

# Neutrino Electron Scattering


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**Flux Measurement and Determination in the Intensity Frontier Era Neutrino Beams**

Dec 6-8, 2012 Pittsburgh, PA

# Overview

- Motivation
  - Flux constraining using  $\nu_{\mu} + e$  scattering
- Signal/Background separation
- MINERvA detector
- Single EM shower reconstruction
- e/ $\gamma$  Separation using dE/dx
- Small sample  comparison
- Estimated statistics in ME
- Summary

# NuMI Beamline

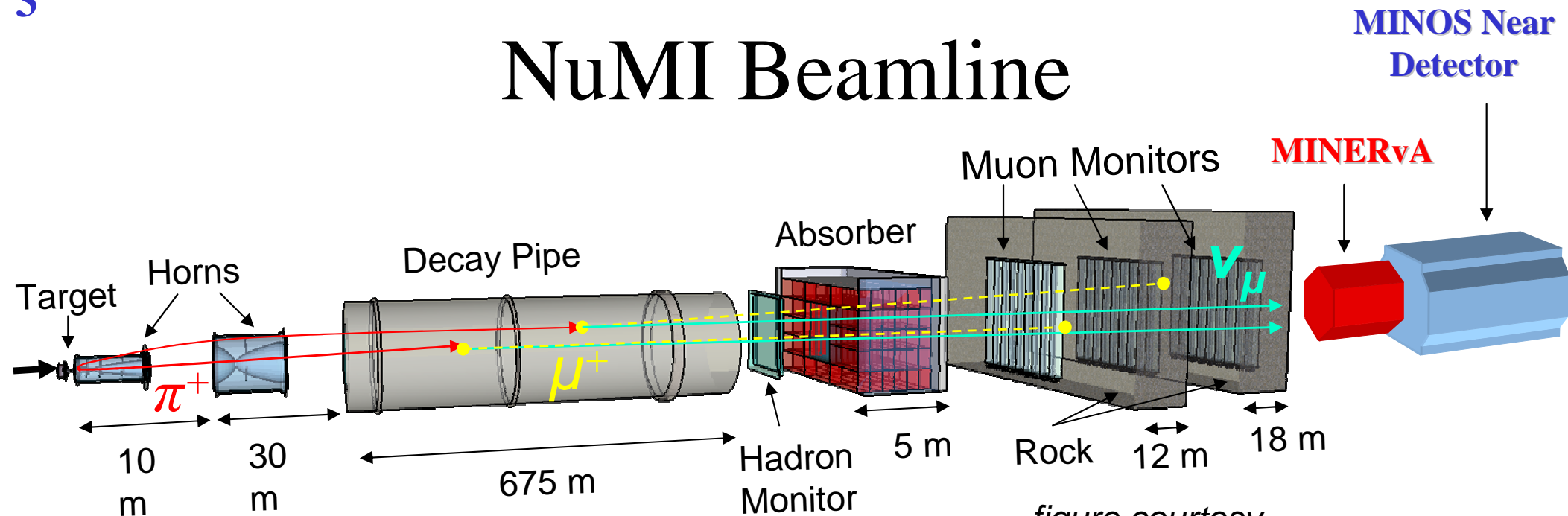
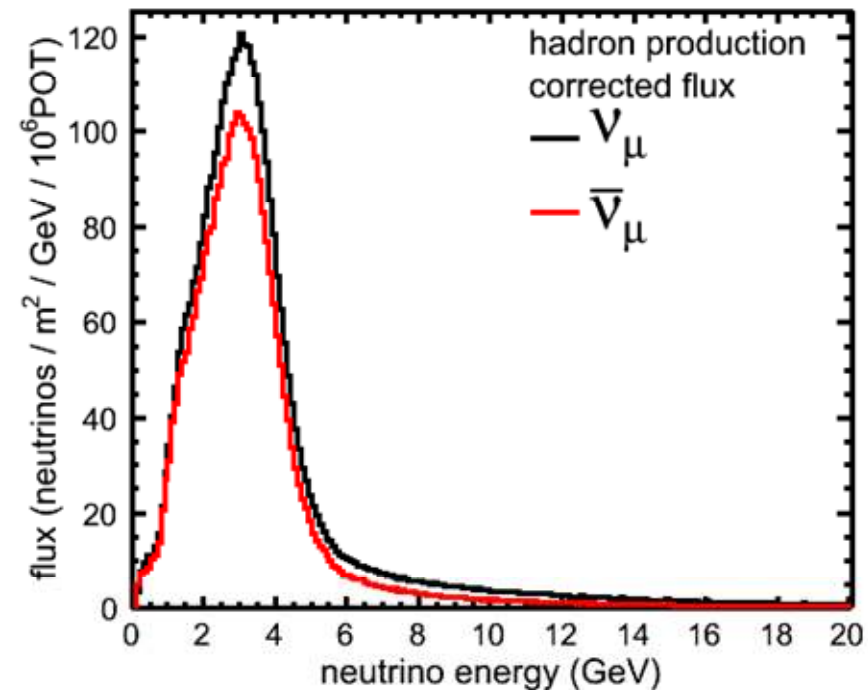


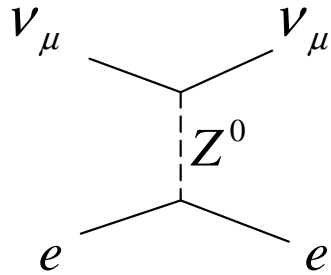
figure courtesy  
Ž. Pavlović

NuMI Low Energy Beam, FTFP



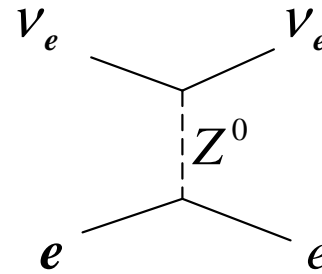
- Movable target to configure beam energy (only low energy target)
  - ME target doesn't move
- Horn current to select sign of neutrino
  - Forward horn current: neutrino dominant beam
  - Reverse horn current: anti-neutrino dominant beam
- My study is based on
  - Low energy beam, forward horn current
  - Contamination of  $\nu_{\mu}$ ,  $\bar{\nu}_e$ , and  $\bar{\nu}_{\mu}$

# Constraining flux using $\nu_\mu + e$ scattering

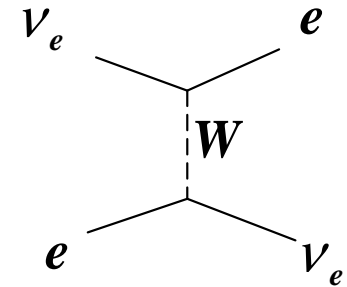


$$\nu_\mu + e^- \rightarrow \nu_\mu + e^-$$

$$\bar{\nu}_\mu + e^- \rightarrow \bar{\nu}_\mu + e^-$$



+

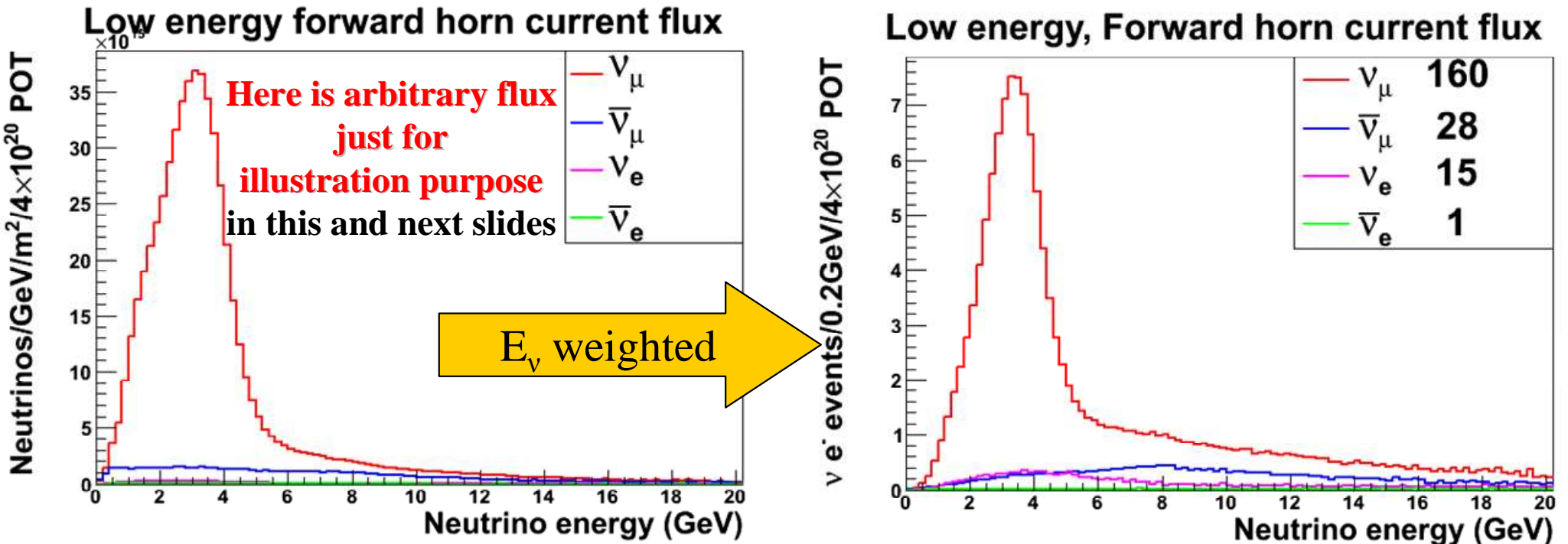


$$\nu_e + e^- \rightarrow \nu_e + e^-$$

$$\bar{\nu}_e + e^- \rightarrow \bar{\nu}_e + e^-$$

- $\nu + e$  scattering is pure leptonic process and theoretically well understood ( $\sim 1\%$  precision)
- $\nu$  scattering on light electron means small center of mass energy, consequently it has tiny cross section ( $\sim 1/2000$  compare to  $\nu N$  scattering)
- Scattering on light electron also means very small  $Q^2$ , which produces very forward electron final state
- In principle, if we measure event rate of this process, we can determine flux ( $R = \Phi\sigma$ )
- But it's not that simple because cross section( $\sigma$ ) and flux( $\Phi$ ) are function of neutrino energy
- And we only measure electron energy
  - Because electron angle is really all forward within detector resolution ( $E\theta^2 < 2m_e$ ), we don't have sensitivity to calculate neutrino energy using 2 body kinematics

