6.2. In the setting of Theorem 6.1, $\liminf _{h \rightarrow 0} \frac{1}{h^{4}} \mathcal{E}^{h}\left(u^{h}\right)$ is bounded below by:

$$
\mathcal{I}_{4}^{O}(V, \mathbb{S})=\frac{1}{2} \int_{\Omega} \mathcal{Q}_{2}\left(x^{\prime}, I\left(x^{\prime}\right)+x_{3} I I I\left(x^{\prime}\right)+I I(x)\right) \mathrm{d} x=\frac{1}{2}\left\|I+x_{3} I I I+I I\right\|_{\mathcal{Q}_{2}}^{2},
$$

where:

$$
\begin{align*}
I\left(x^{\prime}\right) & =\mathbb{S}\left(x^{\prime}\right)+\frac{1}{2} \nabla V\left(x^{\prime}\right)^{\top} \nabla V\left(x^{\prime}\right)-\nabla y_{0}\left(x^{\prime}\right)^{\top} \nabla \int_{-1 / 2}^{1 / 2} \vec{d}_{0}(x) \mathrm{d} x_{3}, \\
I I I\left(x^{\prime}\right) & =\nabla y_{0}\left(x^{\prime}\right)^{\top} \nabla \vec{p}\left(x^{\prime}\right)+\nabla V\left(x^{\prime}\right)^{\top} \nabla \vec{b}_{0},  \tag{6.7}\\
I I(x) & =\frac{x_{3}^{2}}{2} \nabla \vec{b}_{0}\left(x^{\prime}\right)^{\top} \nabla \vec{b}_{0}\left(x^{\prime}\right)+\nabla y_{0}\left(x^{\prime}\right)^{\top} \nabla_{\tan } \vec{d}_{0}(x)-\frac{1}{4} \mathcal{G}_{2}(x)_{2 \times 2} .
\end{align*}
$$

Proof. 1. Towards estimating the energy $\mathcal{E}^{h}\left(u^{h}\right)$, we replace the argument $\nabla u^{h}(x) A^{h}(x)^{-1}$ of the frame invariant density $W$ by:

$$
\begin{align*}
&\left(Q_{0} \bar{A}^{-1}\right)\left(x^{\prime}\right)^{\top} R^{h}\left(x^{\prime}\right)^{\top} \nabla u^{h}(x) A^{h}(x)^{-1} \\
&=\left(Q_{0} \bar{A}^{-1}\right)\left(x^{\prime}\right)^{\top} Q_{0}\left(x^{\prime}\right) A^{h}(x)^{-1}+h \bar{A}\left(x^{\prime}\right)^{-1} Q_{0}\left(x^{\prime}\right)^{\top} P_{0}\left(x^{\prime}, \frac{x_{3}}{h}\right) A^{h}(x)^{-1}  \tag{6.8}\\
&+h^{2} I_{3}^{h}\left(x^{\prime}, \frac{x_{3}}{h}\right) \quad \text { for all } x \in \Omega^{h},
\end{align*}
$$

where $I_{3}^{h}$ is given in (6.11). Calculating the higher order expansion of (2.12):

$$
\begin{align*}
A^{h}(x)^{-1}=\bar{A}\left(x^{\prime}\right)^{-1}+\bar{A}\left(x^{\prime}\right)^{-1}\left(-h A_{1}\left(x^{\prime}, \frac{x_{3}}{h}\right)\right. & +h^{2} A_{1}\left(x^{\prime}, \frac{x_{3}}{h}\right) \bar{A}\left(x^{\prime}\right)^{-1} A_{1}\left(x^{\prime}, \frac{x_{3}}{h}\right) \\
& \left.-\frac{h^{2}}{2} A_{2}\left(x^{\prime}, \frac{x_{3}}{h}\right)\right) \bar{A}\left(x^{\prime}\right)^{-1}+o\left(h^{2}\right), \tag{6.9}
\end{align*}
$$

the expressions in (6.8) can be written as:

$$
\begin{align*}
& \left(Q_{0} \bar{A}^{-1}\right)\left(x^{\prime}\right)^{\top} R^{h}\left(x^{\prime}\right)^{\top} \nabla u^{h}\left(x^{\prime}, h x_{3}\right) A^{h}\left(x^{\prime}, h x_{3}\right)^{-1} \\
& \quad=I d_{3}+h I_{1}\left(x^{\prime}, x_{3}\right)+h^{2}\left(I_{2}\left(x^{\prime}, x_{3}\right)+I_{3}^{h}\left(x^{\prime}, x_{3}\right)\right)+o\left(h^{2}\right) \quad \text { for all } x \in \Omega, \tag{6.10}
\end{align*}
$$

where $I_{1}: \Omega \rightarrow s o(3)$ and $I_{2}: \Omega \rightarrow \mathbb{R}^{3 \times 3}$ are smooth matrix fields, given by:

$$
\begin{aligned}
& I_{1}(x)=\bar{A}\left(x^{\prime}\right)^{-1}\left(Q_{0}\left(x^{\prime}\right)^{\top} P_{0}(x)-\bar{A}\left(x^{\prime}\right) A_{1}(x)\right) \bar{A}\left(x^{\prime}\right)^{-1} \\
& I_{2}(x)=\bar{A}\left(x^{\prime}\right)^{-1}\left(\bar{A}\left(x^{\prime}\right) A_{1}(x) \bar{A}\left(x^{\prime}\right)^{-1} A_{1}(x)-\frac{1}{2} \bar{A}\left(x^{\prime}\right) A_{2}(x)-Q_{0}\left(x^{\prime}\right)^{\top} P_{0}(x) \bar{A}\left(x^{\prime}\right)^{-1} A_{1}(x)\right) \bar{A}\left(x^{\prime}\right)^{-1}
\end{aligned}
$$

The fact that $I_{1}(x) \in s o(3)$ follows from (5.3). Also, we have:

$$
\begin{align*}
& I_{3}^{h}(x)=\bar{A}\left(x^{\prime}\right)^{-1} Q_{0}\left(x^{\prime}\right)^{\top} R^{h}\left(x^{\prime}\right)^{\top} Z^{h}(x) A^{h}\left(x^{\prime}, h x_{3}\right)^{-1}  \tag{6.11}\\
& \quad \rightharpoonup I_{3}(x)=\bar{A}\left(x^{\prime}\right)^{-1} Q_{0}\left(x^{\prime}\right)^{\top} \bar{R}^{\top} Z(x) \bar{A}(x)^{-1} \quad \text { weakly in } L^{2}\left(\Omega, \mathbb{R}^{3 \times 3}\right)
\end{align*}
$$

where we used (6.8) and (6.2) to pass to the limit with $\left(R^{h}\right)^{\top}$. As in the proof of Theorem 2.3, we now identify the "good" sets $\left\{\left|I_{3}^{h}\right|^{2} \leq 1 / h\right\} \subset \Omega$ and employ (6.10) to write there the following Taylor's expansion of $W\left(\nabla u^{h}\left(A^{h}\right)^{-1}\right)$ :

$$
\begin{align*}
W\left(\nabla u ^ { h } \left(x^{\prime},\right.\right. & \left.\left.h x_{3}\right) A^{h}\left(x^{\prime}, h x_{3}\right)^{-1}\right)=W\left(I d_{3}+h I_{1}(x)+h^{2}\left(I_{2}(x)+I_{3}^{h}(x)\right)+o\left(h^{2}\right)\right) \\
& =W\left(e^{-h I_{1}(x)}\left(I d_{3}+h I_{1}(x)+h^{2}\left(I_{2}(x)+I_{3}^{h}(x)\right)\right)+o\left(h^{2}\right)\right) \\
& =W\left(I d_{3}+h^{2}\left(I_{2}-\frac{1}{2} I_{1}^{2}+I_{3}^{h}\right)+o\left(h^{2}\right)\right)  \tag{6.12}\\
& =\frac{h^{4}}{2} \mathcal{Q}_{3}\left(\left(I_{2}-\frac{1}{2} I_{1}^{2}+I_{3}^{h}\right)_{\mathrm{sym}}\right)+o\left(h^{4}\right)
\end{align*}
$$

Above, we repeatedly used the frame invariance of $W$ and the exponential formula:

$$
e^{-h I_{1}}=I d_{3}-h I_{1}+\frac{h^{2}}{2} I_{1}^{2}+\mathcal{O}\left(h^{3}\right)
$$

Since the weak convergence in (6.11) implies convergence of measures $\left|\left|I_{3}^{h}\right|^{2} \geq 1 / h\right| \rightarrow 0$ as $h \rightarrow 0$, with the help of (6.12) we finally arrive at:

$$
\begin{align*}
& \liminf _{h \rightarrow 0} \frac{1}{h^{4}} \frac{1}{h} \int_{\Omega^{h}} W\left(\nabla u^{h}(x) A^{h}(x)^{-1}\right) \mathrm{d} x \geq \liminf _{h \rightarrow 0} \frac{1}{2} \int_{\left|I_{3}^{h}\right|^{2} \leq 1 / h} \mathcal{Q}_{3}\left(I_{2}-\frac{1}{2} I_{1}^{2}+I_{3}^{h}\right) \mathrm{d} x  \tag{6.13}\\
& \geq \frac{1}{2} \int_{\Omega} \mathcal{Q}_{3}\left(I_{2}-\frac{1}{2} I_{1}^{2}+I_{3}\right) \mathrm{d} x=\frac{1}{2} \int_{\Omega} \mathcal{Q}_{2}\left(x^{\prime},\left(\bar{A}\left(I_{2}-\frac{1}{2} I_{1}^{2}+I_{3}\right) \bar{A}\right)_{2 \times 2}\right) \mathrm{d} x
\end{align*}
$$

2. We now compute the effective integrand in (6.13). Firstly, by (2.1) a direct calculation yields:

$$
\begin{equation*}
\left(I_{2}(x)-\frac{1}{2} I_{1}(x)^{2}\right)_{\mathrm{sym}}=\left(I_{2}\right)_{\mathrm{sym}}+\frac{1}{2} I_{1}^{\top} I_{1}=\frac{1}{2} \bar{A}\left(x^{\prime}\right)^{-1}\left(P_{0}(x)^{\top} P_{0}(x)-\frac{1}{2} \mathcal{G}_{2}(x)\right) \bar{A}\left(x^{\prime}\right)^{-1} \tag{6.14}
\end{equation*}
$$

Secondly, to address the symmetric part of the limit $I_{3}$ in (6.11), consider functions $f^{s, h}: \Omega \rightarrow \mathbb{R}^{3}$ :

$$
f^{s, h}(x)=f_{0}^{s}\left(h\left(\bar{R}^{h}\right)^{\top} Z^{h}\left(x^{\prime}, x_{3}+t\right)+S^{h}\left(x^{\prime}\right)\left(Q_{0}\left(x^{\prime}\right)+h P_{0}\left(x^{\prime}, x_{3}+t\right)\right)\right) e_{3} \mathrm{~d} t
$$

By (6.3) it easily follows that:

$$
\begin{equation*}
f^{s, h} \rightarrow S \vec{b}_{0}=\vec{p} \quad \text { strongly in } \quad L^{2}\left(\Omega, \mathbb{R}^{3}\right), \quad \text { as } h \rightarrow 0 \tag{6.15}
\end{equation*}
$$

On the other hand, we write an equivalent form of $f^{s, h}$ and compute the tangential derivatives:

$$
\begin{array}{r}
f^{s, h}(x)=\frac{1}{h^{2} s}\left(y^{h}\left(x^{\prime}, x_{3}+s\right)-y^{h}\left(x^{\prime}, x_{3}\right)\right)-\frac{1}{h} \vec{b}_{0}\left(x^{\prime}\right)-\frac{1}{s}\left(\vec{d}_{0}\left(x^{\prime}, x_{3}+s\right)-\vec{d}_{0}\left(x^{\prime}, x_{3}\right)\right), \\
\partial_{i} f^{s, h}(x)=\frac{1}{s}\left(\bar{R}^{h}\right)^{\top}\left(Z^{h}\left(x^{\prime}, x_{3}+s\right)-Z^{h}\left(x^{\prime}, x_{3}\right)\right) e_{i}+S^{h}\left(x^{\prime}\right) \partial_{i} \vec{b}_{0}\left(x^{\prime}\right) \\
\\
-\frac{1}{s}\left(\partial_{i} \vec{d}_{0}\left(x^{\prime}, x_{3}+s\right)-\partial_{i} \vec{d}_{0}\left(x^{\prime}, x_{3}\right)\right)
\end{array}
$$

for $i=1,2$. In view of (6.2) and (6.3) and equating the tangential derivatives $\partial_{1}, \partial_{2}$, result in:

$$
\bar{R}^{\top}\left(Z\left(x^{\prime}, x_{3}\right)-Z\left(x^{\prime}, 0\right)\right) e_{i}=x_{3}\left(\partial_{i} \vec{p}\left(x^{\prime}\right)-S\left(x^{\prime}\right) \partial_{i} \vec{b}_{0}\left(x^{\prime}\right)\right)+\partial_{i} \vec{d}_{0}\left(x^{\prime}, x_{3}\right)-\partial_{i} \vec{d}_{0}\left(x^{\prime}, 0\right) .
$$

