University of Pittsburgh's Department of Physics and Astronomy PITT PACC Workshop:



Flux Measurement and Determination in the Intensity Frontier Era Neutrino Beams

December 6-8, 2012, Pittsburgh, PA, USA

#### MINERVA Flux:

#### **Current Uncertainties And Future Plans**

Leonidas Aliaga (presented by Vittorio Paolone/UPitt) William & Mary

# Outline

- (Quick) Description of NuMI.
- Our Present Understanding of the NuMI Flux.
- Improving our Flux knowledge
  - Minerva's strategy
- Conclusions and future steps.

#### Hadron production and focusing



#### Total focusing uncertainties...



Focusing uncertainties are expected to be small in comparison with hadron production uncertainties.

Z. Pavlovich, "Observation of disappearance of muon neutrinos in the NuMI beam", PhD thesis, UT Austin 2008





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# But in this talk we're particularly interested in Minerva...

Our main goals are to measure:

- Neutrino-nucleus cross sections of exclusive and inclusive final states.
- The nuclear effects on the v-A interactions and form factors and structure functions.

#### To produce high precision measurements of absolute cross sections...



#### → We need to know our flux

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## Understanding the Flux

- Big discrepancies between predictions from hadronic models.
- We need to go back to the history of every neutrino:
- What happens with the neutrino ancestors and their interactions?



 MINERvA uses a geant4 based MC for Flux simulation and FTFP (FRITIOF Precompound (FTFP) model )

# $\pi^{\scriptscriptstyle +}$ which produce $\nu_{_{\!\!\!\!\mu}}$ passing through MINOS/MINERvA

#### Origin

Parent of  $\pi^+$ 



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#### pC-> $\pi^{+/-}$ X Distributions for LE Configuration



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#### **pC-> K<sup>+/-</sup>X Distributions for LE Configuration**



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#### Predicted Neutrino Flux from target



#### Secondaries That Interact in the Target

Energy Spectrum of Charged Pions and Kaons That Are Neutrino Ancestors

Energy Spectrum of Protons That Are Neutrino Ancestors



#### Improving our Flux Constraint

Multi-prong approach:

- Use the Muon Monitor Data.
- Take Special runs: Vary the beam parameters (horn current, target position).
- Use  $\nu_{_{\mu}}$  atomic electron interactions (see Jaewon's talk)
- Using low v-method (See Arie's talk)
- Using external hadron production data. Results shown here Constraint the MC flux to get the right shape and uncertainty.  $\Phi(E_v) \equiv \Phi(x_F, p_T)$

#### Redundancy will improve our accuracy...

#### Using external Hadron Production



- For  $pC \rightarrow \pi^{+/-} X$ : NA49 @ 158 GeV (CERN), Barton @ 100 GeV (Fermilab) & NA61 @ 31 GeV (CERN) & HARP @ 3, 5, 8, 12 GeV.
- For  $pC \rightarrow K^{+\prime-} X$ : NA49 @ 158 GeV (Tinti's thesis, FERMILAB), MIPP @ 120 GeV ratio  $\pi/K$  (thick: Seun & thin: Lebedev).
- For  $\pi^{+\prime}C \rightarrow \pi^{+\prime}X$ : HARP @ 3, 5, 8, 12 GeV.

#### NA61 & NA49 coverage for charged pions



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pC → pi+ X @158





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#### Comparison pC @ 31 GeV vs pC @ 158 GeV



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#### **Results to Neutrino flux**

Applying corrections to  $\pi^+ C \to \pi^{\pm} X$  of  $E \frac{d^3 \sigma}{dp^3}$ :  $w(x_F, p_T) = \frac{NA49(x_F, p_T, 158GeV)}{FTFP'(x_F, p_T, 158GeV)}$ 





- $U_{\mu}$  the neutrino spectra when we focus  $\pi^+$ .
- ${f V}_{
  m L}$  the anti-neutrino spectra when we focus  $\pi^-$ . 12/07/12 Leonidas Aliaga (presented by V. Paolone)

# Model Spread Uncertainties

- Spread shown between geant4 models for Non-NA49 uncertainties.
- Divide into categories:  $\pi$ , K, p, n and other secondary interactions in target, in horns, decay pipe walls or He, target hall chase.
- A lot of work is required to add more models and gradually replace model spread with existing and new data.



### **Total Uncertainties: Reconstructed E**<sub>v</sub>



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## Hadron Production Test at NA61/SHINE

NA61 (SPS Heavy Ion and Neutrino Experiment) at CERN is a large acceptance hadron spectrometer in the North Area H2 beamline of the SPS.



# Conclusions

For MINERvA it is crucial to have solid flux constraints

with small uncertainties to measure absolute v cross

sections.

- We are employing multiple approaches to constrain the flux.
- Presented here are our first results using external

hadron production data and we expect further

improvement using more measurements (i.e NA61).

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# Backup slides

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#### Components are consumables...

- High radiation, thermal stress, mechanical stress, water leaks affect all components.
- Replaced: 6 times targets, horns and hadron monitors.
- Gradual slope due to radiation damage.



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#### Sources of the focusing uncertainties

- Horn angles. Horn tilt 0.2 mrad.
- Horn offsets. 1.0 mm. Transverse misalignment
- Horn current. 1.0%. Miscalibration.
- Horn current distribution.  $\delta$ =6mm /  $\delta$ =infinitive.
- **Baffle scraping.** 0.25%.
- Protons on target. 2.0 %. Beam intensity, size and position of the beam. MC Beam size: 1.1x1.2 mm2 and target size: 6.4x15mm2.

#### Comparison NA61 and NA49



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#### Comparison NA61 and NA49



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#### QGSP @ different energies



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#### **QGSP FTFP vs Data Comparison**



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#### **QGSP FTFP vs Data Comparison**



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#### **CC** Inclusive Neutrino Reconstructed Energy



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