# Flux $\rightarrow$ Event Rate



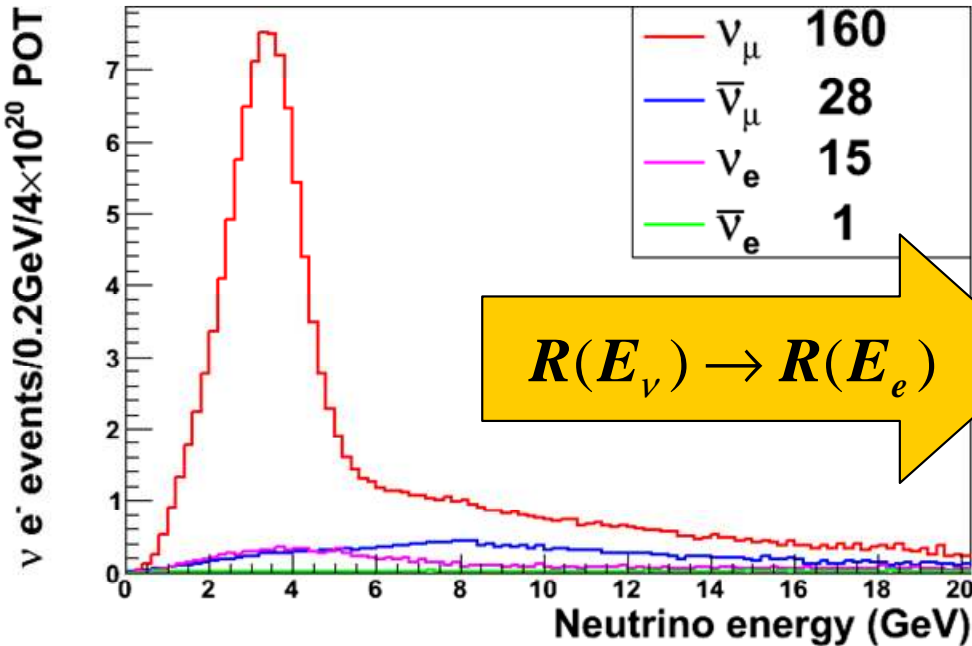
$$\sigma(\nu_\mu e^-) = \frac{G_F^2 s}{\pi} \left[ \left( -\frac{1}{2} + \sin^2 \theta_W \right)^2 + \frac{1}{3} \sin^4 \theta_W \right] \propto E_\nu$$

$$s = 2m_e E_\nu$$

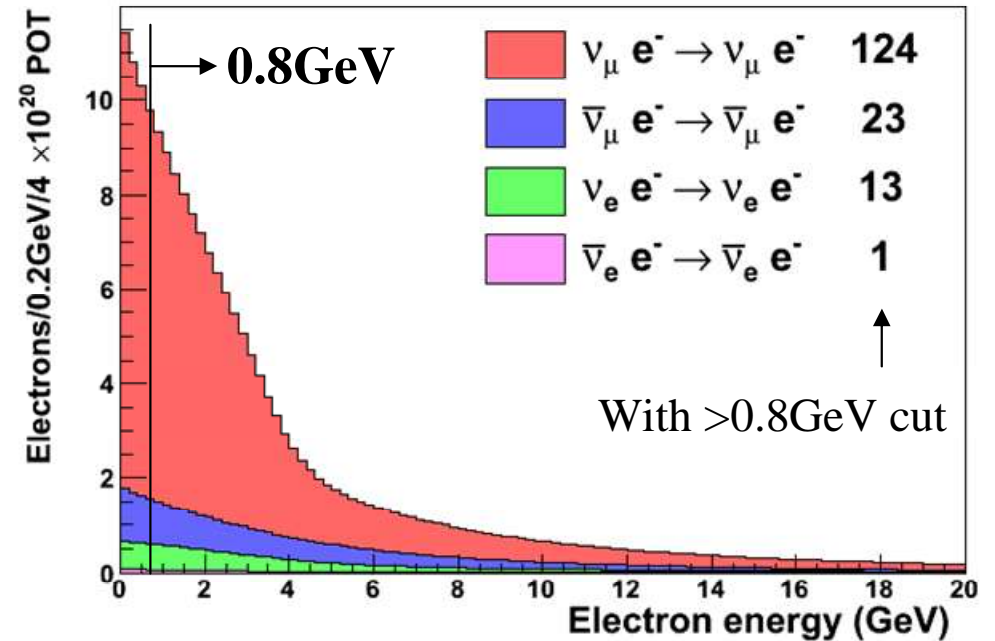
- Total cross section is proportional to beam energy
- High energy tail contribution gets bigger

# Event Rate $\rightarrow$ Electron Spectrum

Low energy, Forward horn current flux



Electron energy (Low energy, Forward horn current)

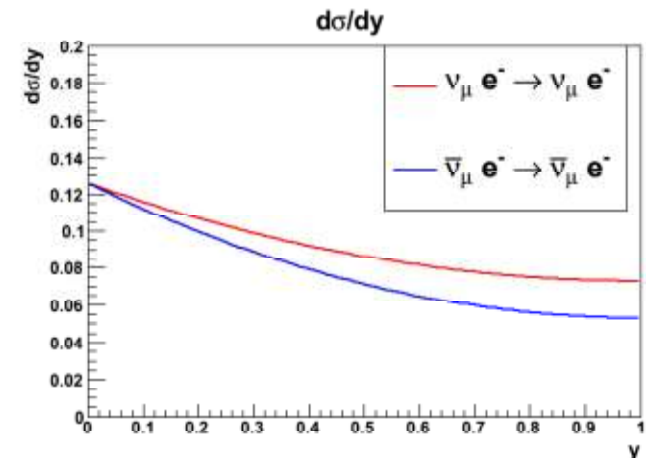


$$\frac{d\sigma(\nu_\mu e^- \rightarrow \nu_\mu e^-)}{dy} = \frac{G_F^2 m_e E_\nu}{2\pi} \left[ \left( \frac{1}{2} - \sin^2 \theta_W \right)^2 + \sin^4 \theta_W (1-y)^2 \right]$$

$$R(E_e) = \int \Phi(E_\nu) \frac{d\sigma(E_\nu, E_e)}{dE_\nu dE_e} dE_\nu$$

$$y = \frac{\text{(electron KE)}}{\text{(neutrino energy)}}$$

- High energy electron from high energy neutrino
- Low energy electron from both low and high energy neutrino
- Note also anti muon neutrino and electron neutrino contribution

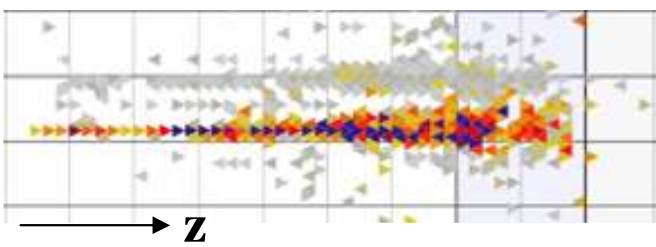
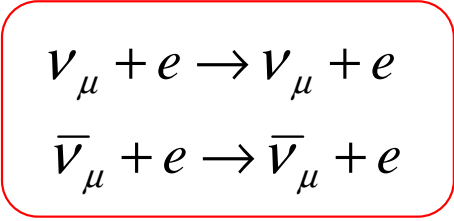


# $\nu_\mu + e^- \rightarrow \nu_\mu + e^-$ and Background

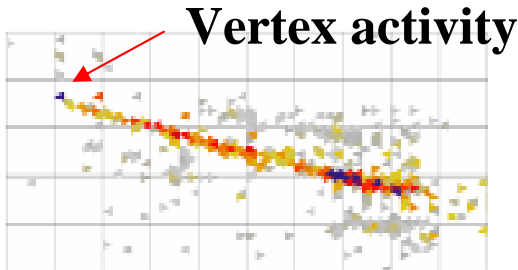
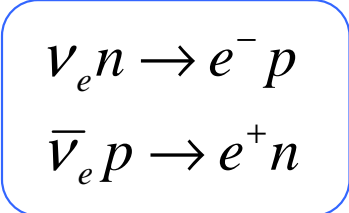
## Events

All **MC** events are shown

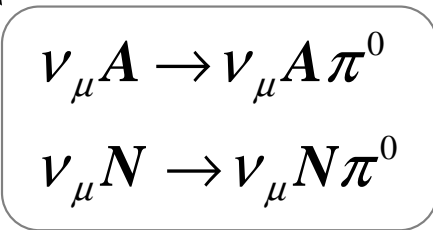
All can look like single electron final state



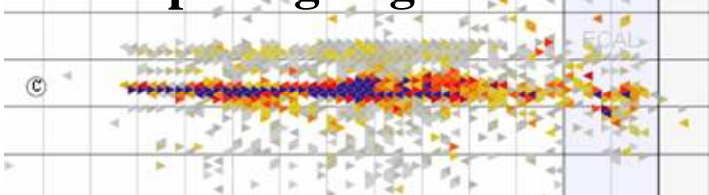
Very forward



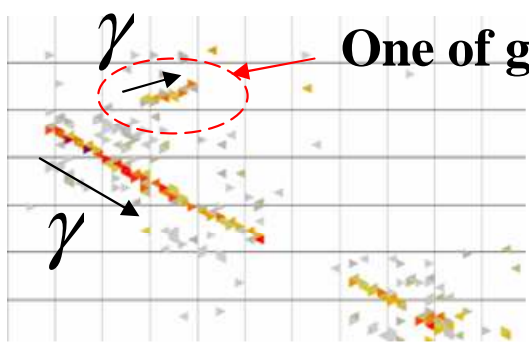
Vertex activity



Small opening angle between two gammas



NC-coherent  $\pi^0$   
NC-resonant  $\pi^0$

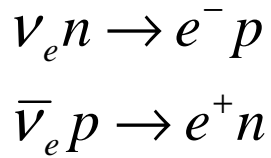


One of gamma is low energy



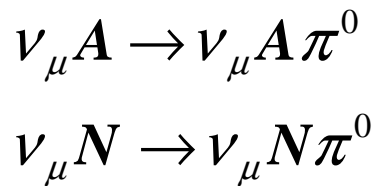
# Background Suppression and Signal Isolation