Further, by (6.11), (6.4) and since $S \in s o(3)$, it follows that:

$$
\begin{align*}
\left(\bar{A}\left(x^{\prime}\right) I_{3}(x) \bar{A}\left(x^{\prime}\right)\right)_{2 \times 2, \text { sym }}=\left(Q_{0}\left(x^{\prime}\right)^{\top} \bar{R}^{\top} Z(x)\right. & )_{2 \times 2, \text { sym }} \\
=\left(Q_{0}\left(x^{\prime}\right)^{\top} \bar{R}^{\top} Z\left(x^{\prime}, 0\right)\right)_{2 \times 2, \text { sym }} & +x_{3}\left(\nabla y_{0}\left(x^{\prime}\right)^{\top} \nabla \vec{p}\left(x^{\prime}\right)+\nabla V\left(x^{\prime}\right)^{\top} \nabla \vec{b}_{0}\right)_{\mathrm{sym}}  \tag{6.16}\\
& +\left(Q_{0}\left(x^{\prime}\right)^{\top} \nabla \vec{d}_{0}(x)\right)_{2 \times 2, \text { sym }}
\end{align*}
$$

On the other hand, taking the $x_{3}$-average and recalling (6.6), we get:

$$
\begin{equation*}
\left(Q_{0}\left(x^{\prime}\right)^{\top} \bar{R}^{\top} Z\left(x^{\prime}, 0\right)\right)_{2 \times 2, \text { sym }}=\mathbb{S}\left(x^{\prime}\right)+\frac{1}{2} \nabla V\left(x^{\prime}\right)^{\top} \nabla V\left(x^{\prime}\right)-\left(\nabla y_{0}\left(x^{\prime}\right)^{\top} \nabla \int_{-1 / 2}^{1 / 2} \vec{d}_{0}(x) \mathrm{d} x_{3}\right)_{\text {sym }} \tag{6.17}
\end{equation*}
$$

3. We now finish the proof of Theorem 6.2. Combining (6.14), (6.16) and (6.17), we see that:

$$
\left(\bar{A}\left(x^{\prime}\right)\left(I_{2}-\frac{1}{2} I_{1}^{2}+I_{3}\right) \bar{A}\left(x^{\prime}\right)\right)_{2 \times 2, \mathrm{sym}}=\left(I\left(x^{\prime}\right)+x_{3} I I I\left(x^{\prime}\right)+I I(x)\right)_{\text {sym }} \quad \text { on } \Omega,
$$

where $I, I I, I I I$ are as in (6.7). In virtue of (6.13), we obtain:

$$
\liminf _{h \rightarrow 0} \frac{1}{h^{4}} \frac{1}{h} \int_{\Omega^{h}} W\left(\nabla u^{h}(x) A^{h}(x)^{-1}\right) \mathrm{d} x \geq \frac{1}{2} \int_{\Omega} \mathcal{Q}_{2}\left(x^{\prime}, I\left(x^{\prime}\right)+x_{3} I I I\left(x^{\prime}\right)+I I(x)\right) \mathrm{d} x .
$$

This yields the claimed lower bound by $\mathcal{I}_{4}^{O}(V, \mathbb{S})$.
For the upper bound statement, define the linear spaces:

$$
\begin{align*}
& \mathscr{V}=\left\{V \in W^{2,2}\left(\omega, \mathbb{R}^{3}\right) ;\left(\nabla y_{0}\left(x^{\prime}\right)^{\top} \nabla V\left(x^{\prime}\right)\right)_{\mathrm{sym}}=0 \text { for all } x^{\prime} \in \omega\right\},  \tag{6.18}\\
& \mathscr{S}=\operatorname{cl}_{L^{2}\left(\omega, \mathbb{R}^{2 \times 2}\right)}\left\{\left(\left(\nabla y_{0}\right)^{\top} \nabla w\right)_{\mathrm{sym}} ; w \in W^{1,2}\left(\omega, \mathbb{R}^{3}\right)\right\} .
\end{align*}
$$

We see that the limiting quantities $V$ and $\mathbb{S}$ in Theorem 6.2 satisfy: $V \in \mathscr{V}, \mathbb{S} \in \mathscr{S}$. The space $\mathscr{V}$ consists of the first order infinitesimal isometries on the smooth minimizing immersion surface $y_{0}(\omega)$, i.e. those Sobolev-regular displacements $V$ that preserve the metric on $y_{0}(\omega)$ up to first order. The tensor fields $\mathbb{S} \in \mathscr{S}$ are the finite strains on $y_{0}(\omega)$, eventually forcing the stretching term in the von Kármán energy $\mathcal{I}_{4}^{O}$ to be of second order.

Theorem 6.3. Assume that $y_{0}$ solves (5.2). Then, for every $(V, \mathbb{S}) \in \mathscr{V} \times \mathscr{S}$ there exists a sequence $\left\{u^{h} \in W^{1,2}\left(\Omega^{h}, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0}$ such that the rescaled sequence $\left\{y^{h}\left(x^{\prime}, x_{3}\right)=u^{h}\left(x^{\prime}, h x_{3}\right)\right\}_{h \rightarrow 0}$ satisfies (i) and (ii) of Theorem 6.1, together with:

$$
\begin{equation*}
\lim _{h \rightarrow 0} \frac{1}{h^{4}} \mathcal{E}^{h}\left(u^{h}\right)=\mathcal{I}_{4}^{O}(V, \mathbb{S}) \tag{6.19}
\end{equation*}
$$

Proof. 1. Given admissible $V$ and $\mathbb{S}$, we first define the $\varepsilon$-recovery sequence $\left\{u^{h} \in W^{1, \infty}\left(\Omega^{h}, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0}$. The ultimate argument for (6.19) will be obtained via a diagonal argument. We set:

$$
\begin{aligned}
u^{h}\left(x^{\prime}, x_{3}\right)= & y_{0}\left(x^{\prime}\right)+h v^{h}\left(x^{\prime}\right)+h^{2} w^{h}\left(x^{\prime}\right)+x_{3} \vec{b}_{0}\left(x^{\prime}\right)+h^{2} \vec{d}_{0}\left(x^{\prime}, \frac{x_{3}}{h}\right) \\
& +h^{3} \vec{k}_{0}\left(x^{\prime}, \frac{x_{3}}{h}\right)+h x_{3} \vec{p}^{h}\left(x^{\prime}\right)+h^{2} x_{3} \vec{q}^{h}\left(x^{\prime}\right)+h^{3} \vec{r}^{h}\left(x^{\prime}, \frac{x_{3}}{h}\right) \quad \text { for all }\left(x^{\prime}, x_{3}\right) \in \Omega^{h} .
\end{aligned}
$$

The smooth vector fields $\vec{b}_{0}$ and $\vec{d}_{0}$ are as in (5.2), (5.3). We now introduce other terms in the above expansion. The sequence $\left\{w^{h} \in \mathcal{C}^{\infty}\left(\bar{\omega}, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0}$ is such that:

$$
\begin{align*}
& \left(\left(\nabla y_{0}\right)^{\top} \nabla\left(w^{h}+\int_{-1 / 2}^{1 / 2} \vec{d}_{0}(\cdot, t) \mathrm{d} t\right)\right)_{\text {sym }} \rightarrow \mathbb{S} \quad \text { strongly in } L^{2}\left(\omega, \mathbb{R}^{2 \times 2}\right) \quad \text { as } h \rightarrow 0,  \tag{6.20}\\
& \lim _{h \rightarrow 0} \sqrt{h}\left\|w^{h}\right\|_{W^{2, \infty}\left(\omega, \mathbb{R}^{3}\right)}=0
\end{align*}
$$

Existence of such a sequence is guaranteed by the fact that $\mathbb{S} \in \mathscr{S}$, where we "slow down" the approximations $\left\{w^{h}\right\}$ to guarantee the blow-up rate of order less that $h^{-1 / 2}$. Further, for a fixed small $\varepsilon>0$, the truncated sequence $\left\{v^{h} \in W^{2, \infty}\left(\omega, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0}$ is chosen according to the standard construction in [17] (see also references therein), in a way that:

$$
\begin{align*}
& v^{h} \rightarrow V \quad \text { strongly in } W^{2,2}\left(\omega, \mathbb{R}^{3}\right) \quad \text { as } h \rightarrow 0, \\
& h\left\|v^{h}\right\|_{W^{2, \infty}\left(\omega, \mathbb{R}^{3}\right)} \leq \varepsilon \quad \text { and } \quad \lim _{h \rightarrow 0} \frac{1}{h^{2}}\left|\left\{x^{\prime} \in \omega ; v^{h}\left(x^{\prime}\right) \neq V\left(x^{\prime}\right)\right\}\right|=0 . \tag{6.21}
\end{align*}
$$

The vector field $\vec{k}_{0} \in \mathcal{C}^{\infty}\left(\bar{\Omega}, \mathbb{R}^{3}\right)$ and sequences $\left\{\vec{p}^{h}, \vec{q}^{h} \in W^{1, \infty}\left(\omega, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0},\left\{\tilde{r}^{h} \in L^{\infty}\left(\Omega, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0}$ are defined by:

$$
\begin{align*}
Q_{0}^{\top} \vec{p}^{h}= & {\left[\begin{array}{c}
-\left(\nabla v^{h}\right)^{\top} \vec{b}_{0} \\
0
\end{array}\right], } \\
Q_{0}^{\top} \vec{q}^{h}= & c\left(x^{\prime},\left(\left(\nabla y_{0}\right)^{\top} \nabla w^{h}\right)_{\mathrm{sym}}+\frac{1}{2}\left(\nabla v^{h}\right)^{\top} \nabla v^{h}\right)-\left[\begin{array}{c}
\left(\nabla v^{h}\right)^{\top} \vec{p}^{h} \\
\frac{1}{2}\left|\vec{p}^{h}\right|^{2}
\end{array}\right]-\left[\begin{array}{c}
\left(\nabla w^{h}\right)^{\top} \vec{b}_{0} \\
0
\end{array}\right], \\
Q_{0}^{\top} \partial_{3} \vec{k}_{0}= & c\left(x^{\prime},\left(\left(\nabla y_{0}\right)^{\top} \nabla_{\tan } \vec{d}_{0}\right)_{\mathrm{sym}}+\frac{x_{3}^{2}}{2}\left(\nabla \vec{b}_{0}\right)^{\top} \nabla \vec{b}_{0}-\frac{1}{4}\left(\mathcal{G}_{2}\right)_{2 \times 2}\right)  \tag{6.22}\\
& -\left[\begin{array}{c}
x_{3}\left(\nabla \vec{b}_{0}\right)^{\top} \partial_{3} \vec{d}_{0} \\
\frac{1}{2}\left|\partial_{3} \vec{d}_{0}\right|^{2}
\end{array}\right]+\left[\begin{array}{c}
\left(\nabla_{\tan } \vec{d}_{0}\right)^{\top} \vec{b}_{0} \\
0
\end{array}\right]+\frac{1}{2} \mathcal{G}_{2} e_{3}-\frac{1}{4}\left(\mathcal{G}_{2}\right)_{33} e_{3}, \\
Q_{0}^{\top} \tilde{r}^{h}= & x_{3} c\left(x^{\prime},\left(\left(\nabla y_{0}\right)^{\top} \nabla \vec{p}^{h}+\left(\nabla v^{h}\right)^{\top} \nabla \vec{b}_{0}\right)_{\mathrm{sym}}\right)-\left[\begin{array}{c}
\left(\nabla v^{h}\right)^{\top} \partial_{3} \vec{d}_{0} \\
\left\langle\vec{p}^{h}, \partial_{3} \vec{d}_{0}\right\rangle
\end{array}\right] .
\end{align*}
$$

Finally, we choose $\left\{\vec{r}^{h} \in W^{1, \infty}\left(\Omega, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0}$ to satisfy:

$$
\begin{equation*}
\lim _{h \rightarrow 0}\left\|\partial_{3} \vec{r}^{h}-\tilde{r}^{h}\right\|_{L^{2}\left(\Omega, \mathbb{R}^{3}\right)}=0 \quad \text { and } \quad \lim _{h \rightarrow 0} \sqrt{h}\left\|\vec{r}^{h}\right\|_{W^{1, \infty}\left(\Omega, \mathbb{R}^{3}\right)}=0 \tag{6.23}
\end{equation*}
$$

2. Observe that for all $\left(x^{\prime}, x_{3}\right) \in \Omega$ there holds::

$$
\begin{aligned}
\nabla u^{h}\left(x^{\prime}, h x_{3}\right)= & Q_{0}+h\left(\left[\nabla v^{h}, \vec{p}^{h}\right]+P_{0}\right)+h^{2}\left(\left[\nabla w^{h}, \vec{q}^{h}\right]+\left[x_{3} \nabla \vec{p}^{h}, \partial_{3} \vec{r}^{h}\right]+\left(\nabla_{\tan } \vec{d}_{0}, \partial_{3} \vec{k}_{0}\right]\right) \\
& +\mathcal{O}\left(h^{3}\right)\left(\left|\nabla \vec{k}_{0}\right|+\left|\nabla \vec{q}^{h}\right|+\left|\nabla \vec{r}^{h}\right|\right) .
\end{aligned}
$$

Consequently, by (6.9) it follows that:

$$
\left(\left(\nabla u^{h}\right)\left(A^{h}\right)^{-1}\right)\left(x^{\prime}, h x_{3}\right)=Q_{0} \bar{A}^{-1}\left(I d_{3}+h \bar{A}^{-1} J_{1}^{h} \bar{A}^{-1}+h^{2} \bar{A}^{-1} J_{2}^{h} \bar{A}^{-1}+J_{h}^{3}\right),
$$

where:

$$
\begin{aligned}
& J_{1}^{h}=Q_{0}^{\top}\left(\left[\nabla v^{h}, \vec{p}^{h}\right]+P_{0}\right)-\bar{A} A_{1}, \\
& J_{2}^{h}=Q_{0}^{\top}\left(\left[\nabla w^{h}, \vec{q}^{h}\right]+\left[x_{3} \nabla \vec{p}^{h}, \partial_{3} \vec{r}^{h}\right]+\left[\nabla \vec{d}_{0}, \partial_{3} \vec{k}_{0}\right]\right)-J_{1}^{h} \bar{A}^{-1} A_{1}-\frac{1}{2} \bar{A} A_{2},
\end{aligned}
$$

and where $J_{1}^{h}, J_{2}^{h}, J_{3}^{h}$ satisfy the uniform bounds (independent of $\varepsilon$ ):

$$
\begin{aligned}
& \left|J_{1}^{h}\right| \leq C\left(1+\left|\nabla v^{h}\right|\right), \\
& \left|J_{2}^{h}\right| \leq C\left(1+\left|\nabla w^{h}\right|+\left|\nabla v^{h}\right|^{2}+\left|\nabla^{2} v^{h}\right|+\left|\nabla \vec{r}^{h}\right|\right), \\
& \left|J_{3}^{h}\right| \leq C h^{3}\left(1+\left|\nabla w^{h}\right|+\left|\nabla^{2} w^{h}\right|+\left|\nabla v^{h}\right|^{2}+\left|\nabla^{2} v^{h}\right|+\left|\nabla v^{h}\right| \cdot\left|\nabla^{2} v^{h}\right|+\left|\nabla \vec{r}^{h}\right|\right)+o\left(h^{2}\right) .
\end{aligned}
$$

In particular, the distance $\operatorname{dist}\left(\left(\nabla u^{h}\right)\left(A^{h}\right)^{-1}, S O(3)\right) \leq\left|\left(\nabla u^{h}\right)\left(A^{h}\right)^{-1}-Q_{0} \bar{A}^{-1}\right|$ is as small as one wishes, uniformly in $x \in \Omega$, for $h$ sufficiently small. Thus, the argument $\left(\nabla u^{h}\right)\left(A^{h}\right)^{-1}$ of the frame invariant density $W$ in $\mathcal{E}^{h}\left(u^{h}\right)$ may be replaced by its polar decomposition factor:

$$
\begin{aligned}
\left(\left(\nabla u^{h}\left(A^{h}\right)^{-1}\right)^{\top}\left(\nabla u^{h}\left(A^{h}\right)^{-1}\right)\right)^{1 / 2} & =\left(I d_{3}+2 h^{2} \bar{A}^{-1}\left(\left(J_{2}^{h}\right)_{\mathrm{sym}}+\frac{1}{2}\left(J_{1}^{h}\right)^{\top} \bar{A}^{-2} J_{1}^{h}\right) \bar{A}^{-1}+\mathcal{R}^{h}\right)^{1 / 2} \\
& =I d_{3}+h^{2} \bar{A}^{-1}\left(\left(J_{2}^{h}\right)_{\mathrm{sym}}+\frac{1}{2}\left(J_{1}^{h}\right)^{\top} \bar{A}^{-2} J_{1}^{h}\right) \bar{A}^{-1}+\mathcal{R}^{h}
\end{aligned}
$$

where $\mathcal{R}^{h}$ stands for any quantity obeying the following bound:

$$
\begin{aligned}
\mathcal{R}^{h}= & \mathcal{O}(h)\left|\left(J_{1}^{h}\right)_{\text {sym }}\right|+\mathcal{O}\left(h^{3}\right)\left(1+\left|\nabla v^{h}\right|\right)\left(1+\left|\nabla w^{h}\right|+\left|\nabla v^{h}\right|^{2}+\left|\nabla^{2} v^{h}\right|+\left|\nabla \vec{r}^{h}\right|\right) \\
& +\mathcal{O}\left(h^{3}\right)\left|\nabla^{2} w^{h}\right|+o\left(h^{2}\right)
\end{aligned}
$$

In conclusion, Taylor's expansion of $W$ at $I d_{3}$ gives:

$$
\begin{align*}
& \frac{1}{h^{4}} \int_{\Omega} W\left(\left(\nabla u^{h}\right)\left(A^{h}\right)^{-1}\left(x^{\prime}, h x_{3}\right)\right) \mathrm{d} x \\
&= \frac{1}{h^{4}} \int_{\Omega} W\left(\left(\left(\nabla u^{h}\left(A^{h}\right)^{-1}\right)^{\top}\left(\nabla u^{h}\left(A^{h}\right)^{-1}\right)\right)^{1 / 2}\left(x^{\prime}, h x_{3}\right)\right) \mathrm{d} x  \tag{6.24}\\
& \leq \frac{1}{2} \int_{\Omega} \mathcal{Q}_{3}\left(\bar{A}^{-1}\left(\left(J_{2}^{h}\right)_{\text {sym }}+\frac{1}{2}\left(J_{1}^{h}\right)^{\top} \bar{A}^{-2} J_{1}^{h}\right) \bar{A}^{-1}+\frac{1}{h^{2}} \mathcal{R}^{h}\right) \mathrm{d} x \\
&+\mathcal{O}\left(h^{2}\right) \int_{\Omega}\left|J_{2}^{h}\right|^{3}+\left|J_{1}^{h}\right|^{6} \mathrm{~d} x+\frac{\mathcal{O}(1)}{h^{4}} \int_{\Omega}\left|\mathcal{R}^{h}\right|^{3} \mathrm{~d} x
\end{align*}
$$

The residual terms above are estimated as in [40], using (6.20), (6.21), (6.23). We have:

$$
h^{2} \int_{\Omega}\left|J_{2}^{h}\right|^{3}+\left|J_{1}^{h}\right|^{6} \mathrm{~d} x \leq h^{2} \int_{\Omega} 1+\left|\nabla w^{h}\right|^{3}+\left|\nabla v^{h}\right|^{6}+\left|\nabla^{2} v^{h}\right|^{3}+\left|\nabla \vec{r}^{h}\right|^{3} \mathrm{~d} x \leq o(1)
$$

as $h^{2} \int_{\Omega}\left|\nabla v^{h}\right|^{6} \mathrm{~d} x \leq C h^{2}\left\|\nabla v^{h}\right\|_{W^{1,2}}^{6}=o(1)$ and $h^{2} \int_{\Omega}\left|\nabla^{2} v^{h}\right|^{3} \mathrm{~d} x \leq \epsilon h \int_{\Omega}\left|\nabla^{2} v^{h}\right|^{2} \mathrm{~d} x=o(1)$. Further:

$$
\begin{aligned}
\frac{1}{h^{4}} \int_{\Omega}\left|\mathcal{R}^{h}\right|^{2} \mathrm{~d} x \leq & \frac{1}{h^{2}} \int_{\Omega}\left|\left(\left(\nabla y_{0}\right)^{\top} \nabla v^{h}\right)_{\mathrm{sym}}\right|^{2} \mathrm{~d} x \\
& +\mathcal{O}\left(h^{2}\right) \int_{\Omega}\left(1+\left|\nabla v^{h}\right|^{2}\right)\left(1+\left|\nabla w^{h}\right|^{2}+\left|\nabla v^{h}\right|^{4}+\left|\nabla^{2} v^{h}\right|^{2}+\left|\nabla \vec{r}^{h}\right|^{2}\right) \mathrm{d} x \\
& +\mathcal{O}\left(h^{2}\right) \int_{\Omega}\left|\nabla^{2} w^{h}\right|^{2} \mathrm{~d} x+o(1) \\
= & o(1)+\mathcal{O}\left(h^{2}\right) \int_{\Omega}\left|\nabla v^{h}\right| \cdot\left|\nabla^{2} v^{h}\right|^{2} \leq C \varepsilon
\end{aligned}
$$

because the last condition in (6.21) implies:

$$
\begin{align*}
& \frac{1}{h^{2}} \int_{\Omega}\left|\left(\left(\nabla y_{0}\right)^{\top} \nabla v^{h}\right)_{\text {sym }}\right|^{2} \mathrm{~d} x \leq \frac{C}{h^{2}}\left\|\nabla^{2} v^{h}\right\|_{L^{\infty}} \int_{\left\{v^{h} \neq V\right\}} \operatorname{dist}^{2}\left(x^{\prime},\left\{v^{h}=V\right\}\right) \mathrm{d} x^{\prime} \\
& \leq \frac{C \varepsilon^{2}}{h^{4}} \int_{\left\{v^{h} \neq V\right\}} \operatorname{dist}^{2}\left(x^{\prime},\left\{v^{h}=V\right\}\right) \mathrm{d} x^{\prime} \leq C \varepsilon^{2} \frac{1}{h^{2}}\left|\left\{v^{h} \neq V\right\}\right|=o(1) \tag{6.25}
\end{align*}
$$

From the two estimates above it also follows that $\frac{1}{h^{4}} \int_{\Omega}\left|\mathcal{R}^{h}\right|^{3} \mathrm{~d} x=o(1)$. Consequently, (6.24) yields:

$$
\begin{equation*}
\left.\limsup _{h \rightarrow 0} \frac{1}{h^{4}} \mathcal{E}^{h}\left(u^{h}\right) \leq C \varepsilon+\limsup _{h \rightarrow 0} \frac{1}{2} \int_{\Omega} \mathcal{Q}_{3}\left(\bar{A}^{-1}\left(\left(J_{2}^{h}\right)_{\text {sym }}+\frac{1}{2}\left(J_{1}^{h}\right)^{\top}\right) \bar{A}^{-2} J_{1}^{h}\right) \bar{A}^{-1}\right) \mathrm{d} x . \tag{6.26}
\end{equation*}
$$

3. Observe now that:

$$
\begin{aligned}
\left(J_{2}^{h}\right)_{\mathrm{sym}}+\frac{1}{2}\left(J_{1}^{h}\right)^{\top} \bar{A}^{-2} J_{1}^{h}= & -\left(\left(\left(\nabla y_{0}\right)^{\top} \nabla v^{h}\right)_{\mathrm{sym}}^{*} \bar{A}^{-1} A_{1}\right)_{\mathrm{sym}} \\
& +\left(Q_{0}^{\top}\left[\nabla w^{h}, \vec{q}^{h}\right]+Q_{0}^{\top}\left[x_{3} \nabla \vec{p}^{h}, \partial_{3} \vec{r}^{h}\right]+Q_{0}^{\top}\left[\nabla \vec{d}_{0}, \partial_{3} \vec{k}_{0}\right]\right)_{\mathrm{sym}} \\
& +\frac{1}{2}\left[\nabla v^{h}, \vec{p}^{h}\right]^{\top}\left[\nabla v^{h}, \vec{p}^{h}\right]+\left(\left[\nabla v^{h}, \vec{p}^{h}\right]^{\top} P_{0}\right)_{\mathrm{sym}}+\frac{1}{2} P_{0}^{\top} P_{0}-\frac{1}{4} \mathcal{G}_{2}
\end{aligned}
$$