$\nu_e$  CCQE



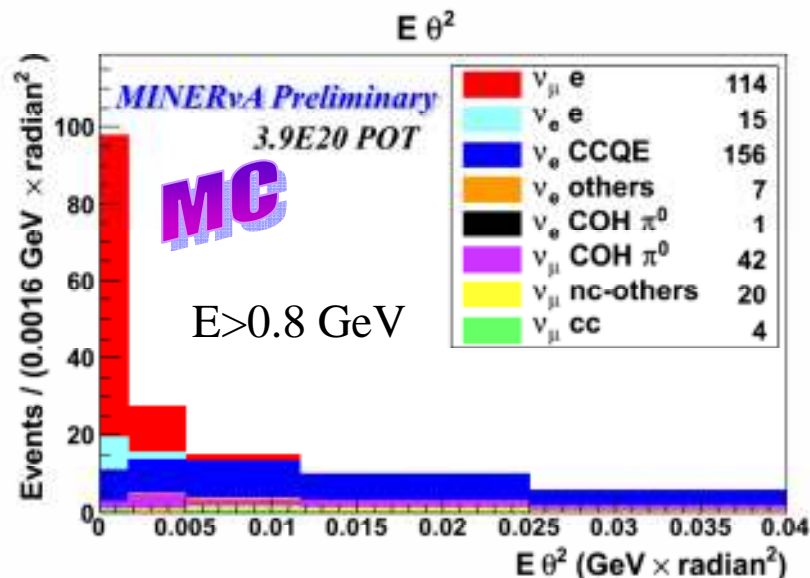
Require no or small vertex activity

NC  $\pi^0$



Gamma's dE/dx at the beginning of shower is different from electron

- Require EM shower energy > 0.8 GeV
  - NC background is very high at lower energy
  - Particle ID is more difficult at lower energy
- After basic background suppression, the signal is isolated using  $E\theta^2$  cut

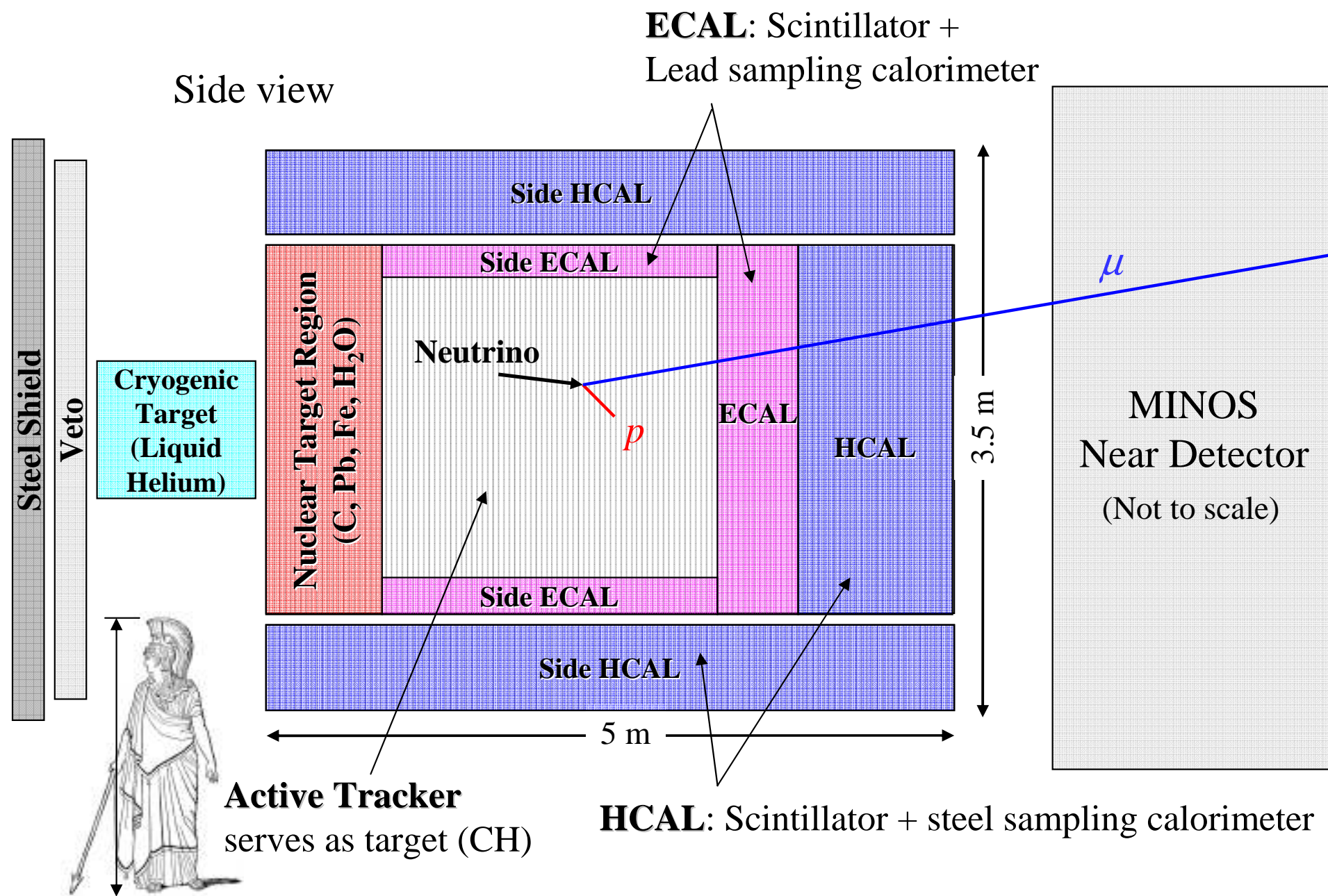


$E$ : Energy of electron candidate  
 $\theta$ : Theta of electron candidates  
w.r.t. beam direction



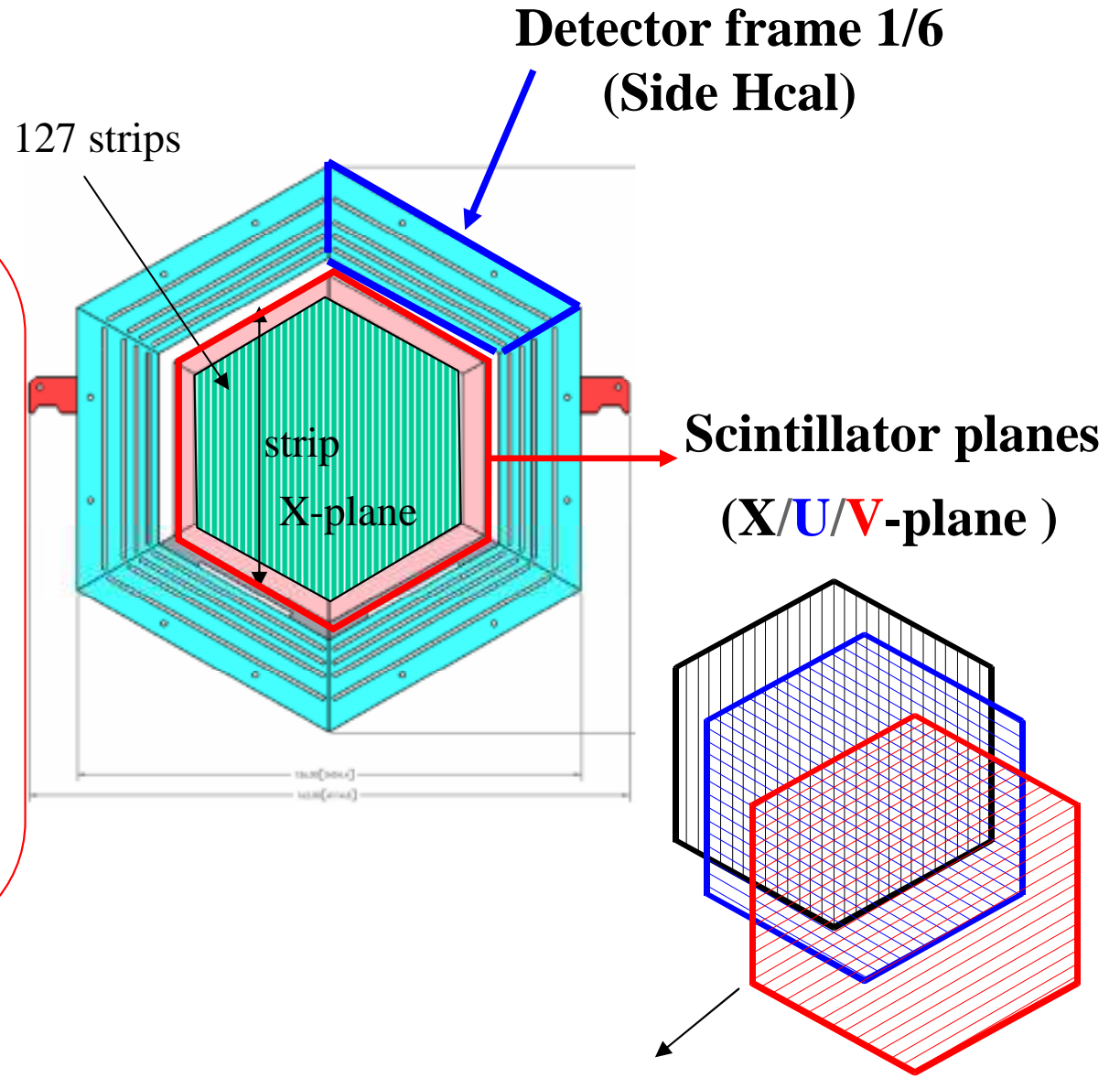
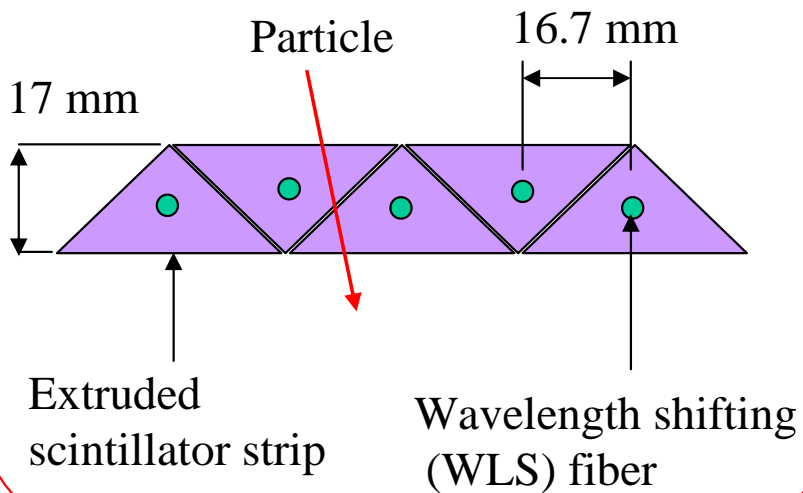
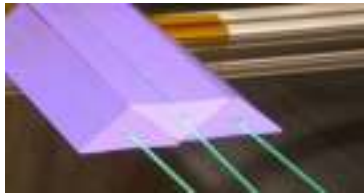
# MINERνA Detector