Replacing $\partial_{3} \vec{r}^{h}$ by $\tilde{r}^{h}$ and using (6.22), it follows that:

$$
\begin{aligned}
& \left(\int_{\Omega} \mathcal{Q}_{3}\left(\bar{A}^{-1}\left(\left(J_{2}^{h}\right)_{\text {sym }}+\frac{1}{2}\left(J_{1}^{h}\right)^{\top} \bar{A}^{-2} J_{1}^{h}\right) \bar{A}^{-1}\right) \mathrm{d} x\right)^{1 / 2} \\
& \quad \leq \\
& \quad\left(\int _ { \Omega } \mathcal { Q } _ { 2 } \left(x^{\prime},\left(\left(\nabla y_{0}\right)^{\top} \nabla w^{h}\right)_{\text {sym }}+\frac{1}{2}\left(\nabla v^{h}\right)^{\top} \nabla v^{h}+x_{3}\left(\left(\nabla y_{0}\right)^{\top} \nabla \vec{p}^{h}+\left(\nabla v^{h}\right)^{\top} \nabla \vec{b}_{0}\right)_{\text {sym }}\right.\right. \\
& \left.\left.\quad+\quad+\frac{x_{3}^{2}}{2}\left(\nabla \vec{b}_{0}\right)^{\top} \nabla \vec{b}_{0}-\frac{1}{4}\left(\mathcal{G}_{2}\right)_{2 \times 2}\right) \mathrm{~d} x\right)^{1 / 2} \\
& \quad+\left\|\left(\left(\nabla y_{0}\right)^{\top} \nabla v^{h}\right)_{\text {sym }}\right\|_{L^{2}(\Omega)}+\left\|\partial_{3} \vec{r}^{h}-\tilde{r}^{h}\right\|_{L^{2}(\Omega)} .
\end{aligned}
$$

The second term above converges to 0 by (6.25) and the third term also converges to 0 , by (6.23). On the other hand, the first term can be split into the integral on the set $\left\{v^{h}=V\right\}$, whose limit as $h \rightarrow 0$ is estimated by $\mathcal{I}_{4}^{O}(V, \mathbb{S})$, and the remaining integral that is bounded by:

$$
\begin{aligned}
& C \int_{\left\{v^{h} \neq V\right\} \times\left(-\frac{1}{2}, \frac{1}{2}\right)} 1+\left|\nabla w^{h}\right|^{2}+\left|\nabla v^{h}\right|^{4}+\left|\nabla^{2} v^{h}\right|^{2}+\left|\nabla \vec{r}^{h}\right|^{3} \mathrm{~d} x \\
& \quad \leq C \epsilon^{2} \frac{1}{h^{2}}\left|\left\{v^{h} \neq V\right\}\right|+C \int_{\left\{v^{h} \neq V\right\}}\left|\nabla v^{h}\right|^{4} \mathrm{~d} x^{\prime} \leq o(1)+C\left|\left\{v^{h} \neq V\right\}\right|^{1 / 2}\left\|\nabla v^{h}\right\|_{L^{8}}^{4}=o(1) .
\end{aligned}
$$

In conclusion, (6.26) becomes (with a uniform constant $C$ that does not depend on $\varepsilon$ ):

$$
\limsup _{h \rightarrow 0} \frac{1}{h^{4}} \mathcal{E}^{h}\left(u^{h}\right) \leq C \varepsilon+\mathcal{I}_{4}^{O}(V, \mathbb{S}) .
$$

A diagonal argument applied to the indicated $\varepsilon$-recovery sequence $\left\{u^{h}\right\}_{h \rightarrow 0}$ completes the proof.
Corollary 6.4. The functional $\mathcal{I}_{4}^{O}$ attains its infimum and there holds:

$$
\lim _{h \rightarrow 0} \frac{1}{h^{4}} \inf \mathcal{E}^{h}=\min \mathcal{I}_{4}^{O} .
$$

The infima in the left hand side are taken over $W^{1,2}\left(\Omega, \mathbb{R}^{3}\right)$ deformations $u^{h}$, whereas the minimum in the right hand side is taken over admissible displacement-strain couples $(V, \mathbb{S}) \in \mathscr{V} \times \mathscr{S}$ in (6.18).

## 7. Further discussion of $\mathcal{I}_{4}^{O}$ and reduction to the non-oscillatory case (NO)

In this section, we identify the appropriate components of the integrand in the energy $\mathcal{I}_{4}^{O}$ as: stretching, bending, curvature and the order-4 excess, the latter quantity being the projection of the entire integrand on the orthogonal complement of $\mathbb{E}_{2}$ in $\mathbb{E}$. This superposition is in the same spirit, as the integrand of $\mathcal{I}_{2}^{O}$ in Theorem 2.3 decoupling into bending and the order- 2 excess, defined as the projection on the orthogonal complement of $\mathbb{E}_{1}$. There, the assumed condition $\int_{-1 / 2}^{1 / 2} \mathcal{G}_{1} \mathrm{~d} x_{3}=0$ served as the compatibility criterion, assuring that the 2 -excess being null results in $\mathcal{I}_{4}^{O}$ coinciding with the non-oscillatory limiting energy $\mathcal{I}_{4}$, written for the effective metric $\bar{G}$ in (EF). Below, we likewise derive the parallel version $\mathcal{I}_{4}$ of $\mathcal{I}_{4}^{O}$, corresponding to the non-oscillatory case, and show that the vanishing of the 4 -excess reduces $\mathcal{I}_{4}^{O}$ to $\mathcal{I}_{4}$ (for the effective metric (EF)), under two new further compatibility conditions (7.9) on $\left(\mathcal{G}_{2}\right)_{2 \times 2}$.

The following formulas will be useful in the sequel:
Lemma 7.1. In the non-oscillatory setting (NO), let $y_{0}, \vec{b}_{0}$ be as in (5.2) and $\tilde{d}_{0}$ as in (5.5). Then:

$$
\left[\partial_{i j} y_{0}, \partial_{i} \vec{b}_{0}, \tilde{d}_{0}\right]\left(x^{\prime}\right)=\left[\partial_{1} y_{0}, \partial_{2} \vec{y}_{0}, \vec{b}_{0}\right]\left(x^{\prime}\right) \cdot\left[\begin{array}{ccc}
\Gamma_{i j}^{1} & \Gamma_{i 3}^{1} & \Gamma_{33}^{1}  \tag{7.1}\\
\Gamma_{i j}^{2} & \Gamma_{i 3}^{2} & \Gamma_{33}^{2} \\
\Gamma_{i j}^{3} & \Gamma_{i 3}^{3} & \Gamma_{33}^{3}
\end{array}\right]\left(x^{\prime}, 0\right) \quad \text { for } i, j=1,2,
$$

for all $x^{\prime} \in \omega$. Consequently, for any smooth vector field $\vec{q}: \omega \rightarrow \mathbb{R}^{3}$ there holds:

$$
\left[\nabla y_{0}\left(x^{\prime}\right)^{\top} \nabla\left(Q_{0}\left(x^{\prime}\right)^{\mathrm{\top},-1} \vec{q}\left(x^{\prime}\right)\right)\right]_{i, j=1,2}=\nabla \vec{q}\left(x^{\prime}\right)_{2 \times 2}-\left[\left\langle\vec{q}\left(x^{\prime}\right),\left[\Gamma_{i j}^{1}, \Gamma_{i j}^{2}, \Gamma_{i j}^{3}\right]\left(x^{\prime}, 0\right)\right\rangle\right]_{i, j=1,2} .
$$

Above, $\left\{\Gamma_{i j}^{k}\right\}$ are the Christoffel symbols of the metric $G$ and the expression in the right and side represents the tangential part of the covariant derivative of the $(0,1)$ tensor field $\vec{q}$ with respect to $G$.
Proof. In view of $\left(\left(\nabla y_{0}\right)^{\top} \nabla \vec{b}_{0}\right)_{\text {sym }}=\frac{1}{2} \partial_{3} G\left(x^{\prime}, 0\right)_{2 \times 2}$ in (5.2) and recalling (3.2), we get:

$$
\left\langle\partial_{i j} y_{0}, \vec{b}_{0}\right\rangle=\frac{1}{2}\left(\partial_{i} G_{j 3}+\partial_{j} G_{i 3}-\partial_{3} G_{i j}\right)\left(x^{\prime}, 0\right) \quad \text { for all } i, j=1,2,
$$

which easily results in:

$$
\left\langle\partial_{i} \vec{b}_{0}, \partial_{j} y_{0}\right\rangle=\frac{1}{2}\left(\partial_{i} G_{j 3}-\partial_{j} G_{i 3}+\partial_{3} G_{i j}\right)\left(x^{\prime}, 0\right) \quad \text { and } \quad\left\langle\partial_{i} \vec{b}_{0}, \vec{b}_{0}\right\rangle=\frac{1}{2} \partial_{i} G_{33}\left(x^{\prime}, 0\right)
$$

Thus (3.3) and the above allow for computing the coordinates in the basis $\partial_{1} y_{0}, \partial_{2} y_{0}, \vec{b}_{0}$ as claimed in (7.1); see also [40, Theorem 6.2] for more details. The second formula results from:

$$
\begin{aligned}
\left\langle\partial_{i} y_{0}, \partial_{j}\left(Q_{0}^{\mathrm{T},-1} \vec{q}\right)\right\rangle & =\left\langle\partial_{i} y_{0}, \partial_{j}\left(Q_{0}^{\mathrm{T},-1}\right) \vec{q}\right\rangle+\left\langle\partial_{i} y_{0}, Q_{0}^{\mathrm{T},-1} \partial_{j} \vec{q}\right\rangle \\
& =-\left\langle\partial_{i} y_{0}, Q_{0}^{\mathrm{T},-1} \partial_{j}\left(Q_{0}^{\mathrm{\top}}\right) Q_{0}^{\mathrm{T},-1} \vec{q}\right\rangle+\left\langle Q_{0}^{-1} \partial_{i} y_{0}, \partial_{j} \vec{q}\right\rangle \\
& =-\left\langle Q_{0}^{-1} \partial_{j}\left(Q_{0}\right) e_{i}, \vec{q}\right\rangle+\left\langle e_{i}, \partial_{j} \vec{q}\right\rangle,
\end{aligned}
$$

which together with (7.1) yields the Lemma.
Lemma 7.2. In the non-oscillatory setting (NO), let $y_{0}, \vec{b}_{0}$ be as in (5.2) and $\tilde{d}_{0}$ as in (5.5). Then the metric-related term $I I$ in (6.7) has the form $I I=\frac{x_{3}^{2}}{2} \bar{I}\left(x^{\prime}\right)$ and for all $x^{\prime} \in \omega$ we have:

$$
\bar{I} \overline{\mathrm{~s} y m}=\left(\nabla \vec{b}_{0}\right)^{\top} \nabla \vec{b}_{0}+\left(\left(\nabla y_{0}\right)^{\top} \nabla \tilde{d}_{0}\right)_{\mathrm{sym}}-\frac{1}{2} \partial_{33} G\left(x^{\prime}, 0\right)_{2 \times 2}=\left[\begin{array}{ll}
R_{1313} & R_{1323}  \tag{7.2}\\
R_{1323} & R_{2323}
\end{array}\right]\left(x^{\prime}, 0\right)
$$

Above, $R_{i j k l}$ are the Riemann curvatures of the metric $G$, evaluated at the midplate points $x^{\prime} \in \omega$.
Proof. We argue as in the proof of [40, Theorem 6.2]. Using (5.3) we arrive at:

$$
\begin{align*}
\left(\left(\nabla y_{0}\right)^{\top} \nabla \tilde{d}_{0}\right)_{\mathrm{sym}}= & -\left[\left\langle\partial_{i j} y_{0}, \vec{d}_{0}\right\rangle\right]_{i, j=1,2}+\frac{1}{2} \partial_{33} G\left(x^{\prime}, 0\right)_{2 \times 2} \\
& +\left[R_{i 2 j 3}-G_{n p}\left(\Gamma_{i 3}^{n} \Gamma_{j 3}^{p}-\Gamma_{i j}^{n} \Gamma_{33}^{p}\right)\right]_{i, j=1,2}\left(x^{\prime}, 0\right) \tag{7.3}
\end{align*}
$$

Directly from (7.1) we hence obtain:

$$
\begin{equation*}
\left\langle\partial_{i j} \vec{y}_{0}, \tilde{d}_{0}\right\rangle=G_{n p} \Gamma_{i j}^{n} \Gamma_{33}^{p}, \quad\left\langle\partial_{i} \vec{b}_{0}, \partial_{j} \vec{b}_{0}\right\rangle=G_{n p} \Gamma_{i 3}^{n} \Gamma_{j 3}^{p}, \tag{7.4}
\end{equation*}
$$

which together with (7.3) yields (7.2).
With the use of Lemma 7.2, it is quite straightforward to derive the ultimate form of the energy $\mathcal{I}_{4}^{O}$ in the non-oscillatory setting. In particular, the proof of the following result is a special case of the proof of Theorem 7.5 below.