- MINERνA detector is made of a stack of “MODULES” (See next slide)



# Detector Module

Scintillator plane consists of extruded scintillator strips and wavelength shifting fibers

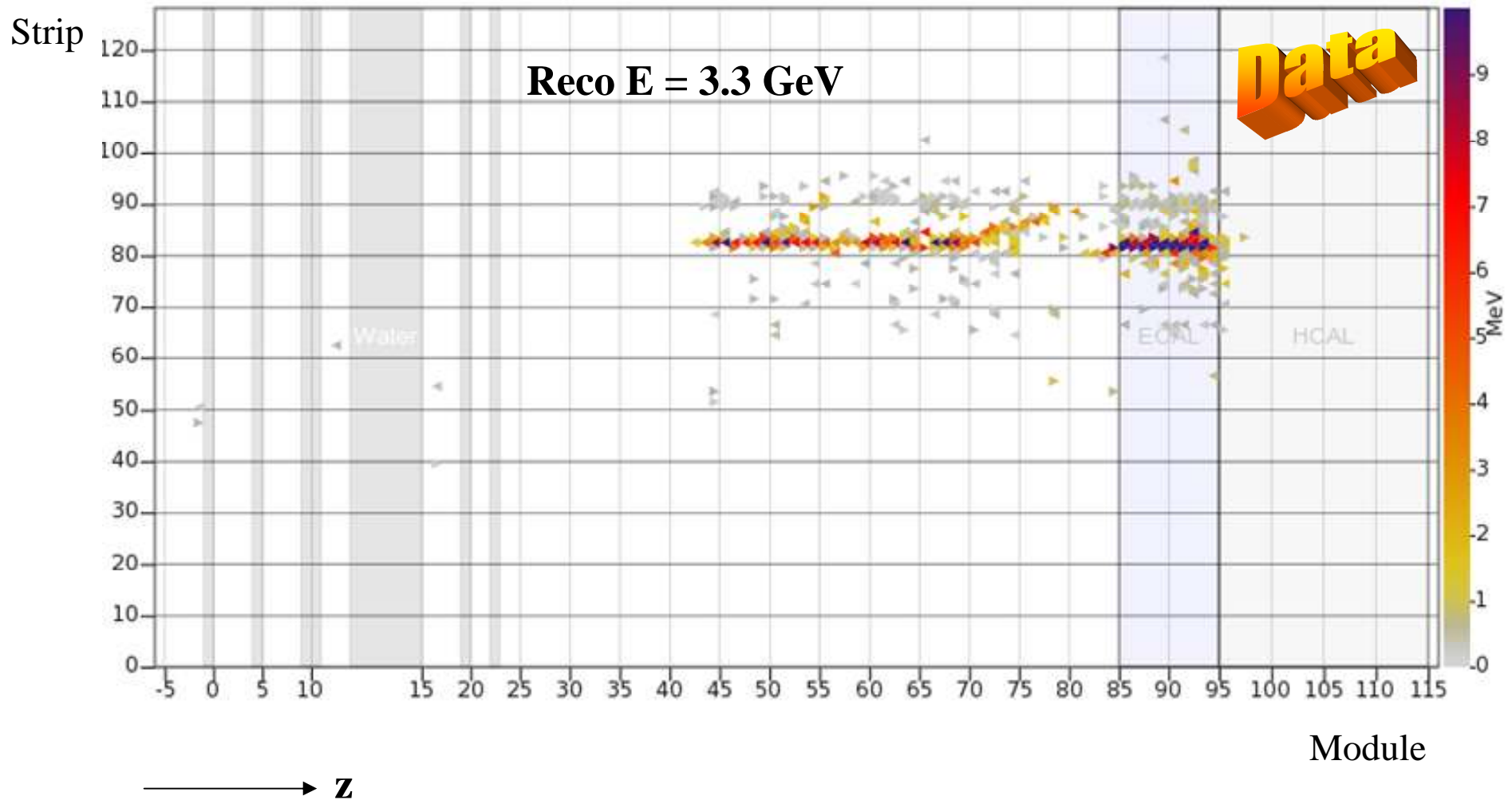


- X, U, V coordinates are combined to make 3D tracking

# $\nu+e \rightarrow \nu+e$ Candidate Event

X-view

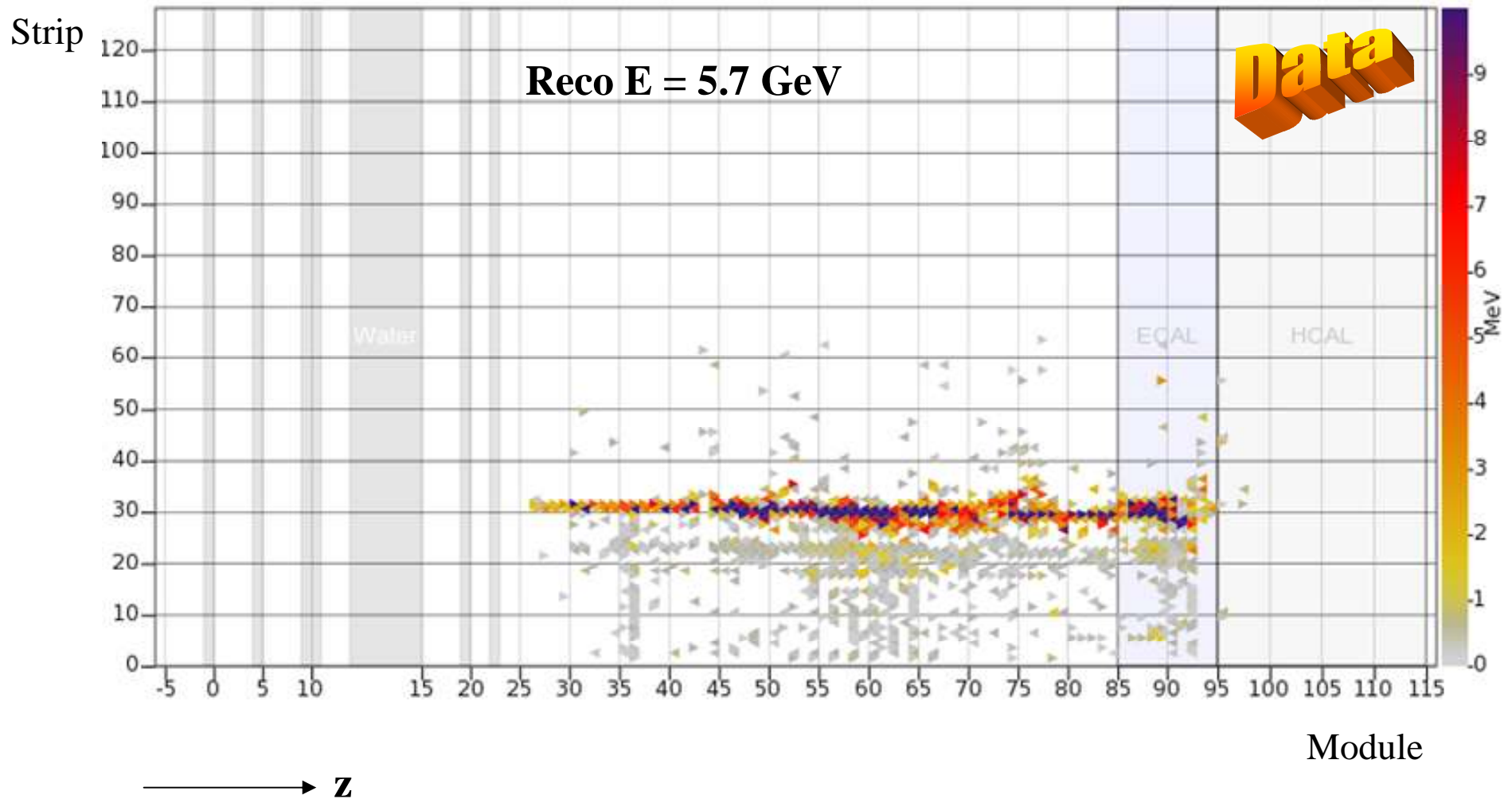
run/subrun/gate/timeslice=2017/2/219/10



# $\nu+e \rightarrow \nu+e$ Candidate Event

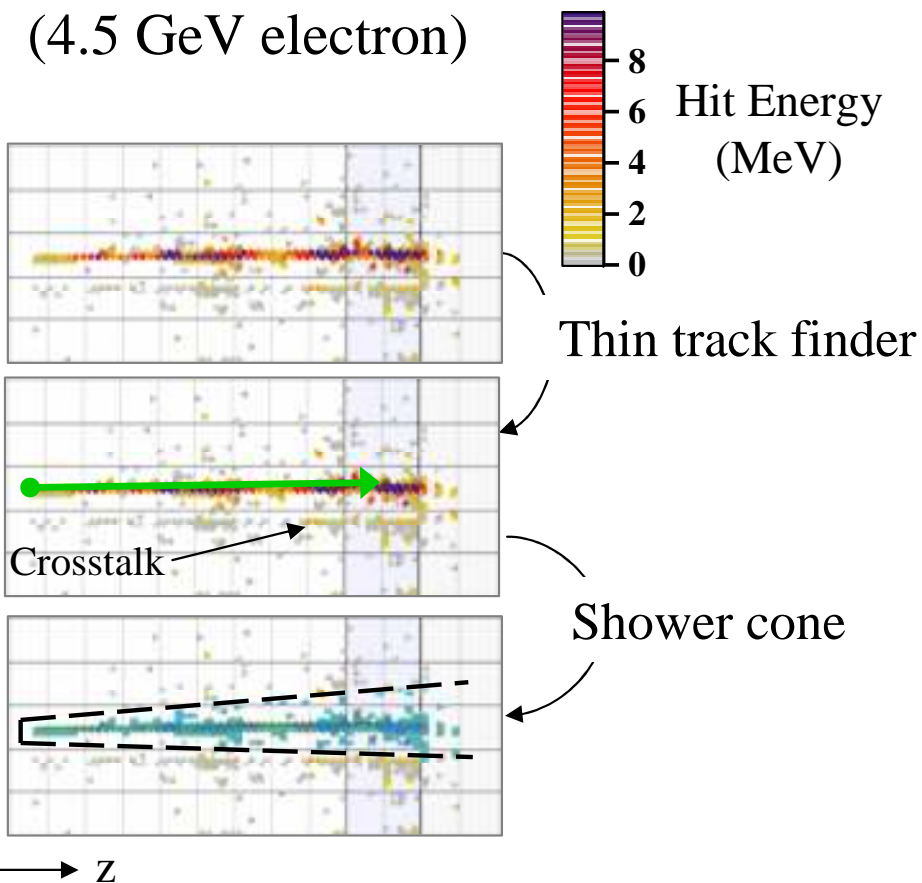
X-view