Theorem 7.3. Assume (NO) and (5.1). The expression (7.6) becomes:

$$
\begin{aligned}
\mathcal{I}_{4}(V, \mathbb{S})= & \frac{1}{2} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime}, \mathbb{S}\left(x^{\prime}\right)+\frac{1}{2} \nabla V\left(x^{\prime}\right)^{\top} \nabla V\left(x^{\prime}\right)+\frac{1}{24} \nabla \vec{b}_{0}\left(x^{\prime}\right)^{T} \nabla \vec{b}_{0}\left(x^{\prime}\right)-\frac{1}{48} \partial_{33} G\left(x^{\prime}, 0\right)_{2 \times 2}\right) \mathrm{d} x^{\prime} \\
& +\frac{1}{24} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime}, \nabla y_{0}\left(x^{\prime}\right)^{\top} \nabla \vec{p}\left(x^{\prime}\right)+\nabla V\left(x^{\prime}\right)^{\top} \nabla \vec{b}_{0}\left(x^{\prime}\right)\right) \mathrm{d} x^{\prime} \\
& +\frac{1}{1440} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime},\left[\begin{array}{ll}
R_{1313} & R_{1323} \\
R_{1323} & R_{2323}
\end{array}\right]\left(x^{\prime}, 0\right)\right) \mathrm{d} x^{\prime},
\end{aligned}
$$

where $R_{i j k l}$ stand for the Riemann curvatures of the metric $G$.
Remark 7.4. In the particular, "flat" case of $G=I d_{3}$ the functional $\mathcal{I}_{4}$ reduces to the classical von Kármán energy below. Indeed, the unique solution to (5.2) is: $y_{0}=i d, \vec{b}_{0}=e_{3}$ and further:

$$
\begin{aligned}
& \mathscr{V}=\left\{V(x)=\left(\alpha x^{\perp}+\vec{\beta}, v(x)\right) ; \alpha \in \mathbb{R}, \vec{\beta} \in \mathbb{R}^{2}, v \in W^{2,2}(\omega)\right\}, \\
& \mathscr{S}=\left\{\operatorname{sym} \nabla w ; w \in W^{1,2}\left(\omega, \mathbb{R}^{2}\right)\right\} .
\end{aligned}
$$

Given $V \in \mathscr{V}$, we have $\vec{p}=(-\nabla v, 0)$ and thus:

$$
\mathcal{I}_{4}(V, \mathbb{S})=\frac{1}{2} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime}, \operatorname{sym} \nabla w+\frac{1}{2}\left(\alpha^{2} I d_{2}+\nabla v \otimes \nabla v\right)\right) \mathrm{d} x^{\prime}+\frac{1}{24} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime}, \nabla^{2} v\right) \mathrm{d} x^{\prime}
$$

Absorbing the stretching $\alpha^{2} I d_{2}$ into sym $\nabla w$, the above energy can be expressed in a familiar form:

$$
\begin{equation*}
\mathcal{I}_{4}(v, w)=\frac{1}{2} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime}, \operatorname{sym} \nabla w+\frac{1}{2} \nabla v \otimes \nabla v\right) \mathrm{d} x^{\prime}+\frac{1}{24} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime}, \nabla^{2} v\right) \mathrm{d} x^{\prime} \tag{7.5}
\end{equation*}
$$

as a function of the out-of-plane scalar displacement $v$ and the in-plane vector displacement $w$.

As done for the Kirchhoff energy $\mathcal{I}_{2}^{O}$ in Theorem 2.3, we now identify conditions allowing $\mathcal{I}_{4}^{O}$ to coincide with $\mathcal{I}_{4}$ of the effective metric $\bar{G}$, modulo the introduced below order- 4 excess term.

Theorem 7.5. In the setting of Theorem 6.2, we have:

$$
\begin{align*}
\mathcal{I}_{4}^{O}(V, \mathbb{S})= & \frac{1}{2} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime}, \mathbb{S}+\frac{1}{2}(\nabla V)^{\top} \nabla V+B_{0}\right) \mathrm{d} x^{\prime} \\
& +\frac{1}{24} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime},(\nabla V)^{\top} \nabla \vec{b}_{0}+\left(\nabla y_{0}\right)^{\top} \nabla \vec{p}+12 B_{1}\right) \mathrm{d} x^{\prime}  \tag{7.6}\\
& +\frac{1}{1440} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime},\left(\nabla \vec{b}_{0}\right)^{\top} \nabla \vec{b}_{0}+\left(\nabla y_{0}\right)^{\top} \nabla \tilde{d}_{0}-\frac{1}{2}\left(\overline{\mathcal{G}}_{2}\right)_{2 \times 2}\right) \mathrm{d} x^{\prime} \\
& +\frac{1}{2} \operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(I I_{\text {sym }}, \mathbb{E}_{2}\right),
\end{align*}
$$

where $\overline{\mathcal{G}}_{1}$ and $\overline{\mathcal{G}}_{2}$ are given in (EF), inducing $\tilde{d}_{0}$ via (5.5) for $\partial_{3} G=\overline{\mathcal{G}}_{1}$, and where we introduce the following purely metric-related quantities:

$$
\begin{align*}
& \operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(I I_{\text {sym }}, \mathbb{E}_{2}\right)=\operatorname{dist}^{2}( \int_{0}^{x_{3}} \nabla\left(\mathcal{G}_{1} e_{3}\right)_{2 \times 2, \text { sym }} \mathrm{d} t \\
&-\left[\left\langle\int_{0}^{x_{3}} \mathcal{G}_{1} e_{3} \mathrm{~d} t,\left[\Gamma_{i j}^{1}, \Gamma_{i j}^{2}, \Gamma_{i j}^{3}\right]\left(x^{\prime}, 0\right)\right\rangle\right]_{i, j=1,2}  \tag{7.7}\\
&\left.+\frac{1}{2} \int_{0}^{x_{3}}\left(\mathcal{G}_{1}\right)_{33} \mathrm{~d} t\left[\Gamma_{i j}^{3}\left(x^{\prime}, 0\right)\right]_{i, j=1,2}-\frac{1}{4}\left(\mathcal{G}_{2}\right)_{2 \times 2}, \mathbb{E}_{2}\right), \\
& B_{0}=\frac{1}{24}\left(\nabla \vec{b}_{0}\right)^{\top} \nabla \vec{b}_{0}-\frac{1}{4} \int_{-1 / 2}^{1 / 2}\left(\mathcal{G}_{2}\right)_{2 \times 2} \mathrm{~d} x_{3}=\frac{1}{24}\left[\overline{\mathcal{G}}_{n p} \Gamma_{i 3}^{n} \Gamma_{j 3}^{p}\right]_{i, j=1,2}-\frac{1}{4} \int_{-1 / 2}^{1 / 2}\left(\mathcal{G}_{2}\right)_{2 \times 2} \mathrm{~d} x_{3}, \\
& B_{1}=\left(\nabla y_{0}\right)^{\top} \nabla\left(\int_{-1 / 2}^{1 / 2} x_{3} \vec{d}_{0} \mathrm{~d} x_{3}\right)-\frac{1}{4} \int_{-1 / 2}^{1 / 2} x_{3}\left(\mathcal{G}_{2}\right)_{2 \times 2} \mathrm{~d} x_{3} \\
&=-\nabla\left(\int_{-1 / 2}^{1 / 2} \frac{x_{3}^{2}}{2} \mathcal{G}_{1} e_{3} \mathrm{~d} x_{3}\right)_{2 \times 2}+\left[\left\langle\int_{-1 / 2}^{1 / 2} \frac{x_{3}^{2}}{2} \mathcal{G}_{1} e_{3} \mathrm{~d} x_{3},\left[\Gamma_{i j}^{1}, \Gamma_{i j}^{2}, \Gamma_{i j}^{3}\right]\left(x^{\prime}, 0\right)\right\rangle\right]_{i, j=1,2}  \tag{7.8}\\
& \quad-\frac{1}{2} \int_{-1 / 2}^{1 / 2} \frac{x_{3}^{2}}{2}\left(\mathcal{G}_{1}\right)_{33} \mathrm{~d} x_{3}\left[\Gamma_{i j}^{3}\left(x^{\prime}, 0\right)\right]_{i, j=1,2}-\frac{1}{4} \int_{-1 / 2}^{1 / 2} x_{3}\left(\mathcal{G}_{2}\right)_{2 \times 2} \mathrm{~d} x_{3},
\end{align*}
$$

By $\left\{\Gamma_{i j}^{k}\right\}$ we denote the Christoffel symbols of the metric $\bar{G}$ in (EF). The third term in (7.6) equals the scaled norm of the Riemann curvatures of the effective metric $\bar{G}$ :

$$
\frac{1}{1440} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime},\left[\begin{array}{ll}
R_{1313} & R_{1323} \\
R_{1323} & R_{2323}
\end{array}\right]\left(x^{\prime}, 0\right)\right) \mathrm{d} x^{\prime} .
$$

The first three terms in $\mathcal{I}_{4}^{O}$ coincide with $\mathcal{I}_{4}$ in Theorem 7.3 for the effective metric $\bar{G}$ in (EF), provided that the following compatibility conditions hold:

$$
\begin{align*}
& \int_{-1 / 2}^{1 / 2}\left(15 x_{3}^{2}-\frac{9}{4}\right) \mathcal{G}_{2}\left(x^{\prime}, x_{3}\right)_{2 \times 2} \mathrm{~d} x_{3}=0 \\
& \begin{aligned}
& \frac{1}{4} \int_{-1 / 2}^{1 / 2} x_{3} \mathcal{G}_{2}\left(x^{\prime}, x_{3}\right)_{2 \times 2} \mathrm{~d} x_{3}+\nabla\left(\int_{-1 / 2}^{1 / 2} \frac{x_{3}^{2}}{2} \mathcal{G}_{1} e_{3} \mathrm{~d} x_{3}\right)_{2 \times 2, \text { sym }} \\
&- {\left[\left\langle\int_{-1 / 2}^{1 / 2} \frac{x_{3}^{2}}{2} \mathcal{G}_{1} e_{3} \mathrm{~d} x_{3},\left[\Gamma_{i j}^{1}, \Gamma_{i j}^{2}, \Gamma_{i j}^{3}\right]\left(x^{\prime}, 0\right)\right\rangle\right]_{i, j=1,2} } \\
&+\frac{1}{2} \int_{-1 / 2}^{1 / 2} \frac{x_{3}^{2}}{2}\left(\mathcal{G}_{1}\right)_{33} \mathrm{~d} x_{3}\left[\Gamma_{i j}^{3}\left(x^{\prime}, 0\right)\right]_{i, j=1,2}=0
\end{aligned} \tag{7.9}
\end{align*}
$$

Proof. We write:

$$
\mathcal{I}_{4}^{O}(V, \mathbb{S})=\frac{1}{2}\left\|I+x_{3} I I I+I I\right\|_{\mathcal{Q}_{2}}^{2}=\frac{1}{2}\left\|I+x_{3} I I I+\mathbb{P}_{2}(I I)\right\|_{\mathcal{Q}_{2}}^{2}+\frac{1}{2} \operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(I I_{\mathrm{sym}}, \mathbb{E}_{2}\right)
$$

and further decompose the first term above along the Legendre projections:

$$
\begin{aligned}
& \| I+x_{3} I I I+\mathbb{P}_{2}(I I)\left\|_{\mathcal{Q}_{2}}^{2}=\right\| \int_{-1 / 2}^{1 / 2}\left(I+x_{3} I I I+I I\right) p_{0}\left(x_{3}\right) \mathrm{d} x_{3} \|_{\mathcal{Q}_{2}}^{2} \\
&+\left\|\int_{-1 / 2}^{1 / 2}\left(I+x_{3} I I I+I I\right) p_{1}\left(x_{3}\right) \mathrm{d} x_{3}\right\|_{\mathcal{Q}_{2}}^{2}+\left\|\int_{-1 / 2}^{1 / 2}\left(I+x_{3} I I I+I I\right) p_{2}\left(x_{3}\right) \mathrm{d} x_{3}\right\|_{\mathcal{Q}_{2}}^{2} \\
&=\underbrace{\left\|I+\int_{-1 / 2}^{1 / 2} I I \mathrm{~d} x_{3}\right\|_{\mathcal{Q}_{2}}^{2}}_{\text {Stretching }}+\frac{1}{12} \underbrace{\left\|I I I+12 \int_{-1 / 2}^{1 / 2} x_{3} I I \mathrm{~d} x_{3}\right\|_{\mathcal{Q}_{2}}^{2}}_{\text {Bending }}+\underbrace{\left\|\int_{-1 / 2}^{1 / 2} p_{2}\left(x_{3}\right) I I \mathrm{~d} x_{3}\right\|_{\mathcal{Q}_{2}}^{2}}_{\text {Curvature }} .
\end{aligned}
$$