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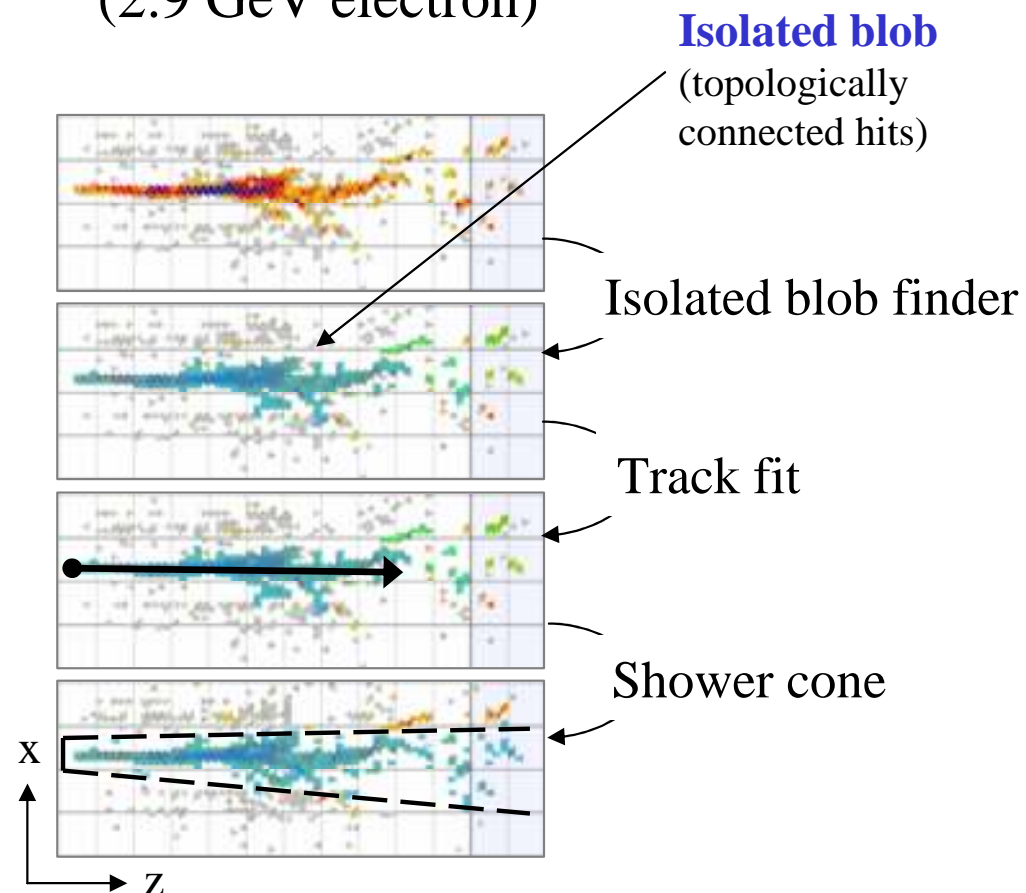


# Single EM Shower Reconstruction

## Track-like shower (4.5 GeV electron)



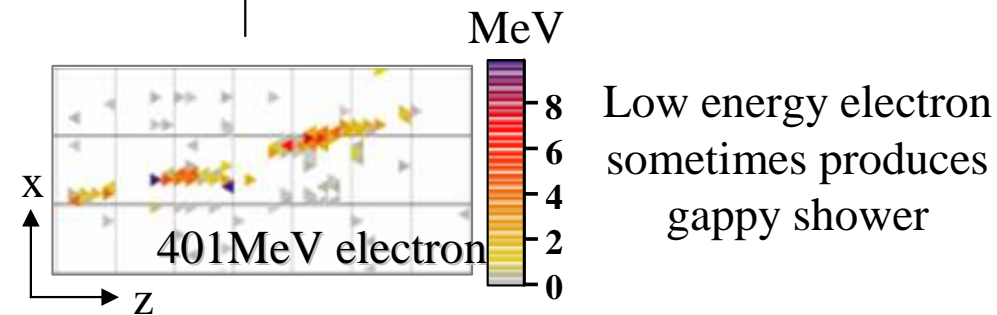
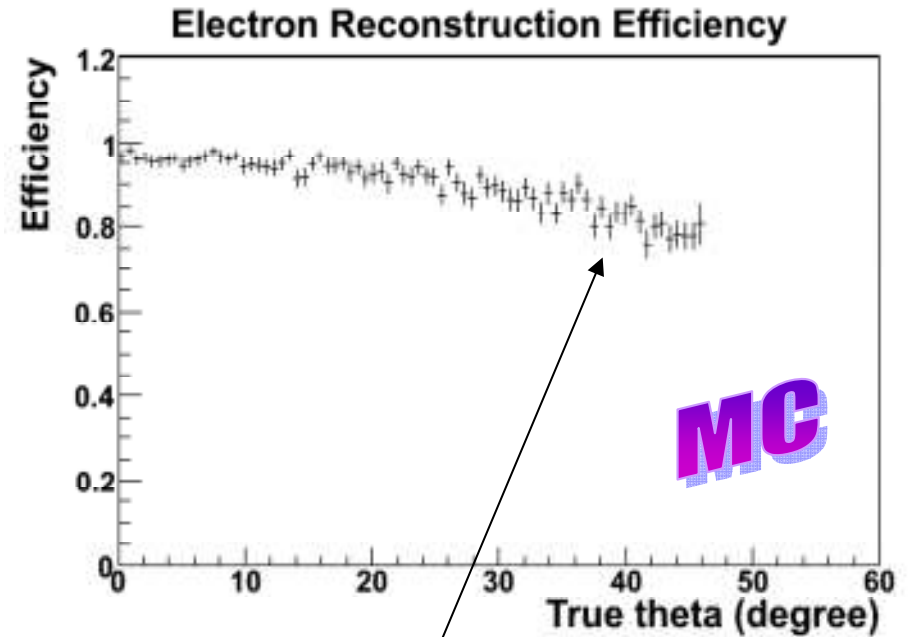
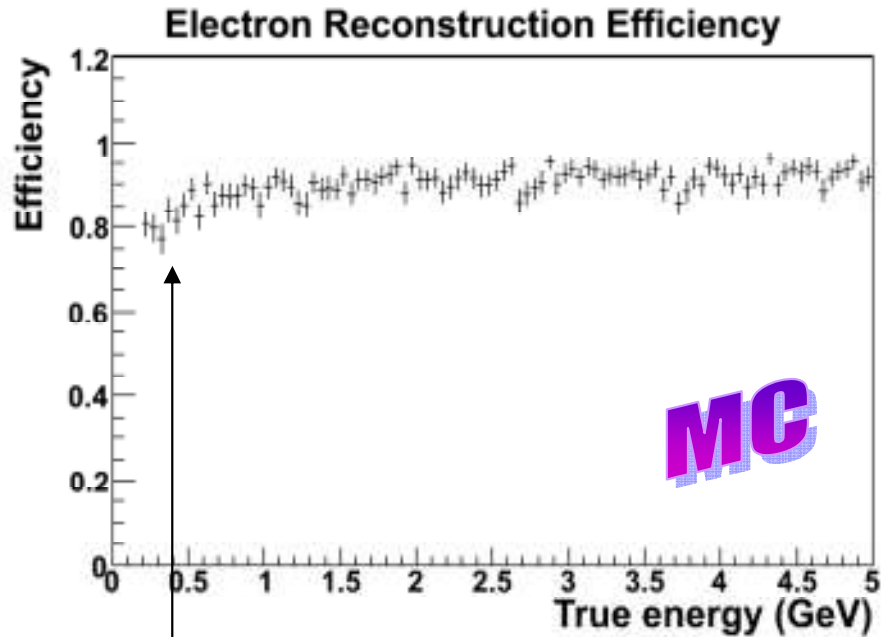
## Fuzzy shower (2.9 GeV electron)



- Once vertex and direction is known, shower cone can be applied
- When (thin) track finder fails on fuzzy shower, isolated blob finder is used and then track fitter can handle fuzzy shower



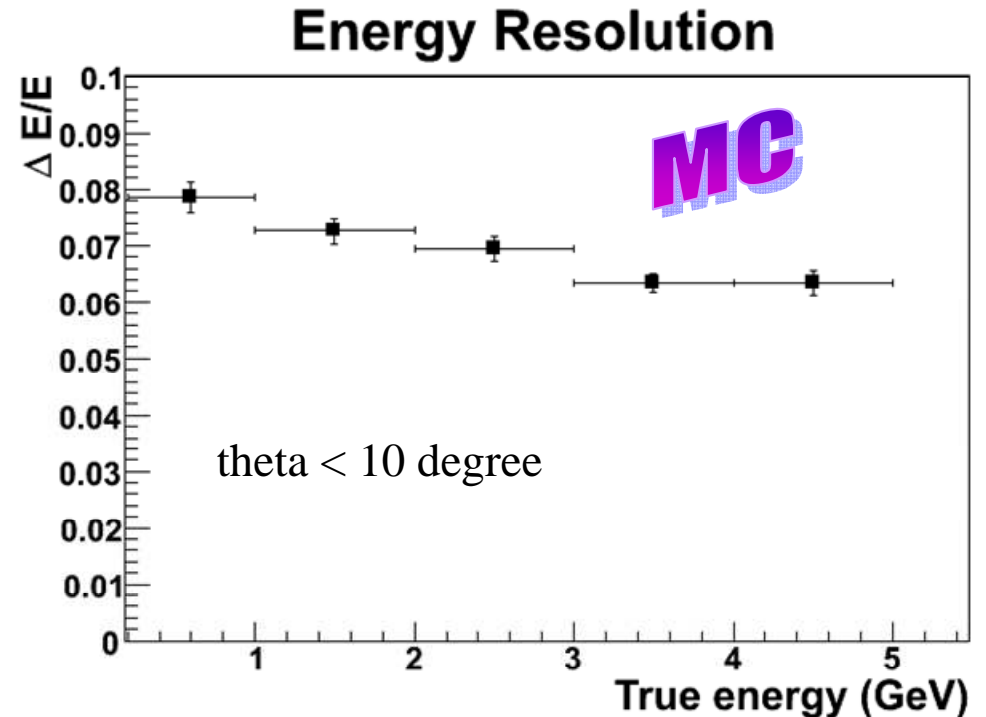
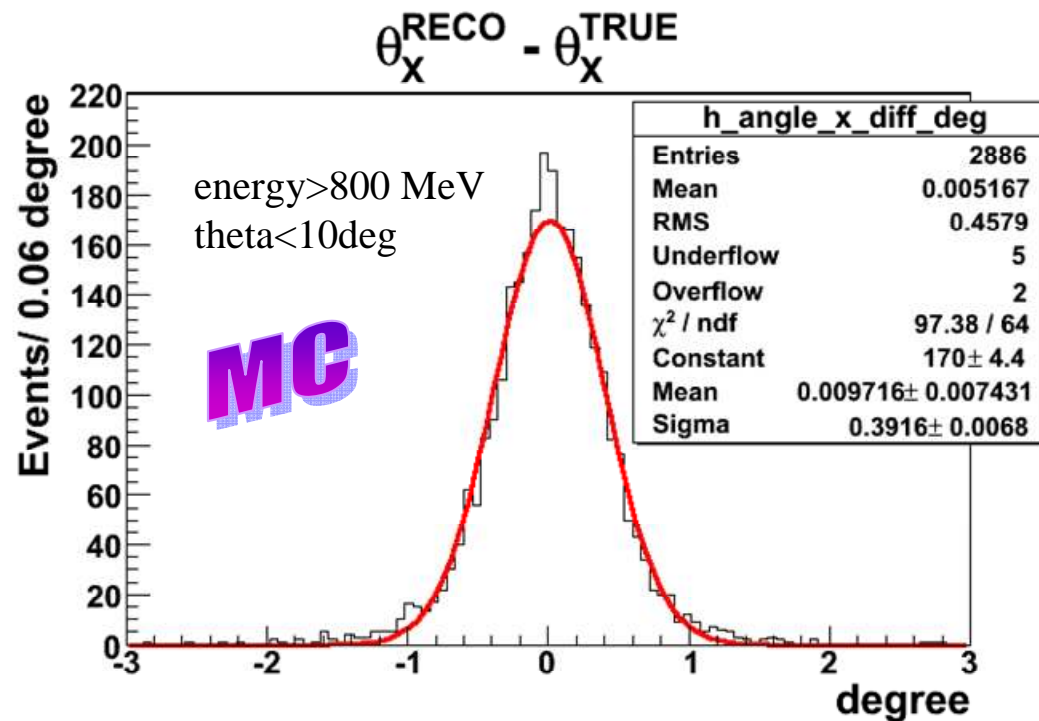
# MC Reconstruction Efficiency



Big theta angle electron tends to exit to sides, which leaves less hits in tracking volume

- Single electron MC is used to calculate efficiency
  - Energy: 0.2 ~ 5 GeV, Theta: 0~45 deg
- **Reconstruction efficiency is 0.96 for small angle (angle <10 degree, energy >400 MeV)**

# MC Angular and Energy Resolution



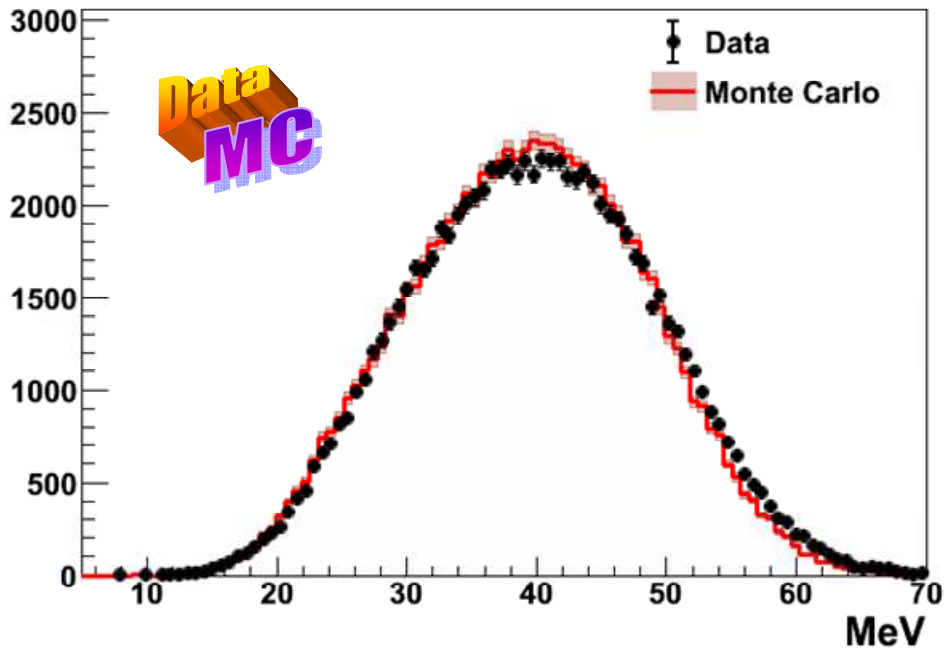
- X-angle resolution  $\sim 0.4$  degree
- Precise angular reconstruction is critical to separate  $\nu_\mu e$  elastic scattering from  $\nu_e$  CCQE
- Energy resolution: 6~ 7%



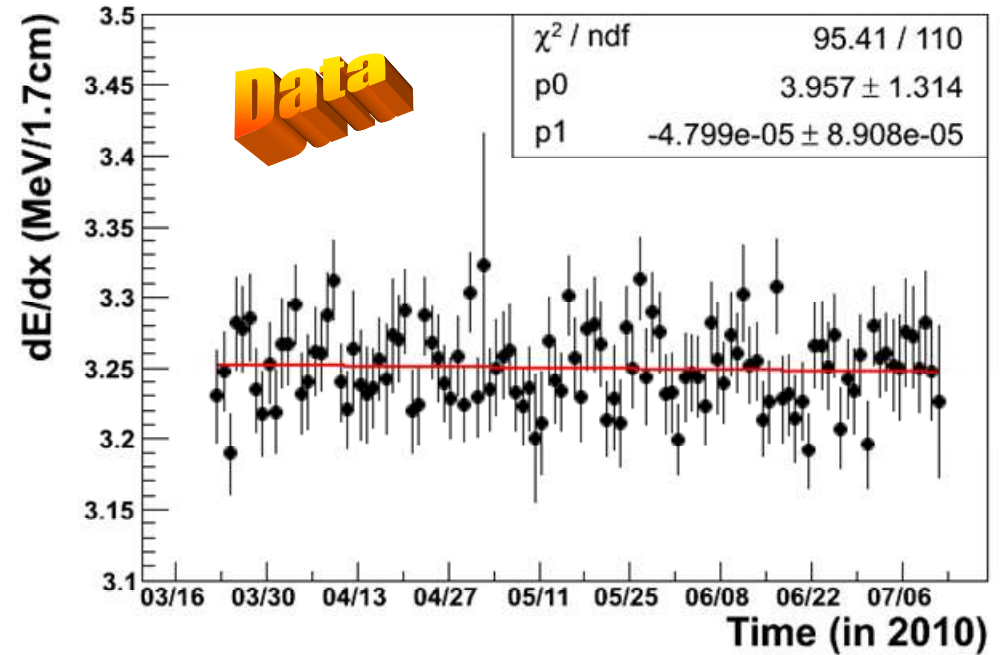
# Calibration Checking using Michel electron