To identify the four indicated terms in $\mathcal{I}_{4}^{O}$, observe that $\int_{-1 / 2}^{1 / 2} x_{3} \int_{0}^{x_{3}} \mathcal{G}_{1} \mathrm{~d} x_{3}=-\int_{-1 / 2}^{1 / 2} \frac{x_{3}^{2}}{2} \mathcal{G}_{1} \mathrm{~d} x_{3}$ and:

$$
\begin{aligned}
& \operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(I I_{\text {sym }}, \mathbb{E}_{2}\right) \\
& \quad=\operatorname{dist}_{\mathcal{Q}_{2}}^{2}\left(\left(\left(\nabla y_{0}\right)^{\top} \nabla\left(Q_{0}^{\mathrm{T},-1} \int_{0}^{x_{3}} \mathcal{G}_{1} e_{3} \mathrm{~d} t-\frac{1}{2} Q_{0}^{\mathrm{T},-1} \int_{0}^{x_{3}}\left(\mathcal{G}_{1}\right)_{33} \mathrm{~d} t e_{3}\right)\right)_{\text {sym }}-\frac{1}{4}\left(\mathcal{G}_{2}\right)_{2 \times 2}, \mathbb{E}_{2}\right) .
\end{aligned}
$$

Thus the formulas in (7.7) and (7.8) follow directly from Lemma 7.1 and (7.4). There also holds:

$$
\begin{aligned}
& \text { Stretching }=\int_{\omega} \mathcal{Q}_{2}\left(x^{\prime}, \mathbb{S}\right.\left.+\frac{1}{2}(\nabla V)^{\top} \nabla V+\frac{1}{24}\left(\nabla \vec{b}_{0}\right)^{\top} \nabla \vec{b}_{0}-\frac{1}{4} \int_{-1 / 2}^{1 / 2}\left(\mathcal{G}_{2}\right)_{2 \times 2} \mathrm{~d} x_{3}\right) \mathrm{d} x^{\prime}, \\
& \text { Bending }=\frac{1}{12} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime},(\nabla V)^{\top} \nabla \vec{b}_{0}+\left(\nabla y_{0}\right)^{\top} \nabla \vec{p}\right. \\
&\left.+12\left(\nabla y_{0}\right)^{\top} \nabla\left(\int_{-1 / 2}^{1 / 2} x_{3} \vec{d}_{0} \mathrm{~d} x_{3}\right)-3 \int_{-1 / 2}^{1 / 2} x_{3}\left(\mathcal{G}_{2}\right)_{2 \times 2} \mathrm{~d} x_{3}\right), \mathrm{d} x^{\prime}
\end{aligned}
$$

$$
\text { Curvature }=\frac{1}{720} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime},\left(\nabla \vec{b}_{0}\right)^{\top} \nabla \vec{b}_{0}+60\left(\nabla y_{0}\right)^{\top} \nabla\left(\int_{-1 / 2}^{1 / 2}\left(6 x_{3}^{2}-\frac{1}{2}\right) \vec{d}_{0} \mathrm{~d} x_{3}\right)\right.
$$

$$
\left.-15 \int_{-1 / 2}^{1 / 2}\left(6 x_{3}^{2}-\frac{1}{2}\right)\left(\mathcal{G}_{2}\right)_{2 \times 2} \mathrm{~d} x_{3}\right) \mathrm{d} x^{\prime}
$$

It is easy to check that with the choice of the effective metric components $\overline{\mathcal{G}}_{1} e_{3}$ and $\left(\overline{\mathcal{G}}_{2}\right)_{2 \times 2}$ and denoting $\tilde{d}_{0}$ the corresponding vector in (5.5), we have:

$$
\text { Curvature }=\frac{1}{720} \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime},\left(\nabla \vec{b}_{0}\right)^{\top} \nabla \vec{b}_{0}+\left(\nabla y_{0}\right)^{\top} \nabla \tilde{d}_{0}-\frac{1}{2}\left(\overline{\mathcal{G}}_{2}\right)_{2 \times 2}\right) \mathrm{d} x^{\prime} .
$$

This proves (7.6). Equivalence of the constraints (7.9) with:

$$
\int_{-1 / 2}^{1 / 2}\left(\mathcal{G}_{2}\right)_{2 \times 2} \mathrm{~d} x_{3}=\frac{1}{12}\left(\overline{\mathcal{G}}_{2}\right)_{2 \times 2} \quad \text { and } \quad\left(B_{1}\right)_{\text {sym }}=0 \quad \text { in } \omega
$$

follows by a direct inspection. We now invoke Lemma 7.2 to complete the proof.

Remark 7.6. Observe that the vanishing of the 4 -excess and curvature terms in $\mathcal{I}_{4}^{O}$ :

$$
I I_{\text {sym }} \in \mathbb{E}_{2} \quad \text { and } \quad \text { Curvature }=0,
$$

are the necessary conditions for $\min \mathcal{I}_{4}^{O}=0$ and they are equivalent to $I I_{\text {sym }} \in \mathbb{E}_{1}$. Consider now a particular case scenario of $\overline{\mathcal{G}}=I d_{3}$ and $\mathcal{G}_{1}=0$, where the spaces $\mathscr{V}$ and $\mathscr{S}$ are given in Remark 7.4, together with $\overrightarrow{d_{0}}=0$. Then, the above necessary condition reduces to: $\left(\mathcal{G}_{2}\right)_{2 \times 2} \in \mathbb{E}_{1}$, namely:

$$
\left(\mathcal{G}_{2}\right)_{2 \times 2}\left(x^{\prime}, x_{3}\right)=x_{3} \mathcal{F}_{1}\left(x^{\prime}\right)+\mathcal{F}_{0}\left(x^{\prime}\right) \quad \text { for all } x=\left(x^{\prime}, x_{3}\right) \in \bar{\Omega}
$$

It is straightforward that, on a simply connected midplate $\omega$, both terms:
Stretching $=\int_{\omega} \mathcal{Q}_{2}\left(x^{\prime}, \operatorname{sym} \nabla w+\frac{1}{2} \nabla v \otimes \nabla v-\frac{1}{4} \mathcal{F}_{0}\right) \mathrm{d} x^{\prime}, \quad$ Bending $=\int_{\omega} \mathcal{Q}_{2}\left(x^{\prime}, \nabla^{2} v+\frac{1}{4} \mathcal{F}_{1}\right) \mathrm{d} x^{\prime}$, can be equated to 0 by choosing appropriate displacements $v$ and $w$, if and only if there holds:

$$
\begin{equation*}
\operatorname{curl} \mathcal{F}_{1}=0, \quad \operatorname{curl}^{\top} \operatorname{curl} \mathcal{F}_{0}+\frac{1}{4} \operatorname{det} \mathcal{F}_{1}=0 \quad \text { in } \omega \tag{7.10}
\end{equation*}
$$

Note that these are precisely the linearised Gauss-Codazzi-Mainardi equations corresponding to the metric $I d_{2}+2 h^{2} \mathcal{F}_{0}$ and shape operator $\frac{1}{2} h \mathcal{F}_{1}$ on $\omega$. We see that these conditions are automatically satisfied in presence of (7.9), when $\left(\mathcal{G}_{2}\right)_{2 \times 2} \in \mathbb{E}_{1}$ actually results in $\left(\mathcal{G}_{2}\right)_{2 \times 2}=0$. An integrability criterion similar to (7.10) can be derived also in the general case, under $I I_{\text {sym }} \in \mathbb{E}_{1}$ and again it automatically holds with (7.9). This last statement will be pursued in the next section.
8. Identification of the $C h^{4}$ scaling regime and coercivity of the limiting energy $\mathcal{I}_{4}$

Theorem 8.1. The energy scaling beyond the von Kármán regime:

$$
\lim _{h \rightarrow 0} \frac{1}{h^{4}} \inf \mathcal{E}^{h}=0
$$

is equivalent to the following condition, on a simply connected $\omega$ :
(i) in the oscillatory case ( $O$ ), in presence of the compatibility conditions (7.9)
$\left[I I_{\text {sym }} \in \mathbb{E}_{2}\right.$ and (8.2) holds with $G$ replaced by the effective metric $\bar{G}$ in (EF). This condition involves $\overline{\mathcal{G}}, \overline{\mathcal{G}}_{1}$ and $\left(\overline{\mathcal{G}}_{2}\right)_{2 \times 2}$ terms of $\bar{G}$.
(ii) in the non-oscillatory case (NO)

$$
\left[\begin{array}{l}
\text { All the Riemann curvatures of the metric } G \text { vanish on } \omega \times\{0\} \text { : }  \tag{8.2}\\
\qquad R_{i j k l}\left(x^{\prime}, 0\right)=0 \quad \text { for all } x^{\prime} \in \omega \quad \text { and all } i, j, k, l=1 \ldots 3 .
\end{array}\right.
$$

Proof. By Corollary 6.4, it suffices to determine the equivalent conditions for $\min \mathcal{I}_{4}=0$. Clearly, $\min \mathcal{I}_{4}=0$ implies (8.2). Vice versa, if (8.2) holds, then:

$$
\frac{1}{24}\left(\nabla \vec{b}_{0}\right)^{\top} \nabla \vec{b}_{0}-\frac{1}{48} \partial_{33} G\left(x^{\prime}, 0\right)=-\frac{1}{24}\left(\left(\nabla y_{0}\right)^{\top} \nabla \tilde{d}_{0}\right)_{\text {sym }}
$$

by Lemma 7.2. Taking $V=\vec{p}=0$ and $\mathbb{S}=\frac{1}{24}\left(\left(\nabla y_{0}\right)^{\top} \nabla \tilde{d}_{0}\right)_{\text {sym }} \in \mathscr{S}$, we get $\mathcal{I}_{4}(V, \mathbb{S})=0$.
We further have the following counterpart of the essential uniqueness of the minimizing isometric immersion $y_{0}$ statement in Theorem 3.2:

Theorem 8.2. In the non-oscillatory setting (NO), assume (8.2). Then $\mathcal{I}_{4}(V, \mathbb{S})=0$ if and only if:

$$
\begin{equation*}
V=S y_{0}+c \quad \text { and } \quad \mathbb{S}=\frac{1}{2}\left(\left(\nabla y_{0}\right)^{\top} \nabla\left(S^{2} y_{0}+\frac{1}{12} \tilde{d}_{0}\right)\right)_{\text {sym }} \quad \text { on } \omega, \tag{8.3}
\end{equation*}
$$

for some skew-symmetric matrix $S \in \operatorname{so(3)}$ and a vector $c \in \mathbb{R}^{3}$.
Proof. We first observe that the bending term $I I I$ in (6.7) is already symmetric, because:

$$
\begin{aligned}
{\left[\left\langle\partial_{i} y_{0}, \partial_{j} \vec{p}\right\rangle+\left\langle\partial_{i} V, \partial_{j} \vec{b}_{0}\right\rangle\right]_{i, j=1,2} } & =\left[\partial_{j}\left(\left\langle\partial_{i} y_{0}, \vec{p}\right\rangle+\left\langle\partial_{i} V, \vec{b}_{0}\right\rangle\right)\right]_{i, j=1,2}-\left[\left\langle\partial_{i j} y_{0}, \vec{p}\right\rangle+\left\langle\partial_{i j} V, \vec{b}_{0}\right\rangle\right]_{i, j=1,2} \\
& =-\left[\left\langle\partial_{i j} y_{0}, \vec{p}\right\rangle+\left\langle\partial_{i j} V, \vec{b}_{0}\right\rangle\right]_{i, j=1,2} \in \mathbb{R}_{\mathrm{sym}}^{2 \times 2}
\end{aligned}
$$

where we used the definition of $\vec{p}$ in (6.4). Recalling (7.2), we see that $\mathcal{I}_{4}(V, \mathbb{S})=0$ if and only if:

$$
\begin{align*}
& \mathbb{S}+\frac{1}{2}(\nabla V)^{\top} \nabla V-\frac{1}{24}\left(\left(\nabla y_{0}\right)^{\top} \nabla \tilde{d}_{0}\right)_{\text {sym }}=0,  \tag{8.4}\\
& \left(\nabla y_{0}\right)^{\top} \nabla \vec{p}+(\nabla V)^{\top} \nabla \vec{b}_{0}=0 .
\end{align*}
$$

Consider the matrix field $S=[\nabla V, \vec{p}] Q_{0}^{-1} \in W^{1,2}(\omega, s o(3))$ as in (6.4). Note that:

$$
\begin{align*}
\partial_{i} S= & {\left[\nabla \partial_{i} V, \partial_{i} \vec{p}\right] Q_{0}^{-1}-[\nabla V, \vec{p}] Q_{0}^{-1}\left(\partial_{i} Q_{0}\right) Q_{0}^{-1}=Q_{0}^{-1, \top} \bar{S}^{i} Q_{0}^{-1} \quad \text { for } i=1,2 } \\
& \text { where } \quad \bar{S}^{i}=Q_{0}^{\top}\left[\nabla \partial_{i} V, \partial_{i} \vec{p}\right]+[\nabla V, \vec{p}]^{\top}\left(\partial_{i} Q_{0}\right) \in L^{2}(\omega, s o(3)) . \tag{8.5}
\end{align*}
$$