Energy scale stability

Michel electron energy

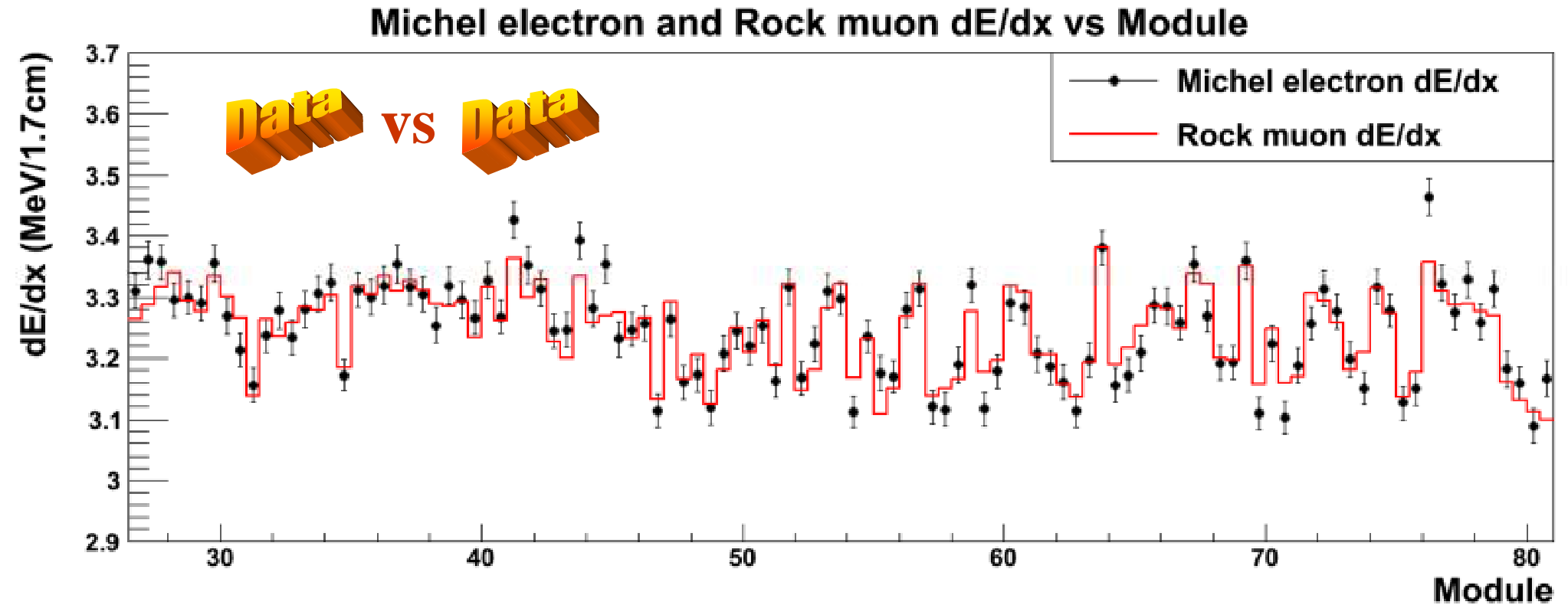


Michel electron dE/dx vs time



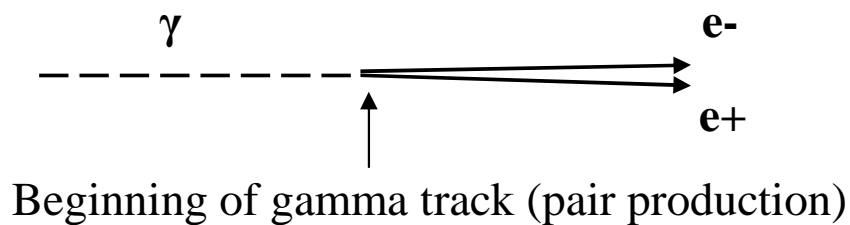
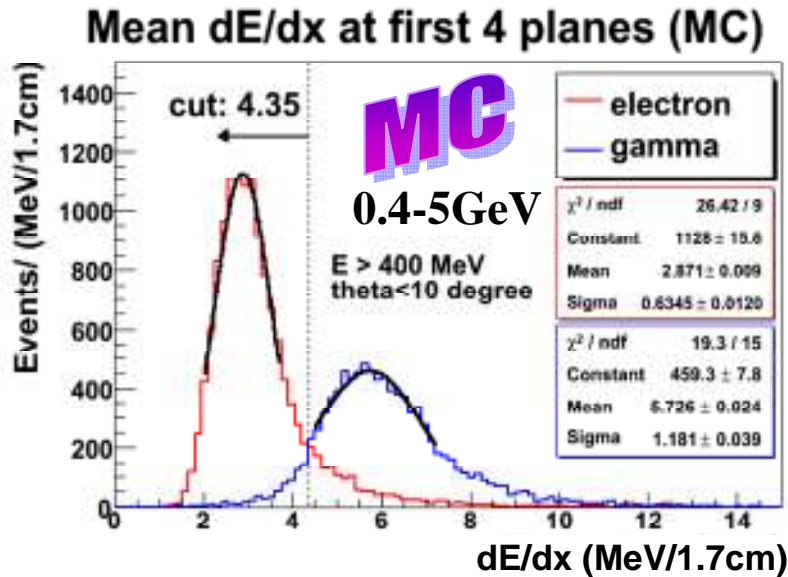
- Michel electron is produced by a muon decay and electron spectrum is predicted by well-understood theory
  - What's really seen is theory spectrum with detector energy resolution and calibration
- Michel electron is nice tool to check calibration
  - Michel energy MC/data comparison
  - EM energy scale is stable over time

# Michel electron $dE/dx$ and Rock muon $dE/dx$ Comparison

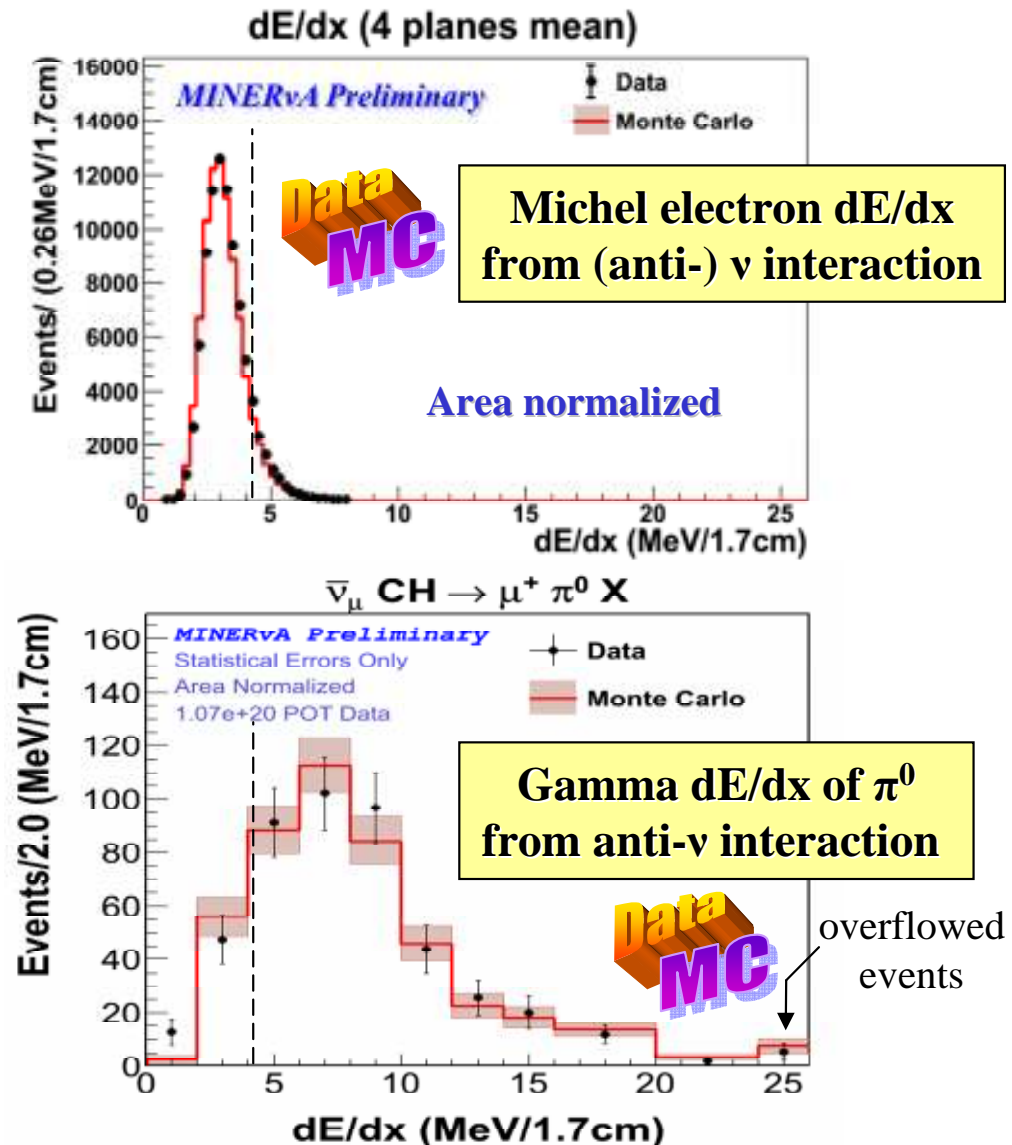


- Rock muon (muons that is produced from neutrino interaction in upstream rock) is also good source of calibration
  - Minimum ionizing particle (MIP)  $dE/dx$  is constant
- $dE/dx$  in each plane reflects plane response
- Module to module variation is consistent between Michel electron  $dE/dx$  and muon  $dE/dx$

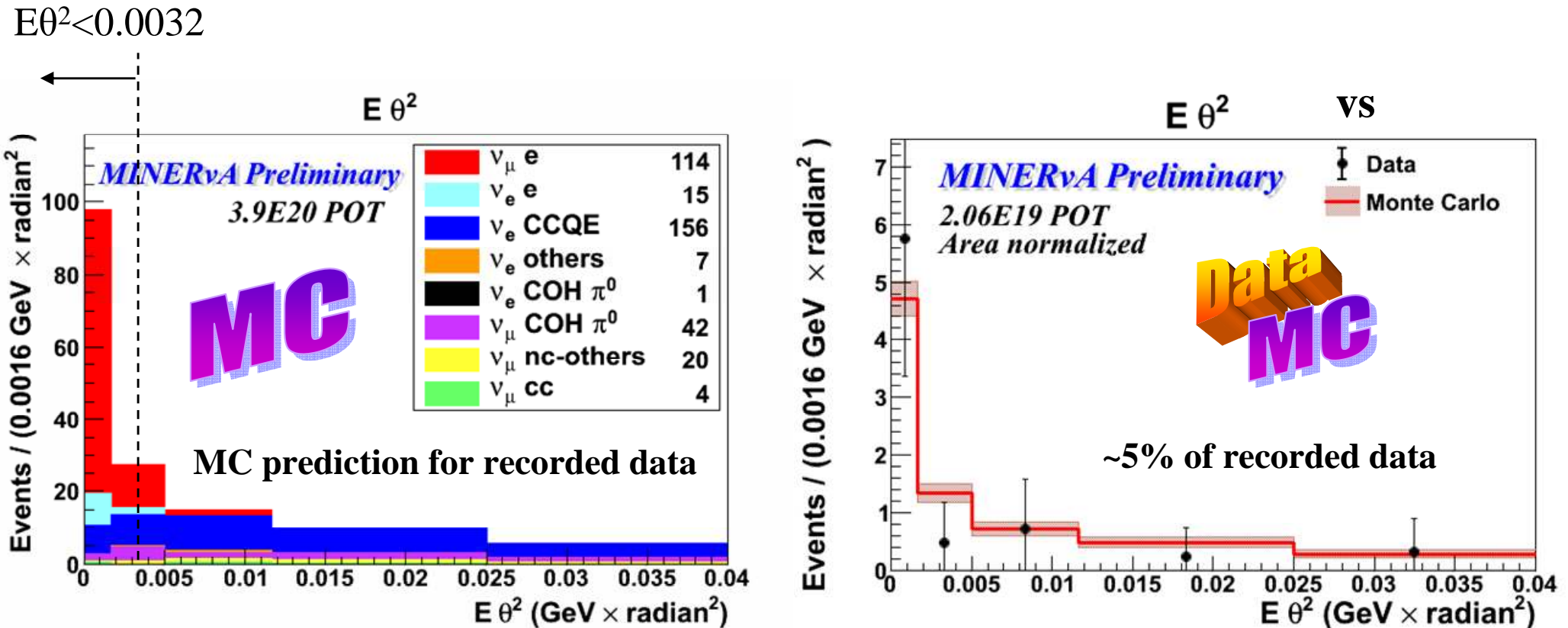
# dE/dx for Electron and Gamma Discrimination