Then we have:

$$
\left\langle\bar{S}^{i} e_{1}, e_{2}\right\rangle=\partial_{i}\left(\left\langle\partial_{2} y_{0}, \partial_{i} V\right\rangle+\left\langle\partial_{2} V, \partial_{i} y_{0}\right\rangle\right)-\left(\left\langle\partial_{12} y_{0}, \partial_{i} V\right\rangle+\left\langle\partial_{12} V, \partial_{i} y_{0}\right\rangle\right)=0
$$

because the first term in the right hand side above equals 0 in view of $V \in \mathscr{V}$, whereas the second term equals $\partial_{2}\left\langle\partial_{1} y_{0}, \partial_{1} V\right\rangle$ for $i=1$ and $\partial_{1}\left\langle\partial_{2} y_{0}, \partial_{2} V\right\rangle$ for $i=2$, both expression being null again in view of $V \in \mathscr{V}$. We now claim that $\left\{\bar{S}^{i}\right\}_{i=1,2}=0$ is actually equivalent to the second condition in (8.4). It suffices to examine the only possibly nonzero components:

$$
\begin{equation*}
\left\langle\bar{S}^{i} e_{3}, e_{j}\right\rangle=\left\langle\partial_{j} y_{0}, \partial_{i} \vec{p}\right\rangle+\left\langle\partial_{j} V, \partial_{i} \vec{b}_{0}\right\rangle=\left(\left(\nabla y_{0}\right)^{\top} \nabla \vec{p}+(\nabla V)^{\top} \nabla \vec{b}_{0}\right)_{i j} \quad \text { for all } i, j=1,2, \tag{8.6}
\end{equation*}
$$

proving the claim.
Consequently, the second condition in (8.4) is equivalent to $S$ being constant, to the effect that $\nabla V=\nabla\left(S y_{0}\right)$, or equivalently that $V-S y_{0}$ is a constant vector. In this case:

$$
\mathbb{S}=\frac{1}{2}\left(\nabla y_{0}\right)^{\top} \nabla\left(S^{2} y_{0}\right)+\frac{1}{24}\left(\left(\nabla y_{0}\right)^{\top} \nabla \tilde{d}_{0}\right)_{\mathrm{sym}}=\frac{1}{2}\left(\left(\nabla y_{0}\right)^{\top} \nabla\left(S^{2} y_{0}+\frac{1}{12} \tilde{d}_{0}\right)\right)_{\mathrm{sym}}
$$

is equivalent to the first condition in (8.4), as $(\nabla V)^{\top} \nabla V=-\left(\nabla y_{0}\right)^{\top} S^{2} \nabla y_{0}$. The proof is done.

From Theorem 8.2 we deduce its quantitative version, that is a counterpart of Theorem 4.1 in the present von Kármán regime:

Theorem 8.3. In the non-oscillatory setting (NO), assume (8.2). Then for all $V \in \mathscr{V}$ there holds:

$$
\begin{equation*}
\operatorname{dist}_{W^{2,2}\left(\omega, \mathbb{R}^{3}\right)}^{2}\left(V,\left\{S y_{0}+c ; S \in s o(3), c \in \mathbb{R}^{3}\right\}\right) \leq C \int_{\omega} \mathcal{Q}_{2}\left(x^{\prime},\left(\nabla y_{0}\right)^{\top} \nabla \vec{p}+(\nabla V)^{\top} \nabla \overrightarrow{b_{0}}\right) \mathrm{d} x^{\prime} \tag{8.7}
\end{equation*}
$$ with a constant $C>0$ that depends on $G, \omega$ and $W$ but it is independent of $V$.

Proof. We argue by contradiction. Since $\mathscr{V}_{\text {lin }} \doteq\left\{S y_{0}+c ; S \in s o(3), c \in \mathbb{R}^{3}\right\}$ is a linear subspace of $\mathscr{V}$ and likewise the expression $I I I$ in (6.7) is linear in $V$, with its kernel equal to $\mathscr{V}_{\text {lin }}$ in virtue of Theorem 8.2, it suffices to take a sequence $\left\{V_{n} \in \mathscr{V}\right\}_{n \rightarrow \infty}$ such that:

$$
\begin{align*}
& \left\|V_{n}\right\|_{W^{2,2}\left(\omega, \mathbb{R}^{3}\right)}=1, \quad V_{n} \perp_{W^{2,2}\left(\omega, \mathbb{R}^{3}\right)} \not \mathscr{V}_{\text {lin }} \quad \text { for all } n, \\
& \text { and: } \quad\left(\nabla y_{0}\right)^{\top} \nabla \vec{p}_{n}+\left(\nabla V_{n}\right)^{\top} \nabla \vec{b}_{0} \rightarrow 0 \quad \text { strongly in } L^{2}\left(\omega, \mathbb{R}^{2 \times 2}\right), \quad \text { as } n \rightarrow \infty . \tag{8.8}
\end{align*}
$$

Passing to a subsequence if necessary and using the definition of $\vec{p}$ in (6.4), it follows that:

$$
\begin{equation*}
V_{n} \rightharpoonup V \quad \text { weakly in } W^{2,2}\left(\omega, \mathbb{R}^{3}\right), \quad \vec{p}_{n} \rightharpoonup \vec{p} \quad \text { weakly in } W^{1,2}\left(\omega, \mathbb{R}^{3}\right) . \tag{8.9}
\end{equation*}
$$

Clearly, $Q_{0}^{\top}[\nabla V, \vec{p}] \in L^{2}(\omega, s o(3))$ so that $V \in \mathscr{V}$, but also $\left(\nabla y_{0}\right)^{\top} \nabla \vec{p}+(\nabla V)^{\top} \nabla \vec{b}_{0}=0$. Thus, Theorem 8.2 and the perpendicularity assumption in (8.8) imply: $V=\vec{p}=0$. We will now prove:

$$
\begin{equation*}
V_{n} \rightarrow 0 \quad \text { strongly in } W^{2,2}\left(\omega, \mathbb{R}^{3}\right), \tag{8.10}
\end{equation*}
$$

which will contradict the first (normalisation) condition in (8.7).
As in (8.5), the assumption $V_{n} \in \mathscr{V}$ implies that for each $x^{\prime} \in \omega$ and $i=1,2$, the following matrix (denoted previously by $\bar{S}^{i}$ ) is skew-symmetric:

$$
\bar{Q}_{0}^{\top}\left[\nabla \partial_{i} V_{n}, \partial \vec{p}_{n}\right]+\left[\nabla V_{n}, \vec{p}_{n}\right]^{\top}\left(\partial_{i} Q_{0}\right) \in s o(3) .
$$

Equating tangential entries and observing (8.8), yields for every $i, j, k=1,2$ :

$$
\left\langle\partial_{j} y_{0}, \partial_{i k} V_{n}\right\rangle+\left\langle\partial_{k} y_{0}, \partial_{i j} V_{n}\right\rangle=-\left(\left\langle\partial_{j} V_{n}, \partial_{i k} y_{0}\right\rangle+\left\langle\partial_{k} V_{n}, \partial_{i j} y_{0}\right\rangle\right) \rightarrow 0 \quad \text { strongly in } L^{2}(\omega) .
$$

Permuting $i, j, k$ we eventually get:

$$
\left\langle\partial_{j} y_{0}, \partial_{i k} V_{n}\right\rangle \rightarrow 0 \quad \text { strongly in } L^{2}(\omega) \quad \text { for all } i, j, k=1,2 .
$$

On the other hand, equating off-tangential entries, we get by (8.8) and (8.9) that for each $i=1,2$ :

$$
\left\langle\vec{b}_{0}, \partial_{i j} V_{n}\right\rangle=-\left(\left(\nabla y_{0}\right)^{\top} \nabla \vec{p}_{n}+\left(\nabla V_{n}\right)^{\top} \nabla \vec{b}_{0}\right)_{i j}-\left\langle\vec{p}_{n}, \partial_{i j} y_{0}\right\rangle \rightarrow 0 \quad \text { strongly in } L^{2}(\omega) .
$$

Consequently, $\left\{Q_{0}^{\top} \partial_{i j} V_{n} \rightarrow 0\right\}_{i, j=1,2}$ in $L^{2}\left(\omega, \mathbb{R}^{3}\right)$, which implies convergence (8.10) as claimed. This ends the proof of (8.7).

Remark 8.4. Although the kernel of the (nonlinear) energy $\mathcal{I}_{4}$, displayed in Theorem 8.2, is finite dimensional, the full coercivity estimate of the form below is false:

$$
\begin{align*}
\min _{S \in s o(3), c \in \mathbb{R}^{3}} & \left(\left\|V-\left(S y_{0}+c\right)\right\|_{W^{2,2}\left(\omega, \mathbb{R}^{3}\right)}^{2}+\left\|\mathbb{S}-\frac{1}{2}\left(\left(\nabla y_{0}\right)^{\top} \nabla\left(S^{2} y_{0}-\frac{1}{12} \tilde{d}_{0}\right)\right)_{\text {sym }}\right\|_{L^{2}\left(\omega, \mathbb{R}^{2 \times 2}\right)}^{2}\right)  \tag{8.11}\\
& \leq C \mathcal{I}_{4}(V, \mathbb{S}) \quad \text { for all }(V, \mathbb{S}) \in \mathscr{V} \times \mathscr{S} .
\end{align*}
$$

For a counterexample, consider the particular case of classical von Kármán functional (7.5), specified in Remark 7.4. Clearly, $\mathcal{I}_{4}(v, w)=0$ if an only if $v(x)=\langle a, x\rangle+\alpha$ and $w(x)=\beta x^{\perp}-\frac{1}{2}\langle a, x\rangle a+\gamma$,
for some $a \in \mathbb{R}^{2}$ and $\alpha, \beta, \gamma \in \mathbb{R}$. Note that (8.7) reflects then the Poincaré inequality: $\int_{\omega} \mid \nabla v-$ $\left.f_{\omega} \nabla v\right|^{2} \mathrm{~d} x^{\prime} \leq C \int_{\omega}\left|\nabla^{2} v\right|^{2} \mathrm{~d} x^{\prime}$, whereas (8.11) takes the form:

$$
\begin{equation*}
\min _{a \in \mathbb{R}^{2}}\left(\int_{\omega}|\nabla v-a|^{2} \mathrm{~d} x^{\prime}+\int_{\omega}\left|\operatorname{sym} \nabla w+\frac{1}{2} a \otimes a\right|^{2} \mathrm{~d} x^{\prime}\right) \leq C \mathcal{I}_{4}(v, w) \tag{8.12}
\end{equation*}
$$

Let $\omega=B_{1}(0)$. Given $v \in W^{2,2}(\omega)$ such that $\operatorname{det} \nabla^{2} v=0$, let $w$ satisfy: $\operatorname{sym} \nabla w=-\frac{1}{2} \nabla v \otimes \nabla v$, which results in vanishing of the first term in (7.5). Neglecting the first term in the left hand side of (8.12), leads in this context to the following weaker form, which we below disprove:

$$
\begin{equation*}
\min _{a \in \mathbb{R}^{2}} \int_{\omega}|\nabla v \otimes \nabla v-a \otimes a|^{2} \mathrm{~d} x^{\prime} \leq C \int_{\omega}\left|\nabla^{2} v\right|^{2} \mathrm{~d} x^{\prime} \tag{8.13}
\end{equation*}
$$

Define $v_{n}(x)=n\left(x_{1}+x_{2}\right)+\frac{1}{2}\left(x_{1}+x_{2}\right)^{2}$ for all $x=\left(x_{1}, x_{2}\right) \in \omega$. Then $\nabla v_{n}=\left(n+x_{1}+x_{2}\right)(1,1)$ and $\operatorname{det} \nabla^{2} v_{n}=0$. Minimization in (8.13) becomes: $\min _{a \in \mathbb{R}^{2}} \int_{\omega}\left|\left(n+x_{1}+x_{2}\right)^{2}(1,1) \otimes(1,1)-a \otimes a\right|^{2} \mathrm{~d} x^{\prime}$ and an easy explicit calculation yields the necessary form of the minimizer: $a=\delta(1,1)$. Thus, the same minimization can be equivalently written and estimated in:

$$
4 \cdot \min _{\delta \in \mathbb{R}} \int_{\omega}\left|\left(n+x_{1}+x_{2}\right)^{2}-\delta^{2}\right|^{2} \mathrm{~d} x^{\prime} \sim 4 n^{2} \rightarrow \infty \quad \text { as } n \rightarrow \infty .
$$

On the other hand, $\left|\nabla^{2} v_{n}\right|^{2}=4$ at each $x^{\prime} \in \omega$. Therefore, the estimate (8.13) cannot hold.

## 9. Beyond the von Kármán regime: an example

Given a function $\phi \in \mathcal{C}^{\infty}\left(\left(-\frac{1}{2}, \frac{1}{2}\right), \mathbb{R}\right)$, consider the conformal metric:

$$
G\left(x^{\prime}, x_{3}\right)=e^{2 \phi\left(x_{3}\right)} I d_{3} \quad \text { for all } x=\left(x^{\prime}, x_{3}\right) \in \Omega^{h} .
$$

The midplate metric $\overline{\mathcal{G}}_{2 \times 2}=e^{2 \phi(0)} I d_{2}$ has a smooth isometric immersion $y_{0}=e^{\phi(0)} i d_{2}: \omega \rightarrow \mathbb{R}^{2}$ and thus by Theorem 2.4 there must be:

$$
\inf \mathcal{E}^{h} \leq C h^{2}
$$

By a computation, we get that the only possibly non-zero Christoffel symbols of $G$ are: $\Gamma_{11}^{3}=\Gamma_{22}^{3}=$ $-\phi^{\prime}\left(x_{3}\right)$ and $\Gamma_{13}^{1}=\Gamma_{23}^{2}=\Gamma_{33}^{3}=\phi^{\prime}\left(x_{3}\right)$, while the only possibly nonzero Riemann curvatures are:

$$
\begin{equation*}
R_{1212}=-\phi^{\prime}\left(x_{3}\right)^{2} e^{2 \phi\left(x_{3}\right)}, \quad R_{1313}=R_{2323}=-\phi^{\prime \prime}\left(x_{3}\right) e^{2 \phi\left(x_{3}\right)} \tag{9.1}
\end{equation*}
$$

Consequently, the results of this paper provide the following hierarchy of possible energy scalings:
(a) $\left\{c h^{2} \leq \inf \mathcal{E}^{h} \leq C h^{2}\right\}_{h \rightarrow 0}$ with $c, C>0$. This scenario is equivalent to $\phi^{\prime}(0) \neq 0$. The functionals $\frac{1}{h^{2}} \mathcal{E}^{h}$ as in Theorems 2.1, 2.3 and 2.4 exhibit the indicated compactness properties and $\Gamma$-converge to the following energy $\mathcal{I}_{2}$ defined on the set of deformations: $\left\{y \in W^{2,2}\left(\omega, \mathbb{R}^{3}\right) ;(\nabla y)^{\top} \nabla y=I d_{2}\right\}:$

$$
\mathcal{I}_{2}(y)=\frac{1}{24} \int_{\omega} \mathcal{Q}_{2}\left(\Pi_{y}-\phi^{\prime}(0) I d_{2}\right) \mathrm{d} x^{\prime}
$$

Here $\mathcal{Q}_{2}\left(F_{2 \times 2}\right)=\min \left\{D^{2} W\left(I d_{3}\right)(\tilde{F}, \tilde{F}) ; \tilde{F} \in \mathbb{R}^{3 \times 3}\right.$ with $\left.\tilde{F}_{2 \times 2}=F_{2 \times 2}\right\}$.
(b) $\left\{c h^{4} \leq \inf \mathcal{E}^{h} \leq C h^{4}\right\}_{h \rightarrow 0}$ with $c, C>0$. This scenario is equivalent to $\phi^{\prime}(0)=0$ and $\phi^{\prime \prime}(0) \neq$ 0 . The unique (up to rigid motions) minimizing isometric immersion is then $i d_{2}: \omega \rightarrow \mathbb{R}^{2}$ and the functionals $\frac{1}{h^{4}} \mathcal{E}^{h}$ have the compactness and $\Gamma$-convergence properties as in Theorems
6.1, 6.2 and 6.3. The following limiting functional $\mathcal{I}_{4}$ is defined on the set of displacements $\left\{(v, w) \in W^{2,2}(\omega, \mathbb{R}) \times W^{1,2}\left(\omega, \mathbb{R}^{2}\right)\right\}$ as in Remark 7.4:

$$
\begin{aligned}
\mathcal{I}_{4}(v, w)= & \frac{1}{2} \int_{\omega} \mathcal{Q}_{2}\left(\operatorname{sym} \nabla w+\frac{1}{2} \nabla v \otimes \nabla v-\frac{1}{24} \phi^{\prime \prime}(0) I d_{2}\right) \mathrm{d} x^{\prime} \\
& +\frac{1}{24} \int_{\omega} \mathcal{Q}_{2}\left(\nabla^{2} v\right) \mathrm{d} x^{\prime}+\frac{1}{1440} \phi^{\prime \prime}(0)^{2}|\omega| \mathcal{Q}_{2}\left(I d_{2}\right) .
\end{aligned}
$$

(c) $\left\{\inf \mathcal{E}^{h} \leq C h^{6}\right\}_{h \rightarrow 0}$ with $C>0$. This scenario is equivalent to $\phi^{\prime}(0)=0$ and $\phi^{\prime \prime}(0)=0$ and in fact we have the following more precise result below.
Theorem 9.1. Let $G\left(x^{\prime}, x_{3}\right)=e^{2 \phi\left(x_{3}\right)} I d_{3}$, where $\phi^{(k)}(0)=0$ for $k=1 \ldots n-1$ up to some $n>2$. Then: $\inf \mathcal{E}^{h} \leq C h^{2 n}$ and:

$$
\begin{equation*}
\lim _{h \rightarrow 0} \frac{1}{h^{2 n}} \inf \mathcal{E}^{h} \geq c_{n} \phi^{(n)}(0)^{2}|\omega| \mathcal{Q}_{2}\left(I d_{2}\right) \tag{9.2}
\end{equation*}
$$

where $c_{n}>0$. In particular, if $\phi^{(n)}(0) \neq 0$ then we have: ch $h^{2 n} \leq \inf \mathcal{E}^{h} \leq C h^{2 n}$ with $c, C>0$.
Proof. 1. For the upper bound, we compute:

$$
\begin{aligned}
\mathcal{E}^{h}\left(e^{\phi(0)} i d_{3}\right) & =\frac{1}{h} \int_{\Omega^{h}} W\left(e^{\phi(0)-\phi\left(x_{3}\right)} I d_{3}\right) \mathrm{d} x=\frac{1}{2 h} \int_{\Omega^{h}} \mathcal{Q}_{3}\left(\phi^{(n)}(0) \frac{x_{3}^{n}}{n!} I d_{3}\right)+\mathcal{O}\left(h^{2 n+2}\right) \mathrm{d} x \\
& =h^{2 n}\left(\frac{\phi^{(n)}(0)^{2}}{(n!)^{2}} \frac{1}{(2 n+1) 2^{2 n+1}}|\omega| \mathcal{Q}_{3}\left(I d_{3}\right)+o(1)\right) \leq C h^{2 n},
\end{aligned}
$$

where we used the fact that $e^{\phi(0)-\phi\left(x_{3}\right)}=1-\phi^{(n)}(0) \frac{x_{3}^{n}}{n!}+\mathcal{O}\left(\left|x_{3}\right|^{n+1}\right)$.
2. To prove the lower bound (9.2), let $\left\{u^{h} \in W^{1,2}\left(\Omega^{h}, \mathbb{R}^{3}\right)\right\}_{h \rightarrow 0}$ satisfy $\mathcal{E}^{h}\left(u^{h}\right) \leq C h^{2 n}$. Then:

$$
\begin{aligned}
\mathcal{E}^{h}\left(u^{h}\right) & \geq \frac{c}{h} \int_{\Omega^{h}} \operatorname{dist}^{2}\left(\nabla u^{h}, e^{\phi\left(x_{3}\right)} S O(3)\right) \mathrm{d} x \\
& \geq \frac{c}{h} \int_{\Omega^{h}} \operatorname{dist}^{2}\left(\nabla u^{h}, e^{\phi(0)} S O(3)\right) \mathrm{d} x-\frac{\bar{c}}{h} \int_{\Omega^{h}}\left|\phi^{(n)}(0) \frac{x_{3}^{n}}{n!}+\mathcal{O}\left(h^{n+1}\right)\right|^{2} \mathrm{~d} x,
\end{aligned}
$$

which results in: $\frac{1}{h} \int_{\Omega^{h}} \operatorname{dist}^{2}\left(e^{-\phi(0)} \nabla u^{h}, S O(3)\right) \mathrm{d} x \leq C h^{2 n}$. Similarly as in Lemma 5.2 and Corollary 5.3 , it follows that there exist approximating rotation fields $\left\{R^{h} \in W^{1,2}(\omega, S O(3))\right\}_{h \rightarrow 0}$ such that:

$$
\begin{equation*}
\frac{1}{h} \int_{\Omega^{h}}\left|\nabla u^{h}-e^{\phi(0)} R^{h}\right|^{2} \mathrm{~d} x \leq C h^{2 n}, \quad \int_{\omega}\left|\nabla R^{h}\right|^{2} \mathrm{~d} x \leq C h^{2 n-2} \tag{9.3}
\end{equation*}
$$

As in sections 2 and 6 , we define the following displacement and deformation fields:

$$
\begin{aligned}
& y^{h}\left(x^{\prime}, x_{3}\right)=\left(\bar{R}^{h}\right)^{\top}\left(u^{h}\left(x^{\prime}, h x_{3}\right)-f_{\Omega^{h}} u^{h}\right) \in W^{1,2}\left(\Omega, \mathbb{R}^{3}\right), \text { where } \bar{R}^{h}=\mathbb{P}_{S O(3)} f_{\Omega^{h}} e^{-\phi(0)} \nabla u^{h}(x) \mathrm{d} x, \\
& V^{h}\left(x^{\prime}\right)=\frac{1}{h^{n-1}} \int_{-1 / 2}^{1 / 2} y^{h}\left(x^{\prime}, x_{3}\right)-e^{\phi(0)}\left(i d_{2}+h x_{3} e_{3}\right) \mathrm{d} x_{3} \in W^{1,2}\left(\omega, \mathbb{R}^{3}\right) .
\end{aligned}
$$

In view of (9.3), we obtain then the following convergences (up to a not relabelled subsequence):

$$
\begin{aligned}
& y^{h} \rightarrow e^{\phi(0)} i d_{2} \quad \text { in } W^{1,2}\left(\omega, \mathbb{R}^{3}\right), \quad \frac{1}{h} \partial_{3} y^{h} \rightarrow e^{\phi(0)} e_{3} \text { in } L^{2}\left(\omega, \mathbb{R}^{3}\right), \\
& V^{h} \rightarrow V \in W^{2,2}\left(\omega, \mathbb{R}^{3}\right) \quad \text { in } W^{1,2}\left(\omega, \mathbb{R}^{3}\right), \quad \frac{1}{h}\left(\nabla V^{h}\right)_{2 \times 2, \text { sym }} \rightharpoonup \operatorname{sym} \nabla w \quad \text { weakly in } L^{2}\left(\omega, \mathbb{R}^{2 \times 2}\right) .
\end{aligned}
$$

This allows to conclude the claimed lower bound:

$$
\begin{aligned}
\liminf _{h \rightarrow 0} \frac{1}{h^{2 n}} \mathcal{E}^{h}\left(u^{h}\right) & \geq \frac{1}{2}\left\|e^{-\phi(0)} \operatorname{sym} \nabla w-x_{3} e^{-\phi(0)} \nabla^{2} V^{3}-\phi^{(n)}(0) \frac{x_{3}^{n}}{n!} I d_{2}\right\|_{\mathcal{Q}_{2}}^{2} \\
& \geq \frac{1}{2}\left\|\phi^{(n)}(0) \frac{x_{3}^{n}}{n!} I d_{2}-\mathbb{P}_{1}\left(\phi^{(n)}(0) \frac{x_{3}^{n}}{n!} I d_{2}\right)\right\|_{\mathcal{Q}_{2}}^{2} \\
& =\frac{1}{2} \frac{\phi^{(n)}(0)^{2}}{(n!)^{2}} \cdot \int_{-1 / 2}^{1 / 2}\left(x_{3}^{n}-\mathbb{P}_{1}\left(x_{3}^{n}\right)\right)^{2} \mathrm{~d} x_{3} \cdot|\omega| \mathcal{Q}_{2}\left(I d_{2}\right) \\
& =c_{n} \cdot \phi^{(n)}(0)^{2}|\omega| \mathcal{Q}_{2}\left(I d_{2}\right),
\end{aligned}
$$

as in (9.2), with the following constant $c_{n}$ :

$$
c_{n}=\frac{1}{2^{2 n+1}(n!)^{2}} \begin{cases}\frac{(n-1)^{2}}{(2 n+1)(n+2)^{2}} & \text { for } n \text { odd } \\ \frac{n^{2}}{(2 n+1)(n+1)^{2}} & \text { for } n \text { even } .\end{cases}
$$

Observe that $c_{2}=\frac{1}{1440}$, consistently with the previous direct application of Theorem 6.2.

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