- Neutral current  $\pi^0$  is decayed into energetic gamma + tiny energy gamma
- dE/dx at the beginning of shower is different for electron and gamma
  - Electron loses energy like MIP (Minimum Ionization Particle)
  - Gamma loses energy like twice MIP

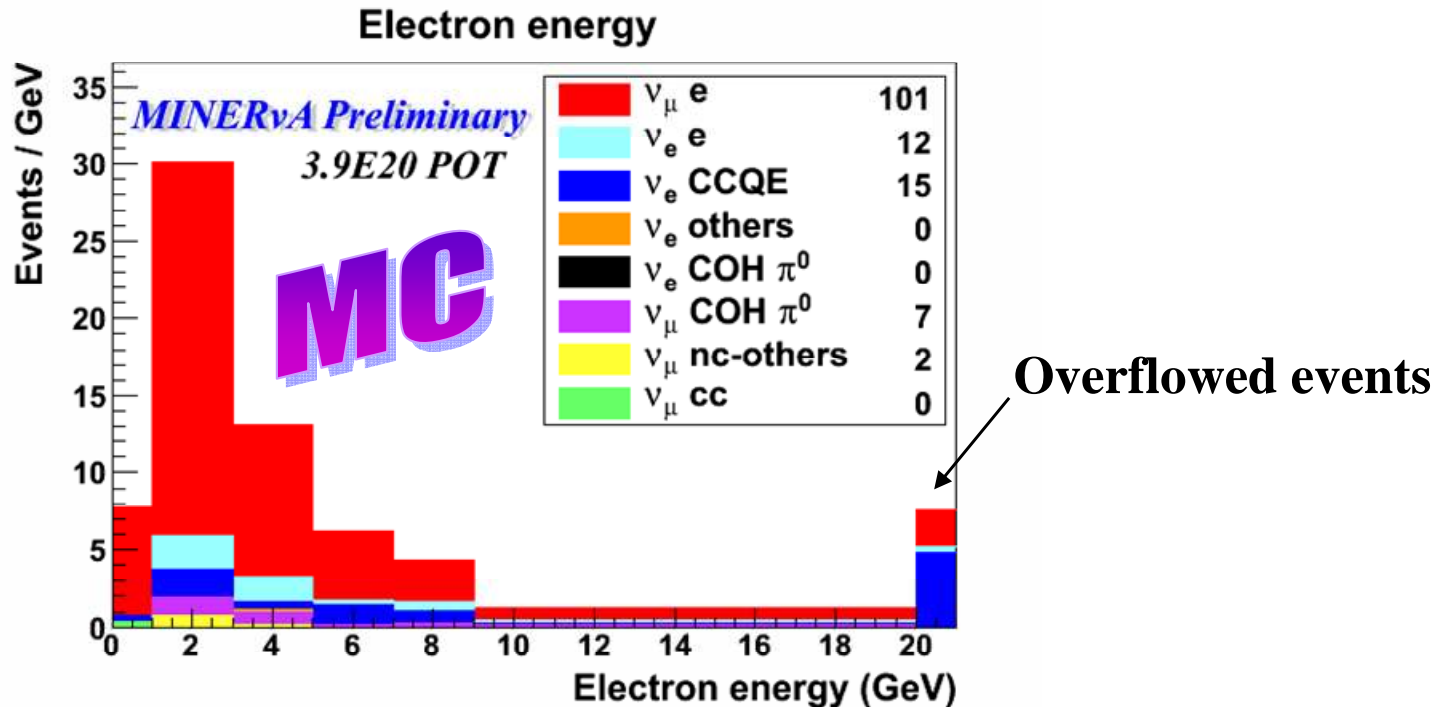


# Small Sample Data/MC comparison



- Neutrino beam, MC (left): 3.9E20 POT, data (right): 2.06E19 POT
- Small data (~ 5% to full data) is used for comparison
- Peak in low  $E\theta^2$  is found in data

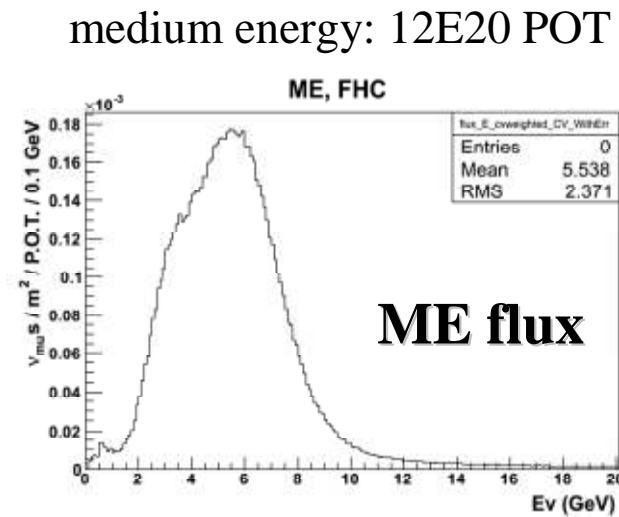
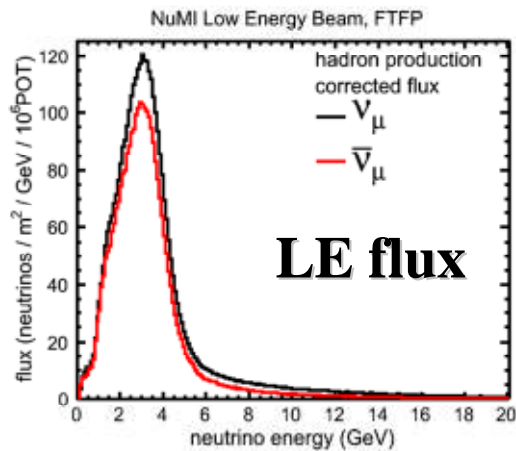
# Electron spectrum of $\nu+e^-$ Elastic Scattering



- $E\theta^2 < 0.0032$  cut is applied to get  $\nu_\mu + e^-$  rich sample
- Purity: 0.82, efficiency: 0.6
- Expected signal ( $\nu_\mu + e^-$ ,  $\nu_e + e^-$ ) is 112 events with 24 background events for expected full data set
- It gives  $\sim 10\%$  statistical error on absolute flux constraint

# What do we expect in ME?

- Medium energy beam expect roughly 10 times statistics compared to low energy beam
  - Assuming:
    - low energy: 4E20 POT
    - medium energy: 12E20 POT

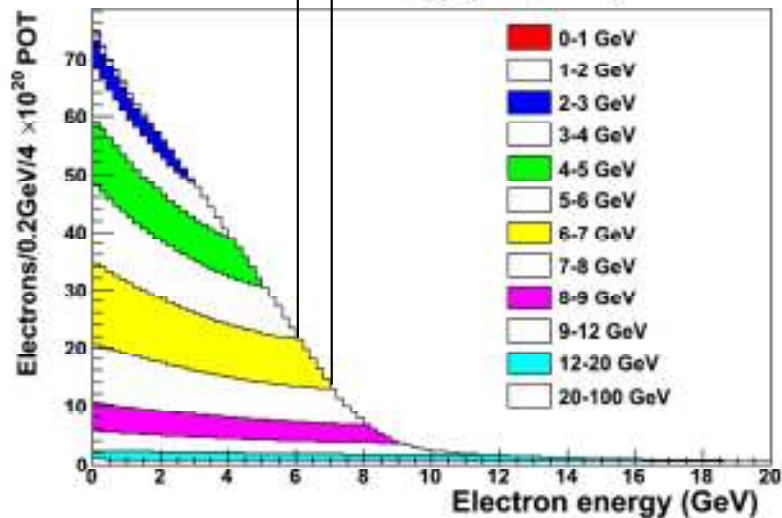
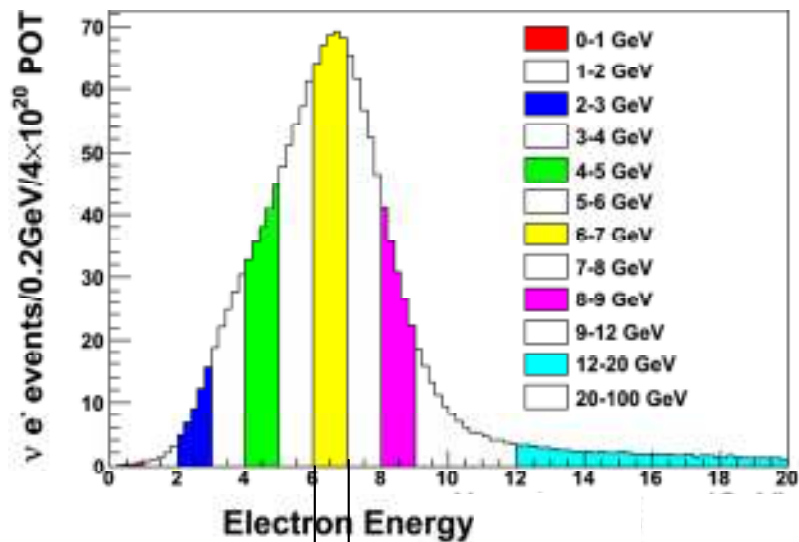


- If we assume we have similar signal/background ratio as LE:
  - Signal/Background = 173/47 (LE)
  - Signal/Background = 1730/470 (ME, scaled from LE)
  - Statistical error =  $\sim 2.7\%$



# Flux Shape?

Event Rate (using arbitrary flux)



- In principle, deconvolution (electron spectrum  $\rightarrow$  flux shape) is possible but it's not easy
- It's easier to compare in electron spectrum space
- Two different flux model predictions can be tested with data and we can tell which one is more consistent



# Summary

- Absolute flux constraining based on  $\nu_{\mu}+e^{-}$  scattering is shown useful in LE along with other methods of flux measurement
- It'll be more powerful method in higher statistics ME